

Sometimes What Glitters is Gold

The Detection of an Optical Counterpart to a
Neutron Star-Neutron Star Merger

Ben Shappee & Maria Drout

GW170817 Papers

Query Results from the ADS Database

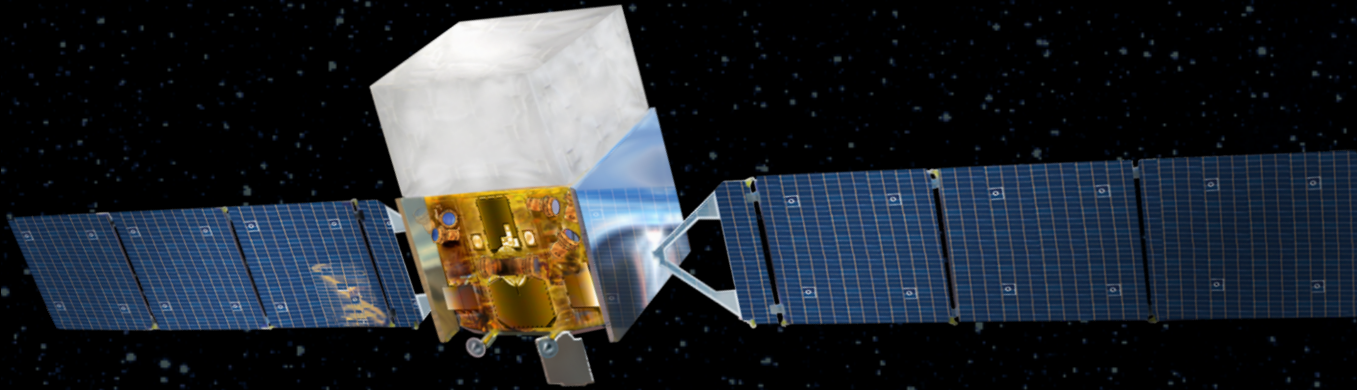
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#	Bibcode Authors	Cites Title	Date	List of Links Access Control Help			
1	<input type="checkbox"/> 2017arXiv171005432S Shappee, B. J.; Simon, J. D.; Drout, M. R.; Piro, A. L.; Morrell, N.; Prieto, J. L.; Kasen, D.; Holoien, T. W.-S.; Kollmeier, J. A.; Kelson, D. D.; and 26 coauthors	3.000	10/2017	A Z	X	R C	U
2	<input type="checkbox"/> 2017arXiv171005452C Coulter, D. A.; Foley, R. J.; Kilpatrick, C. D.; Drout, M. R.; Piro, A. L.; Shappee, B. J.; Siebert, M. R.; Simon, J. D.; Ulloa, N.; Kasen, D.; and 7 coauthors	2.000	10/2017	A Z	X	R C	U
3	<input type="checkbox"/> 2017arXiv171005443D Drout, M. R.; Piro, A. L.; Shappee, B. J.; Kilpatrick, C. D.; Simon, J. D.; Contreras, C.; Coulter, D. A.; Foley, R. J.; Siebert, M. R.; Morrell, N.; and 34 coauthors	2.000	10/2017	A Z	X	R C	U
4	<input type="checkbox"/> 2017arXiv171005434K Kilpatrick, Charles D.; Foley, Ryan J.; Kasen, Daniel; Murguia-Berthier, Ariadna; Ramirez-Ruiz, Enrico; Coulter, David A.; Drout, Maria R.; Piro, Anthony L.; Shappee, Benjamin J.; Boutsia, Konstantina; and 11 coauthors	2.000	10/2017	A Z	X	R C	U
5	<input type="checkbox"/> 2017arXiv171005901M	1.000	10/2017	A Z	X	R C	U

The Discovery Actors



Carnegie Observatories



+ Juna, Barry, Carlos, Konstantina, Francesco, David
and observations from 12 other OCIW/LCO scientists and 18 other scientists

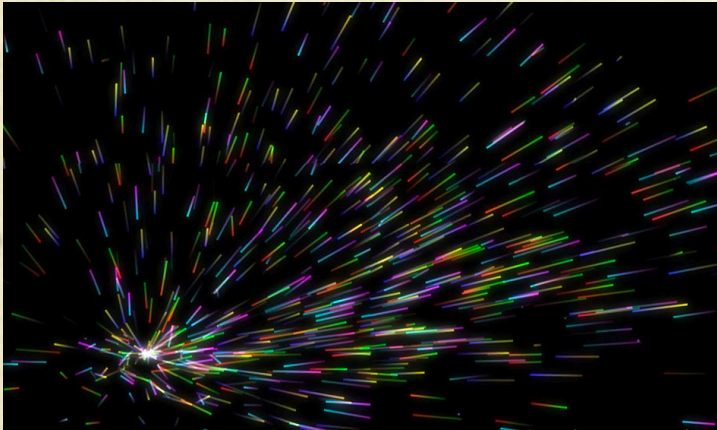
UC Santa Cruz



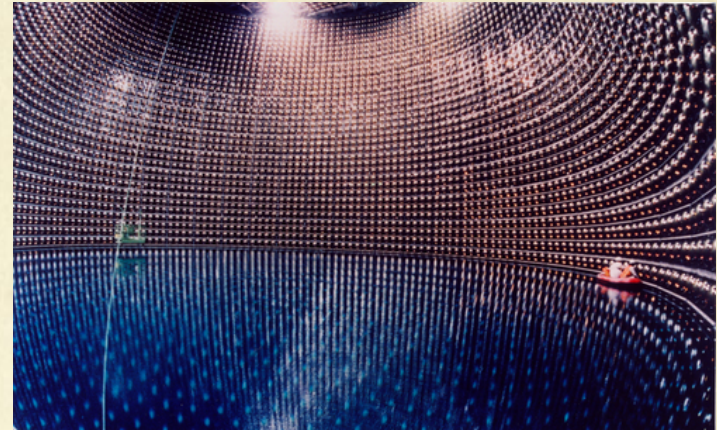
+ 70 observatories and 1500 physicists/astronomers around the world

Multi-messenger Astronomy

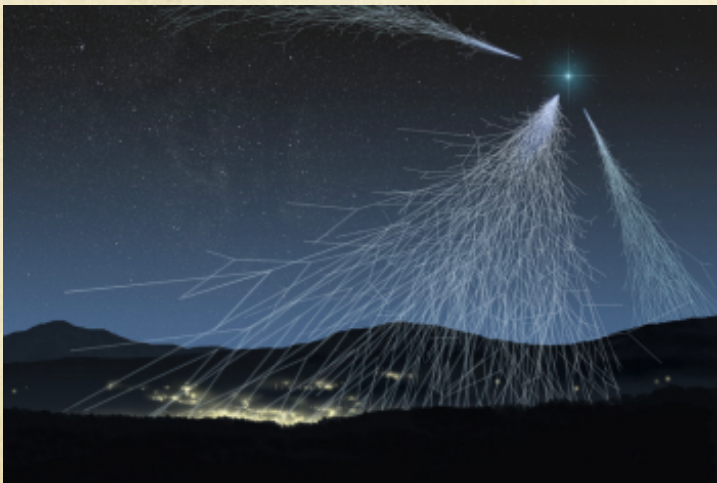
Multi-messenger Astronomy



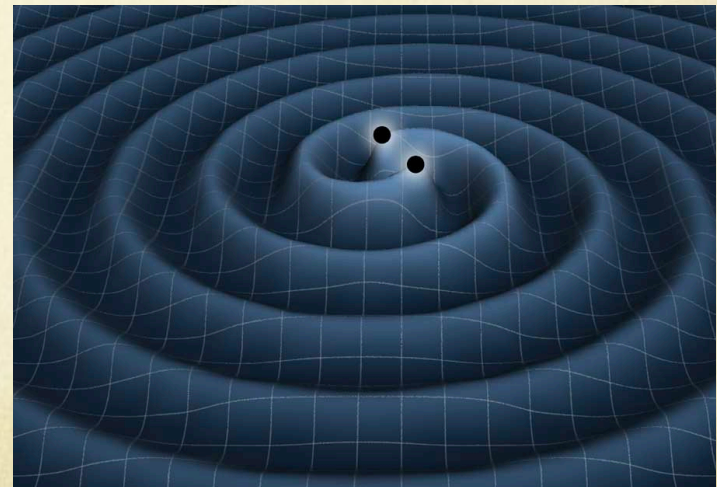
Photons



Neutrinos



Cosmic-Rays



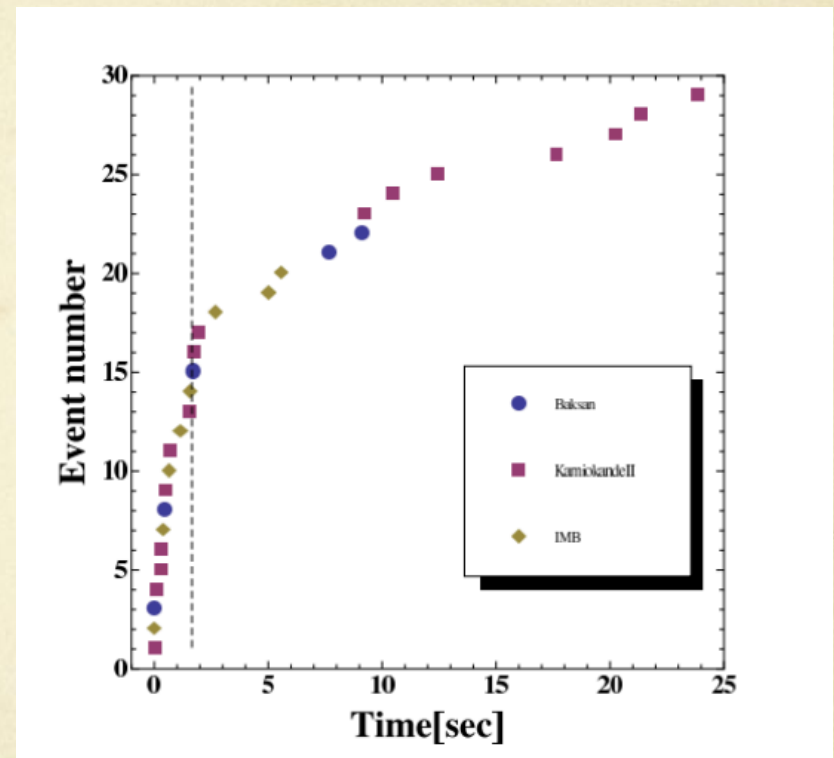
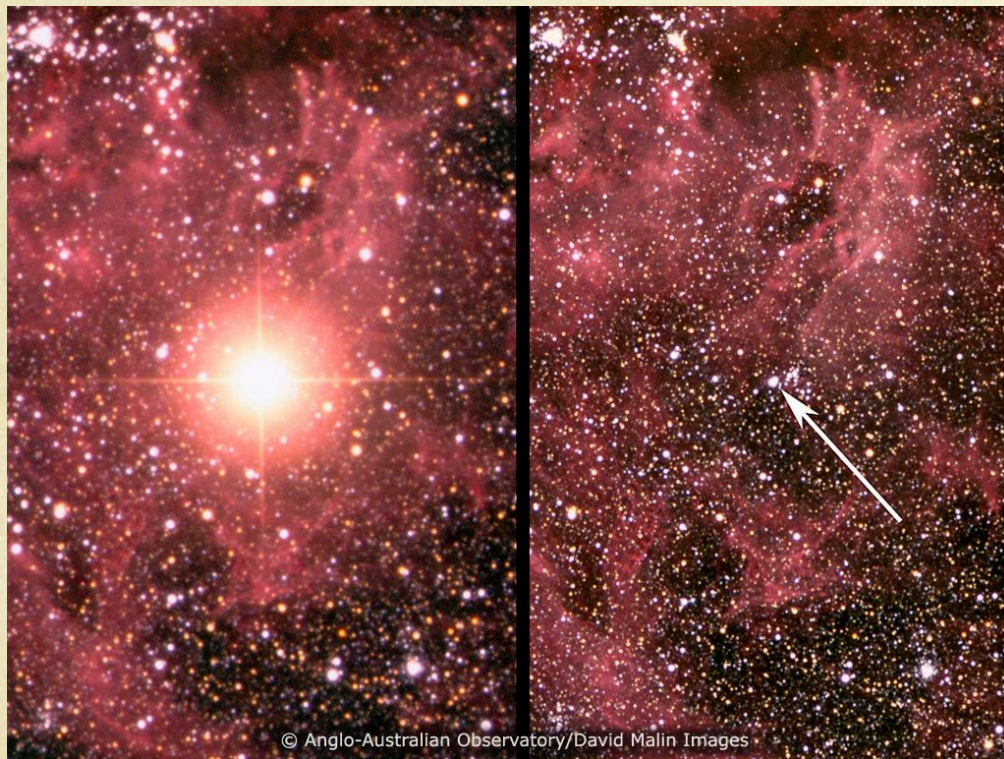
Gravitational Waves

Multi-messenger Astronomy

Success Stories

Multi-messenger Astronomy Success Stories

SN1987A: photons + neutrinos



Vissani et al. (2010)

Multi-messenger Astronomy

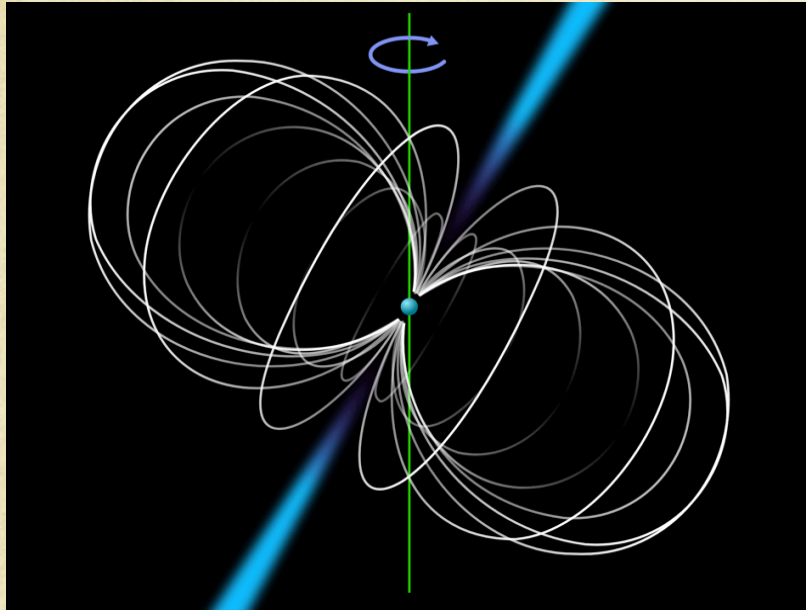
Gravitational Wave

Gravitational Waves



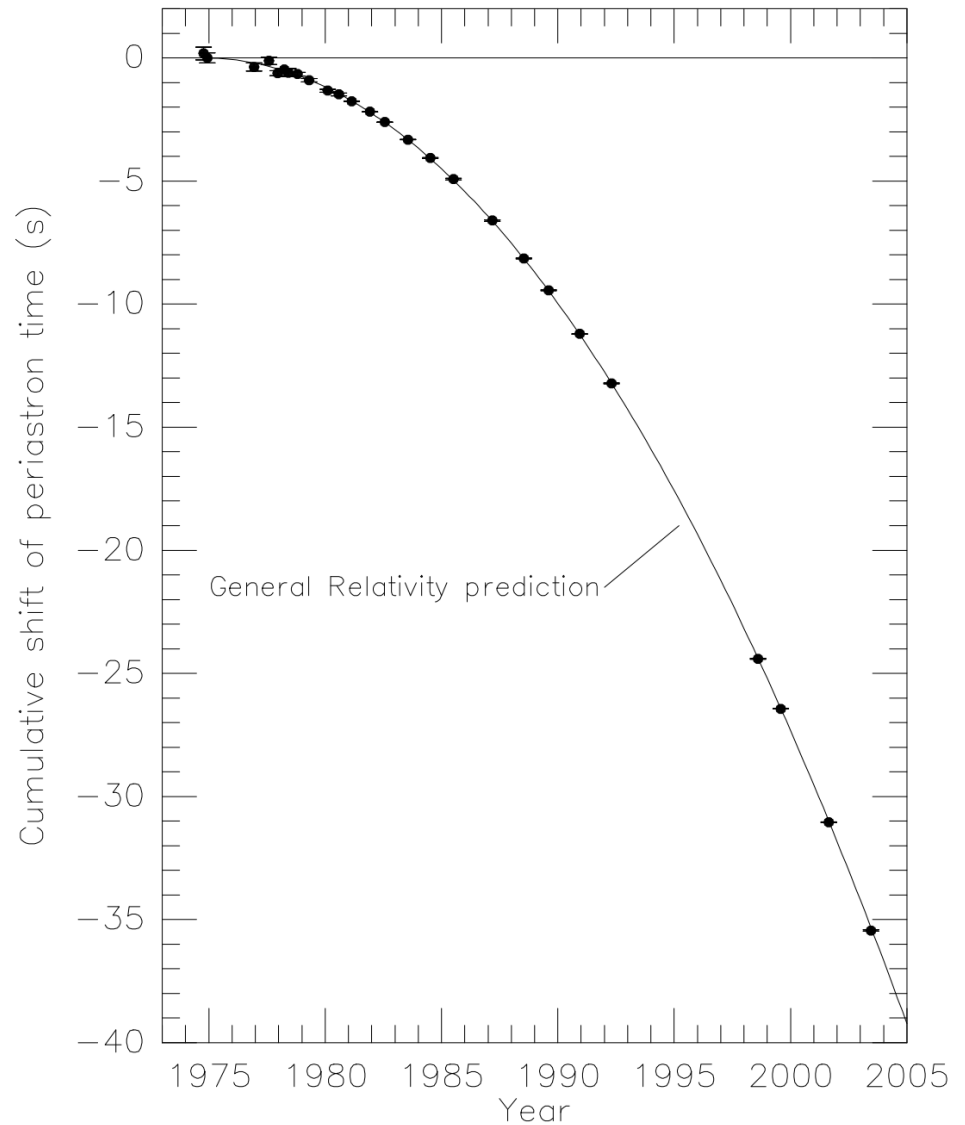
$$t_{\text{GW}} = \frac{3}{85} \frac{a_1}{c} \left(\frac{a_1^3 c^6}{G^3 m_0 m_1 M} \right) (1 - e_1^2)^{7/2}$$
$$\simeq 1.6 \times 10^{13} \text{ yr} \left(\frac{2 M_{\odot}^3}{m_0 m_1 M} \right) \left(\frac{a_1}{0.1 \text{ AU}} \right)^4 (1 - e_1^2)^{7/2},$$

Hulse-Taylor binary



Measured Orbital Parameters for PSR B1913+16

Fitted Parameter	Value
$a_p \sin i$ (s)	2.3417725 (8)
e	0.6171338 (4)
T_0 (MJD)	52144.90097844 (5)
P_b (d)	0.322997448930 (4)
ω_0 (deg)	292.54487 (8)
$\langle \dot{\omega} \rangle$ (deg/yr)	4.226595 (5)
γ (s)	0.0042919 (8)
\dot{P}_b (10^{-12} s/s)...	-2.4184 (9)



Hulse & Taylor 1975, Weisberg and Taylor 2005

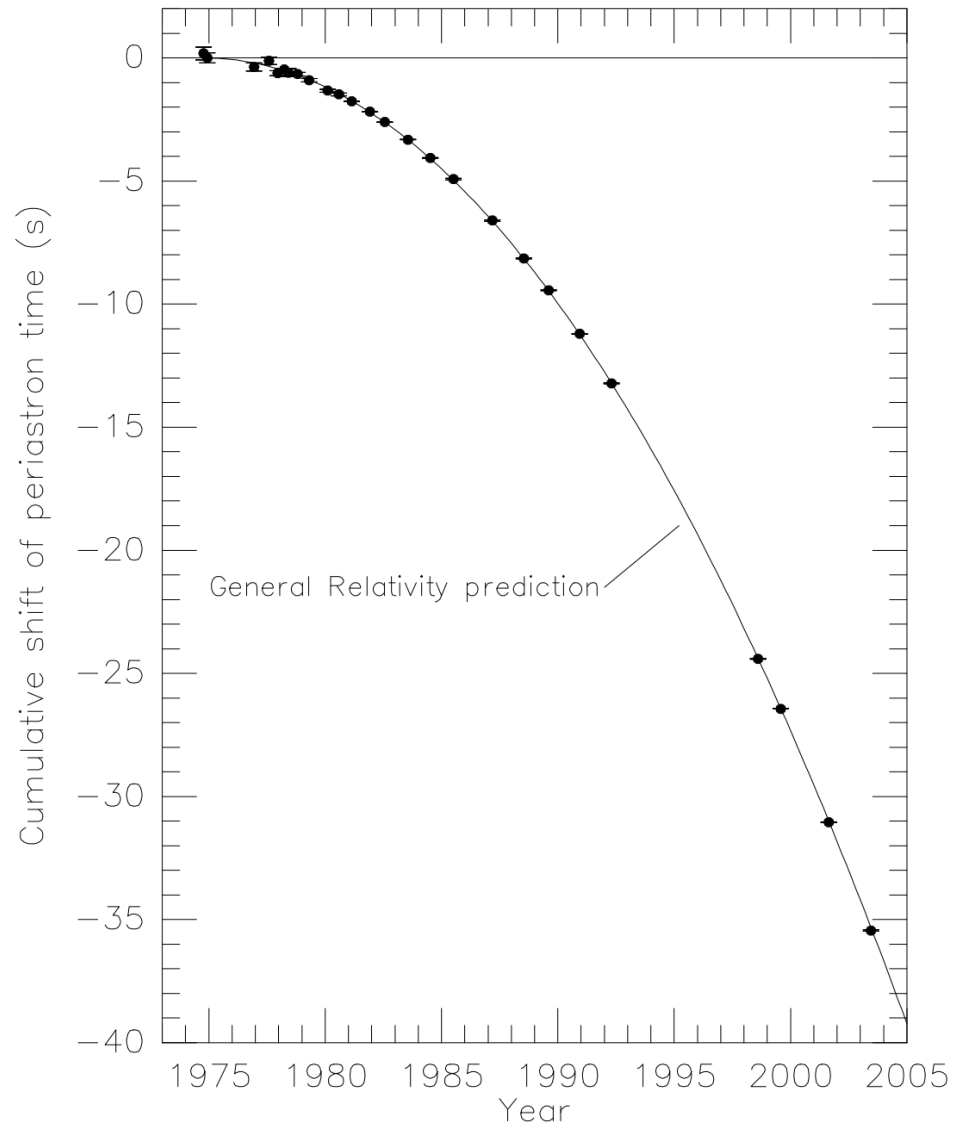
Hulse-Taylor binary

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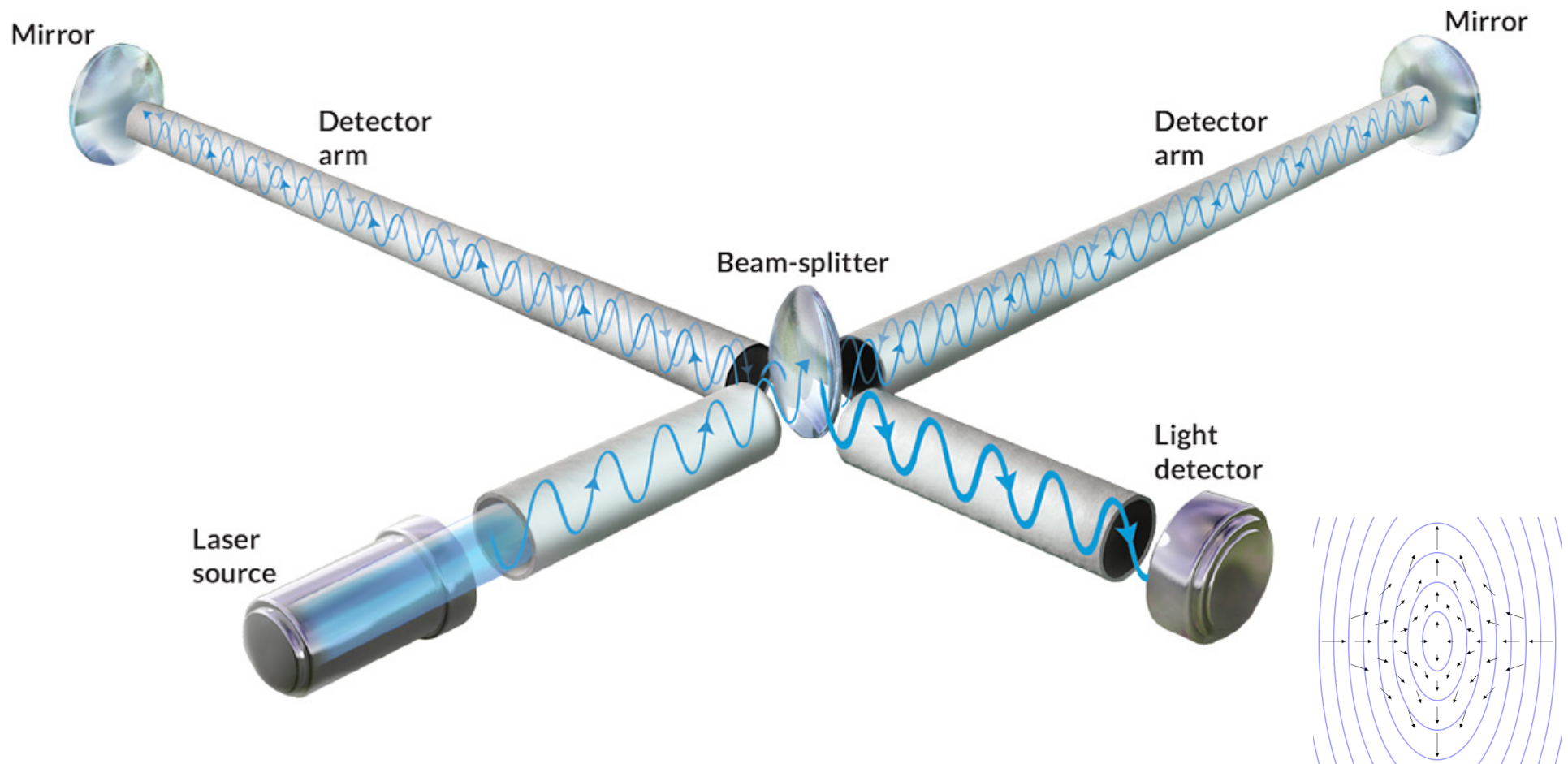
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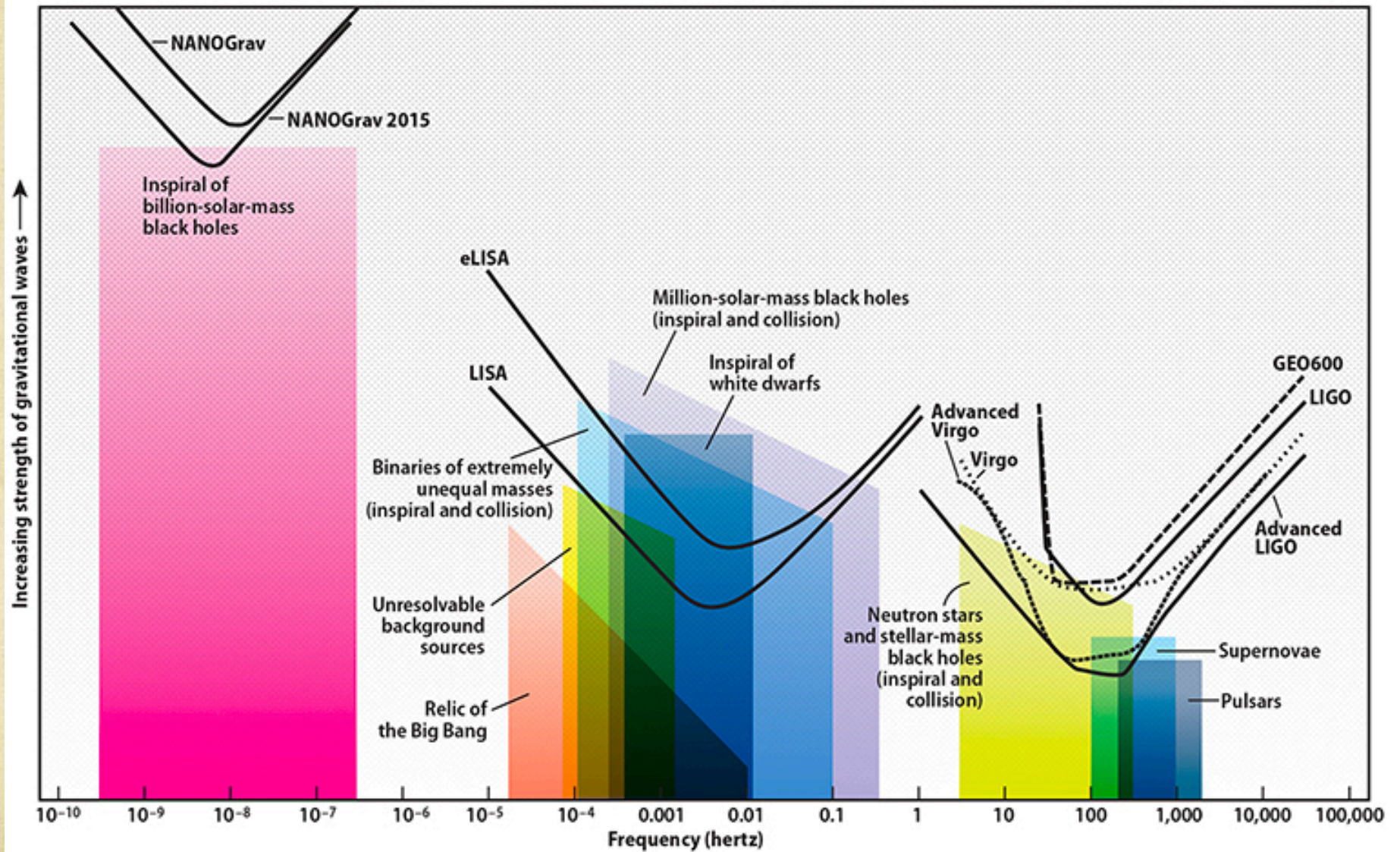
Detecting Gravitational Waves

Detecting Gravitational Waves

Fabry-Perot-Michelson Interferometers



Detecting Gravitational Waves



Astronomy: Roen Kelly, after C. Moore, R. Cole, and C. Berry (Institute of Astronomy, Univ. of Cambridge)

Multi-messenger Astronomy

Motivation

Multi-messenger Astronomy

Motivation

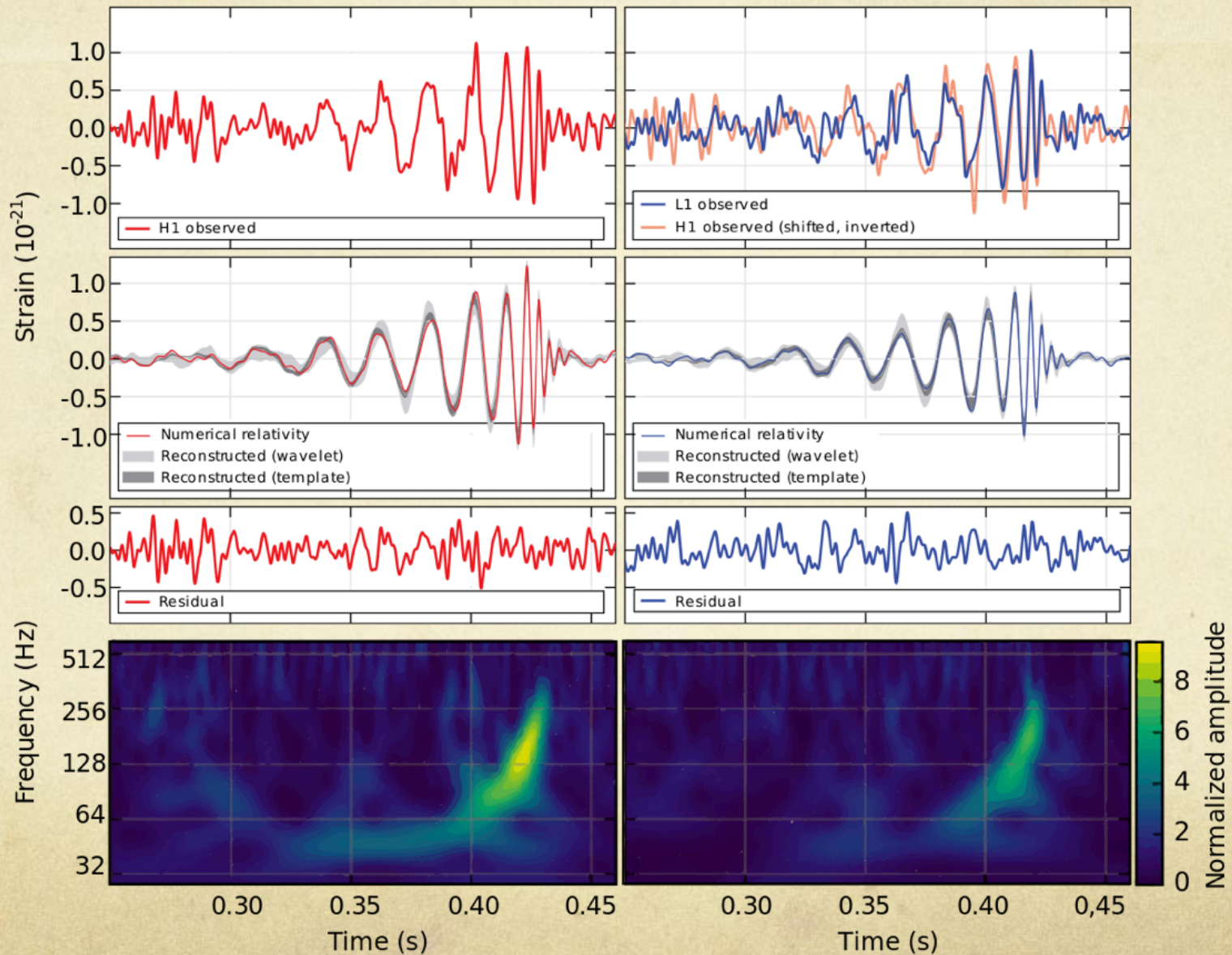
Gravitational Wave data only provides specific information

1. Precise Location/Distance Determination
 - Provides environmental information
 - Breaks degeneracy with inclination
2. Measure Influence/Test Merger Models
 - Origin of short-duration GRBs
 - Origin of r-process elements
3. Fundamental Physics
 - Speed of gravity
 - Lorentz invariance, equivalence principle, non-GR gravity
4. Cosmology
 - “Standard Siren” Hubble diagram

GW150914

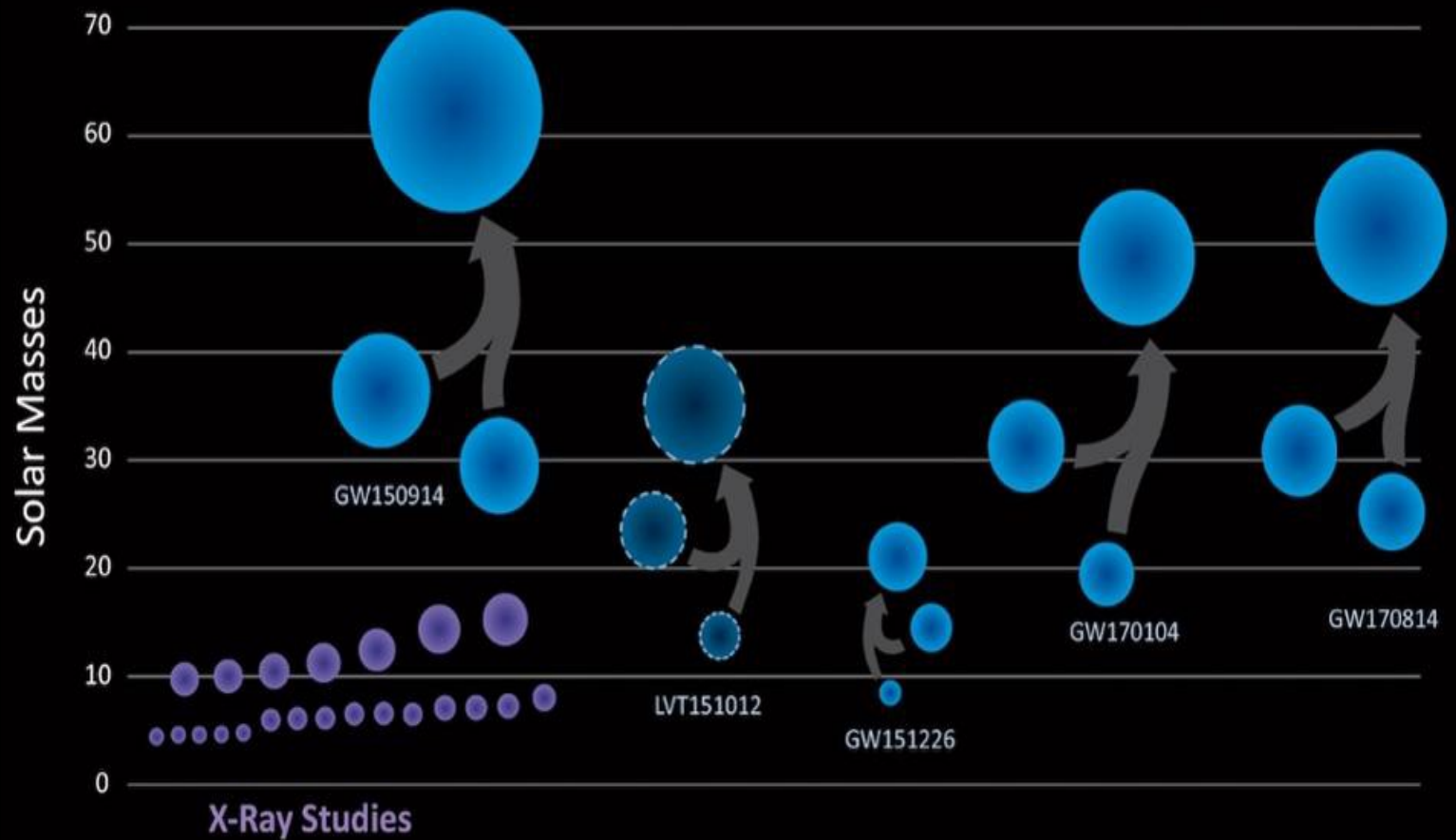
Hanford, Washington (H1)

Livingston, Louisiana (L1)



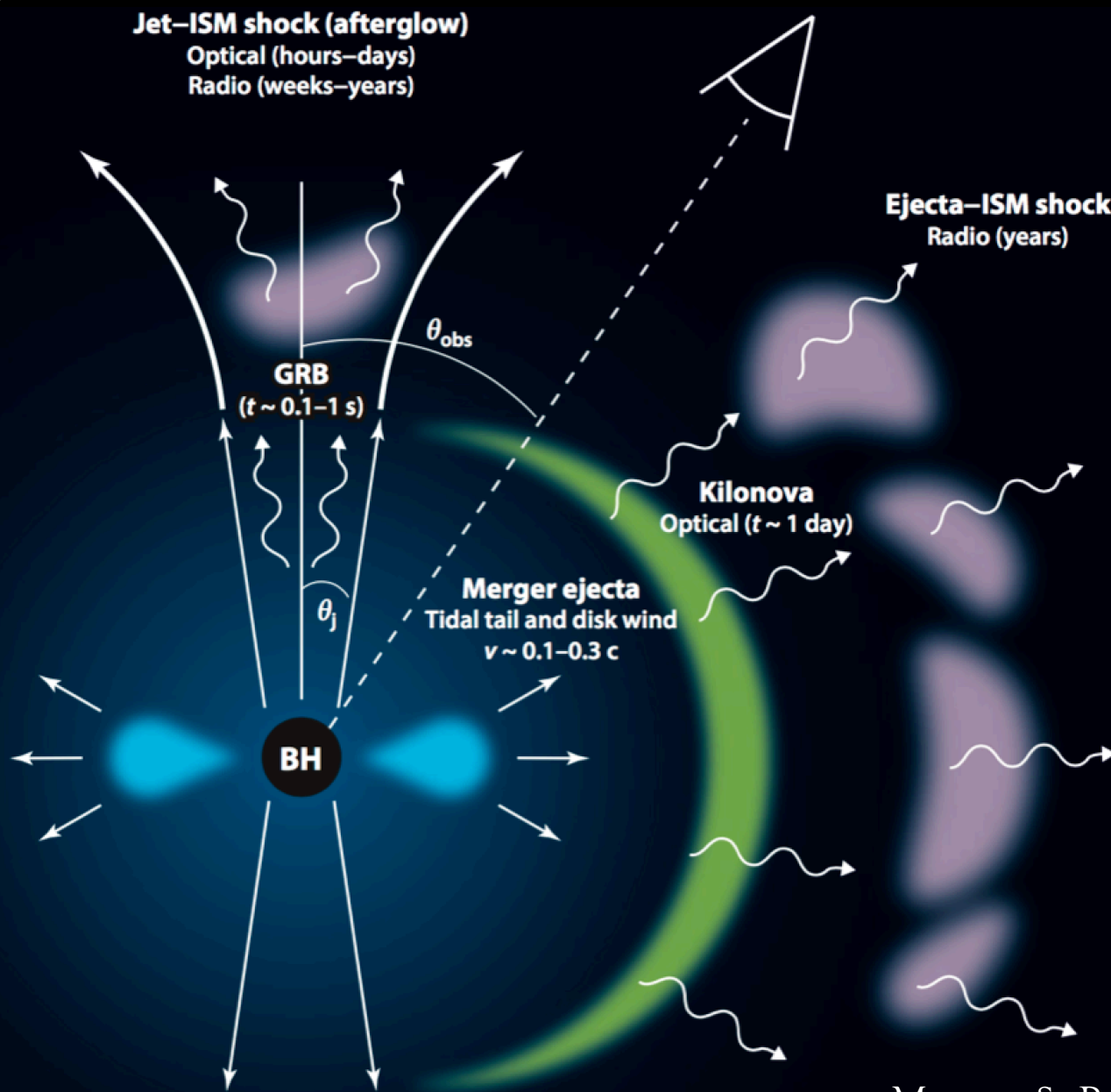
GW150814

Black Holes of Known Mass



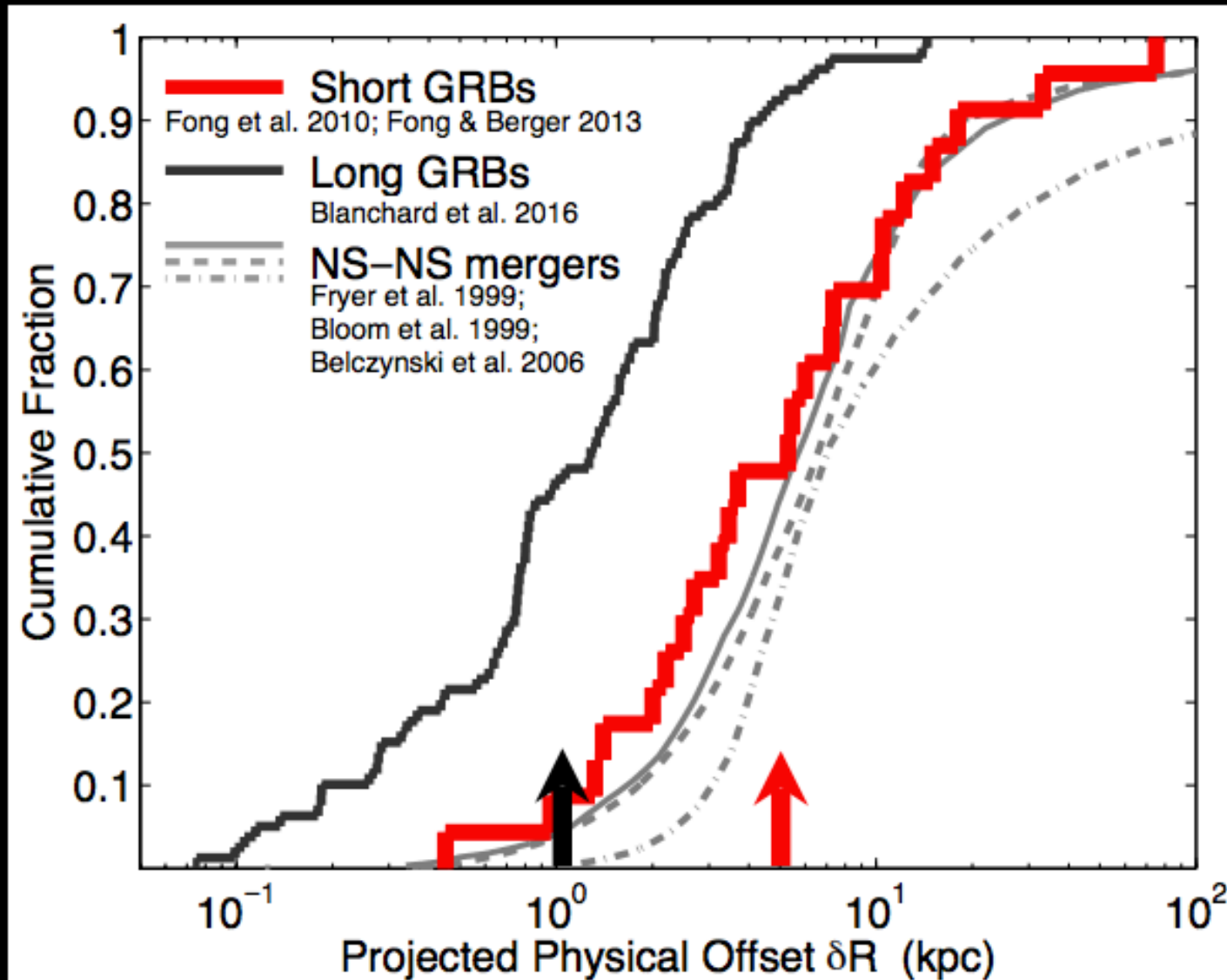
LIGO/VIRGO

EM Signals from Neutron-Star Mergers



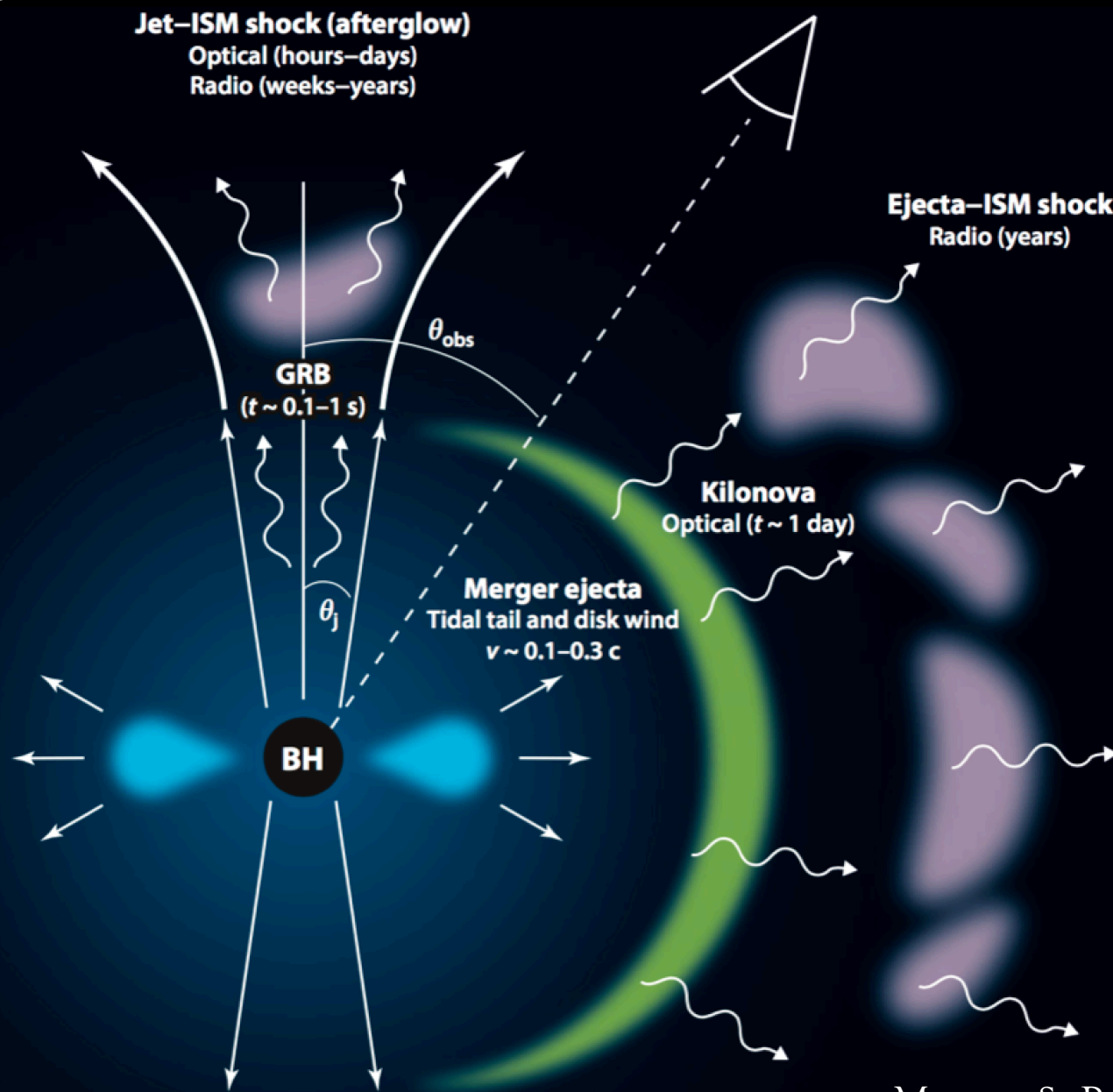
Metzger & Berger (2012)

Short-duration GRBs as NS mergers



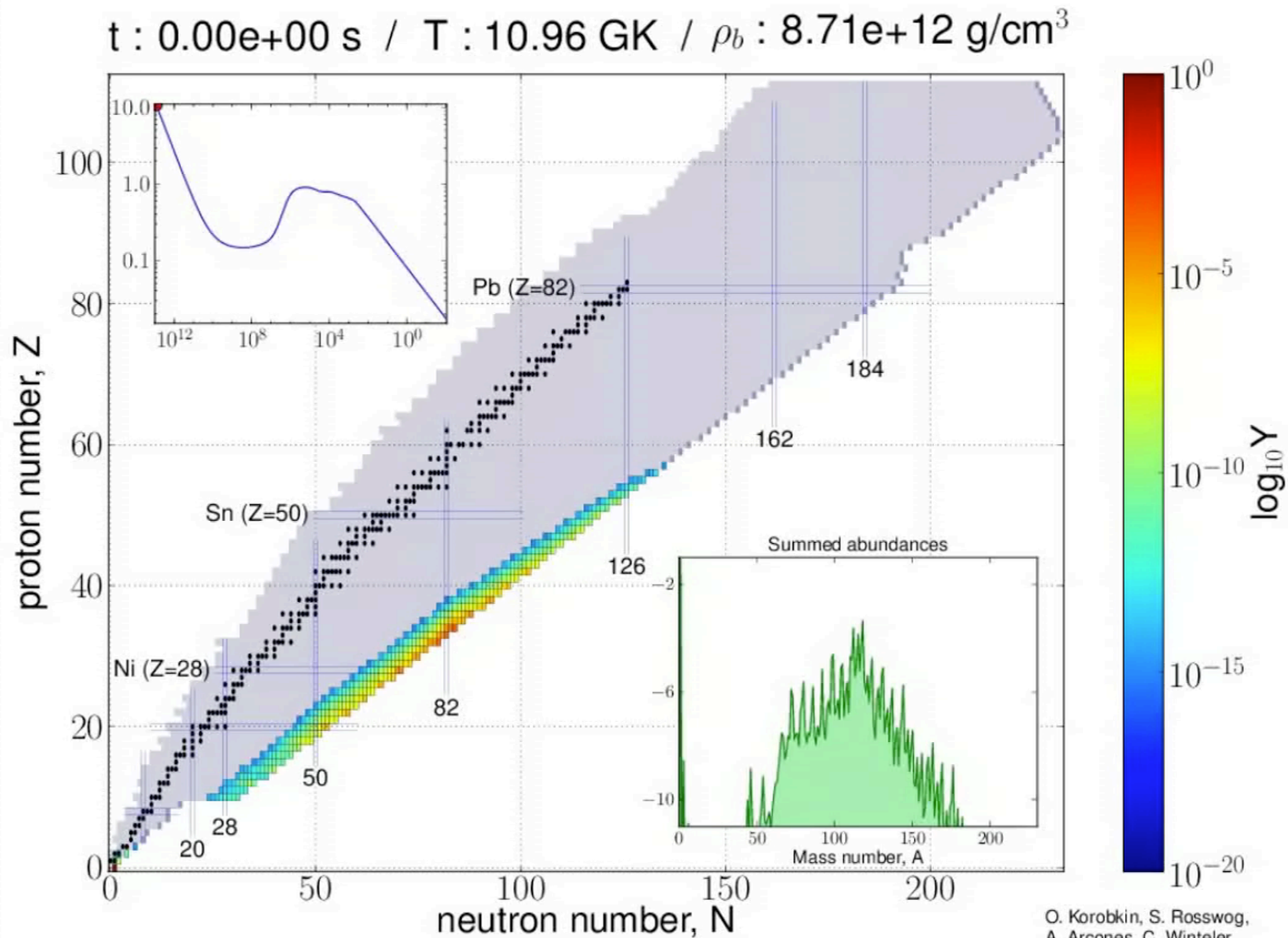
Fong et al. 2010; Fong & Berger 2013

EM Signals from Neutron-Star Mergers



Metzger & Berger (2012)

r-process nucleosynthesis



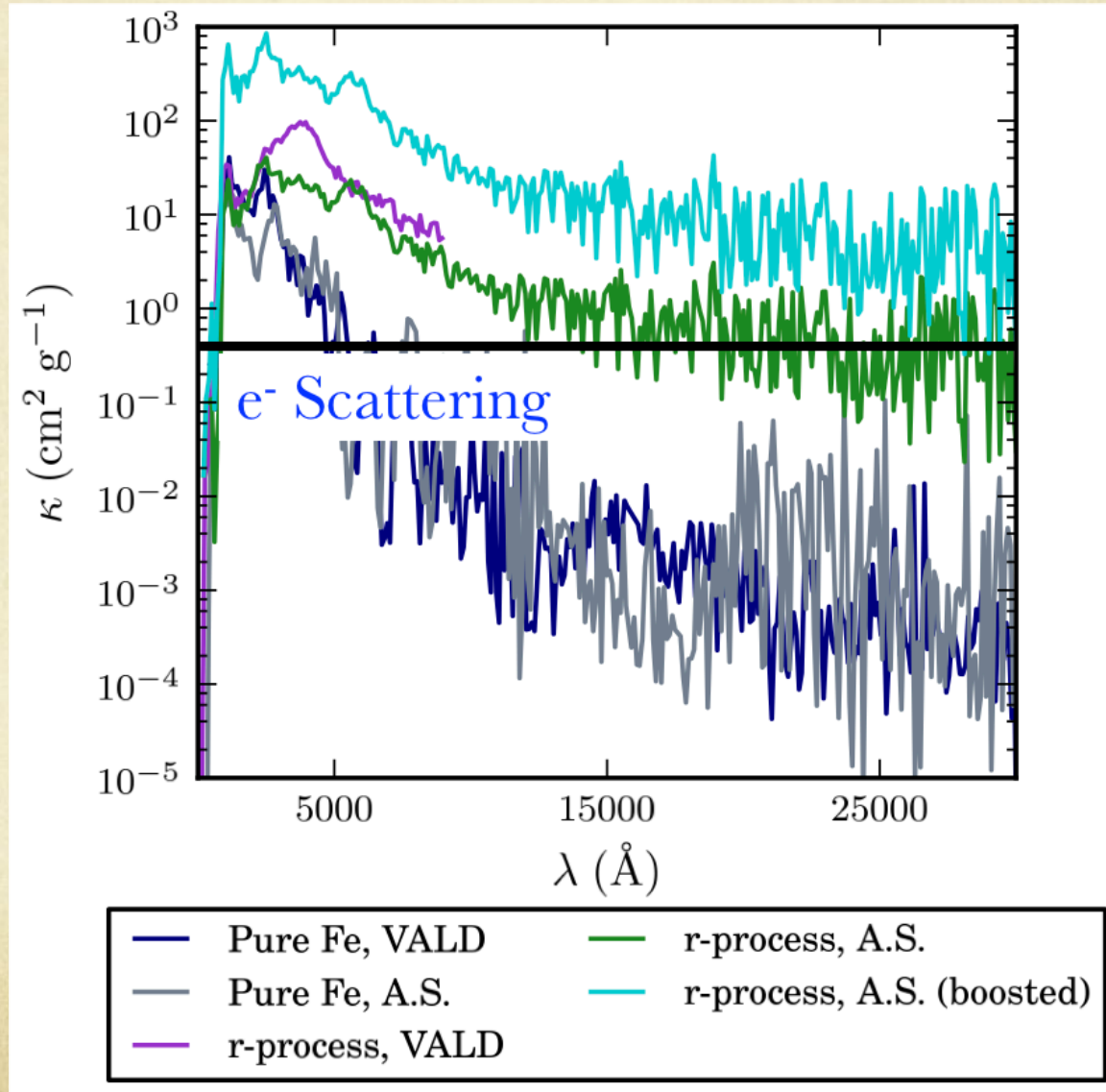
O. Korobkin, S. Rosswog,
A. Arcones, C. Winteler,
arXiv:1206.2379

Kilonova

What will they look like?

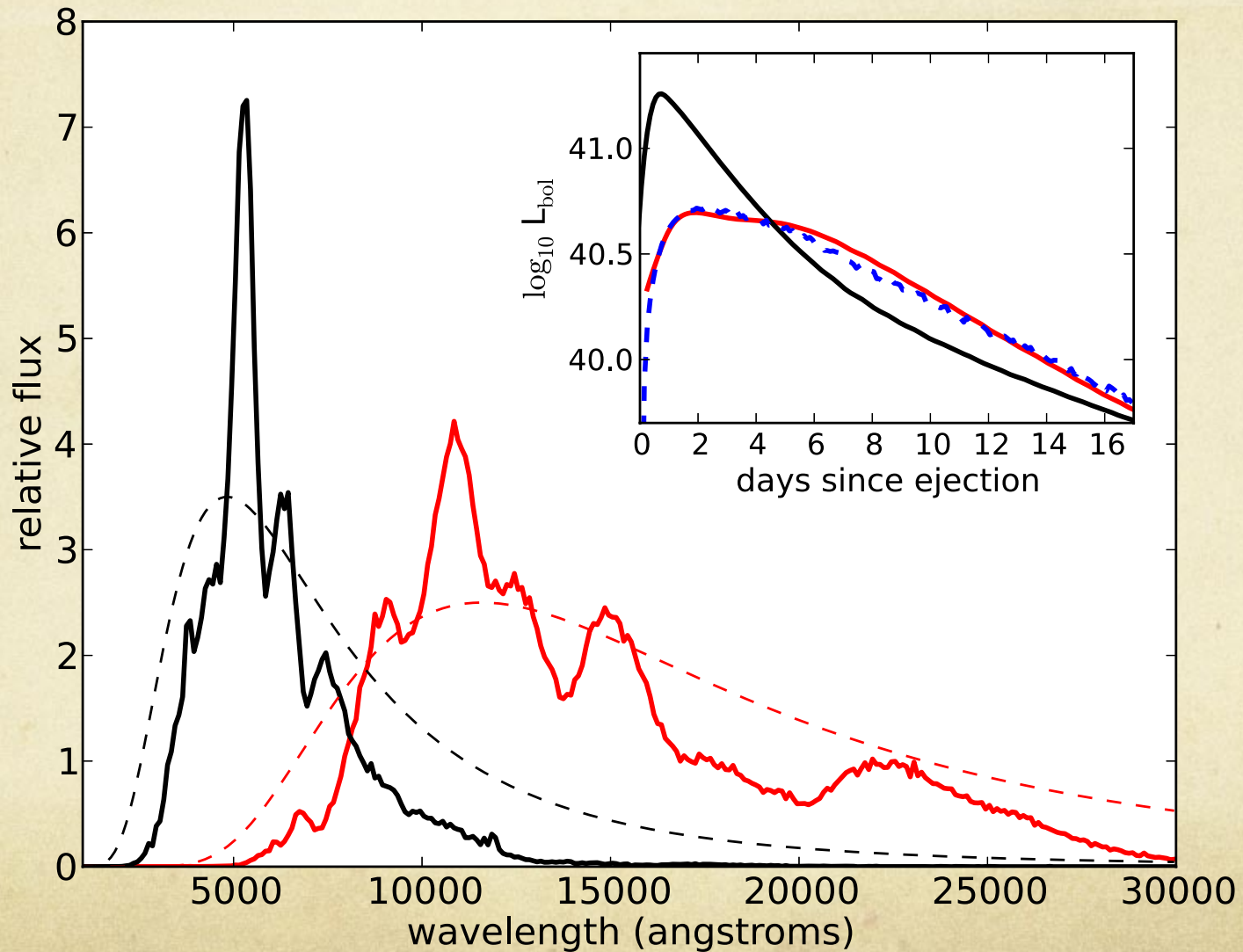
Kilonova

What will they look like?



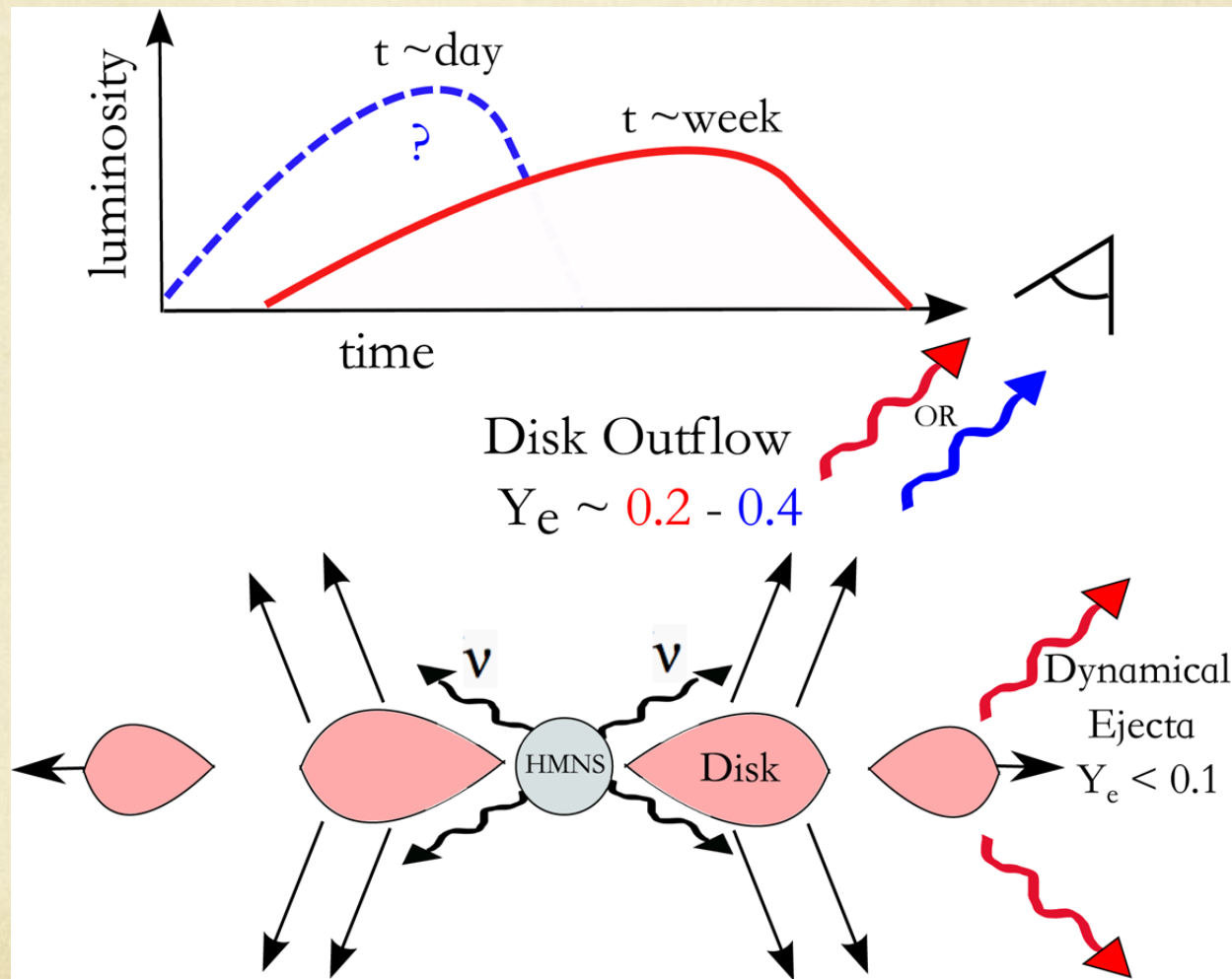
Kilonova

What will they look like?



Kilonova

What will they look like?



Multi-messenger Astronomy

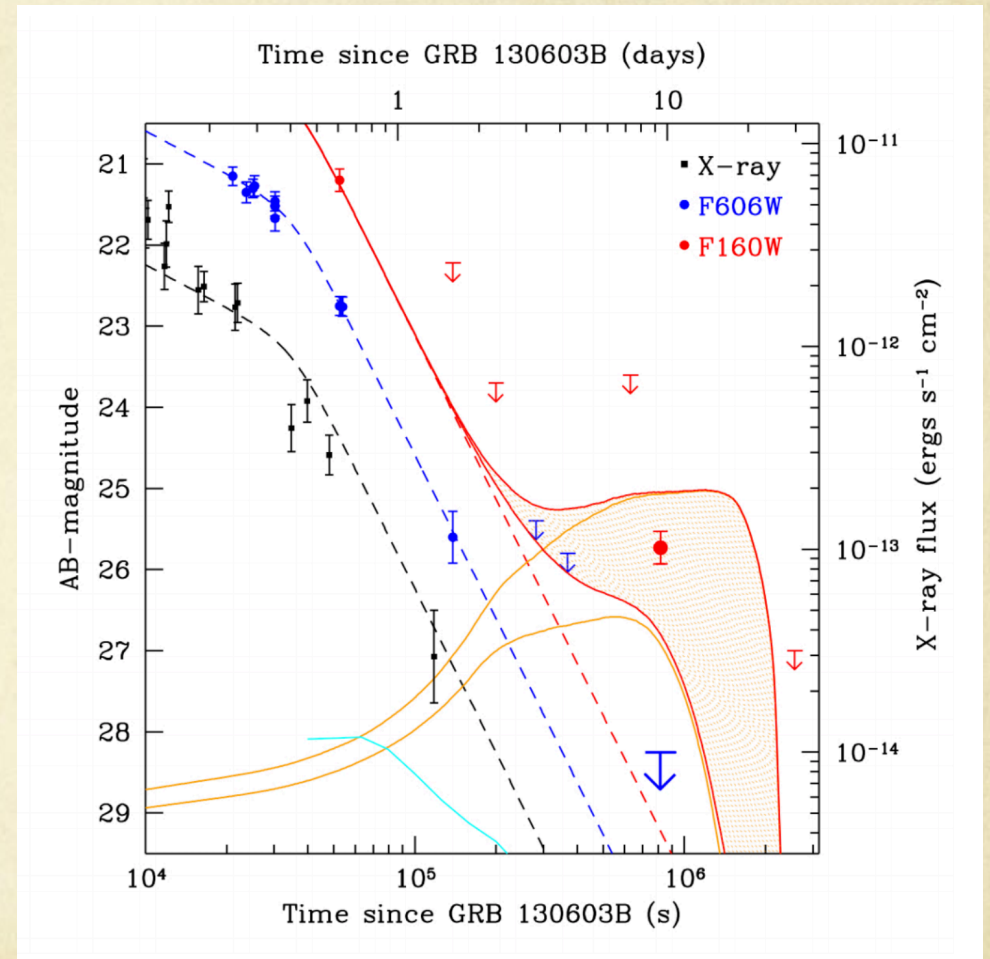
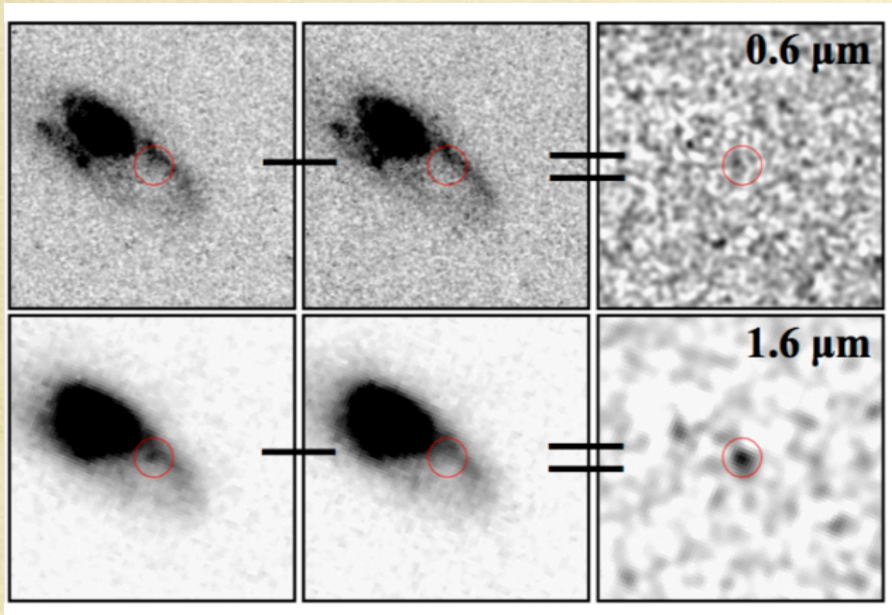
Motivation

Gravitational Wave data only provides specific information

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 - “Standard Siren” Hubble diagram

Aug 17th, 2017

GRB 130603B a Kilonova?



Tanvir et al. 2013, Berger et al. 2013

Aug 17th, 2017

Aug 17, 2017

//

TITLE: GCN CIRCULAR
NUMBER: 21505
SUBJECT: LIGO/Virgo G298048: Fermi GBM trigger 524666471/170817529: LIGO/Virgo Identification of a possible gravitational-wave counterpart
DATE: 17/08/17 13:21:42 GMT
FROM: Reed Clasey Essick at MIT <ressick@mit.edu>

The LIGO Scientific Collaboration and the Virgo Collaboration report:

The online CBC pipeline (gstlal) has made a preliminary identification of a GW candidate associated with the time of Fermi GBM trigger 524666471/170817529 at gps time 1187008884.47 (Thu Aug 17 12:41:06 GMT 2017) with RA=186.62deg Dec=-48.84deg and an error radius of 17.45deg.

The candidate is consistent with a neutron star binary coalescence with False Alarm Rate of $\sim 1/10,000$ years.

An offline analysis is ongoing. Any significant updates will be provided by a new Circular.

[GCN OPS NOTE(17aug17): Per author's request, the LIGO/VIRGO ID was added to the beginning of the Subject-line.]

Aug 17, 2017

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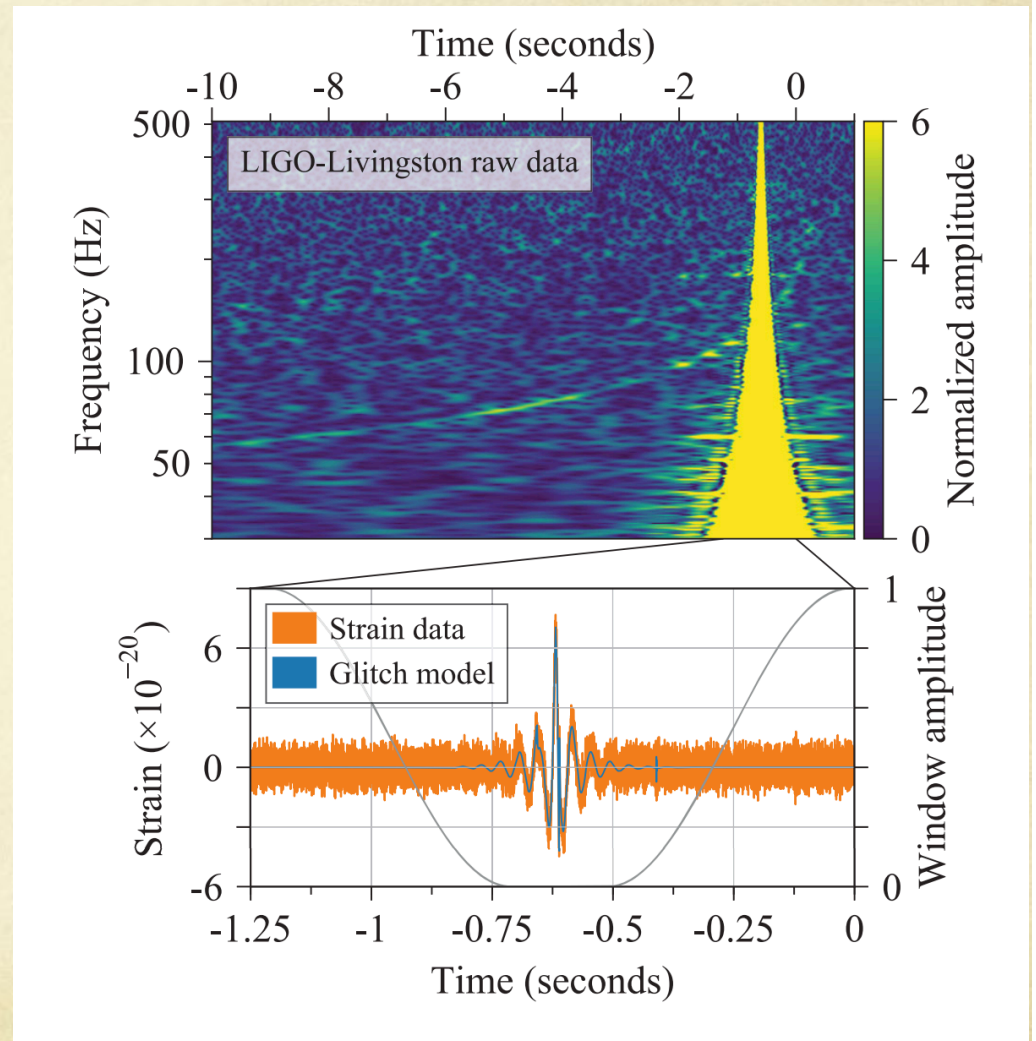
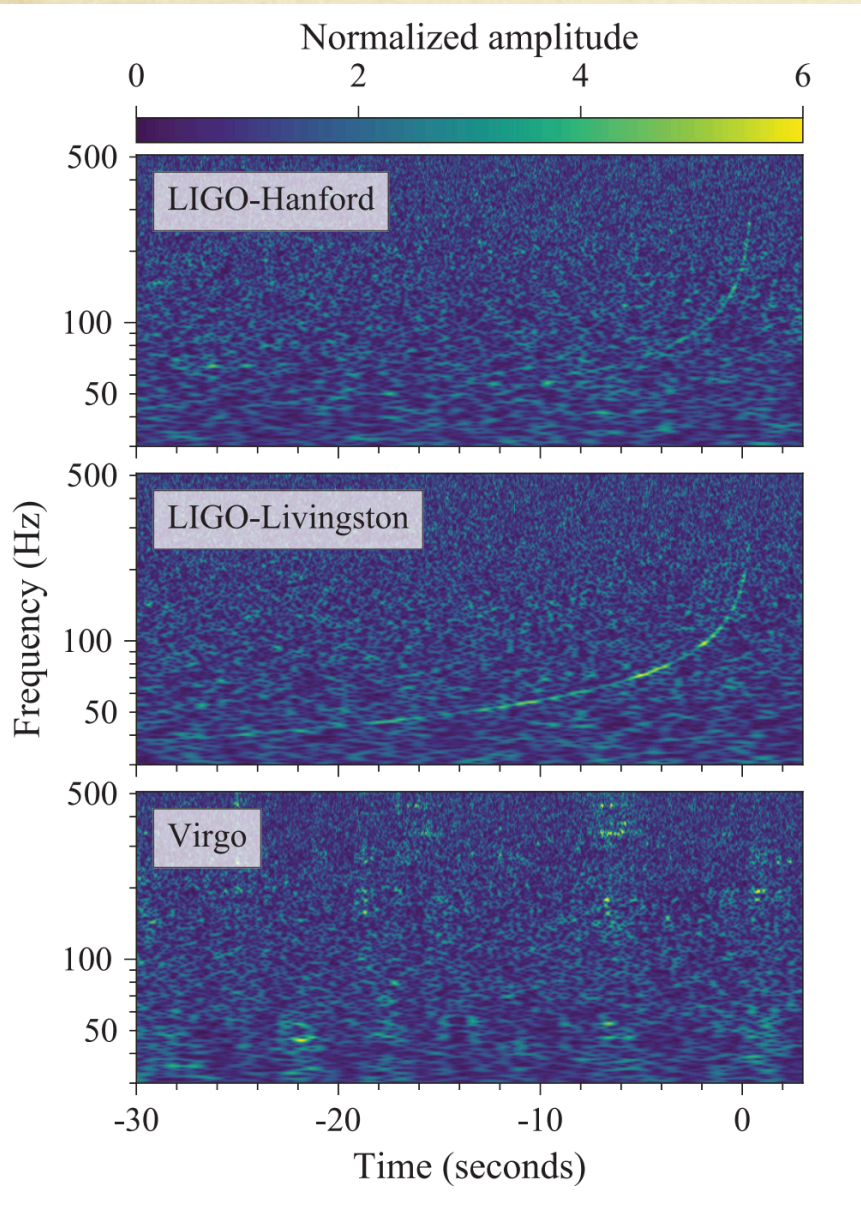
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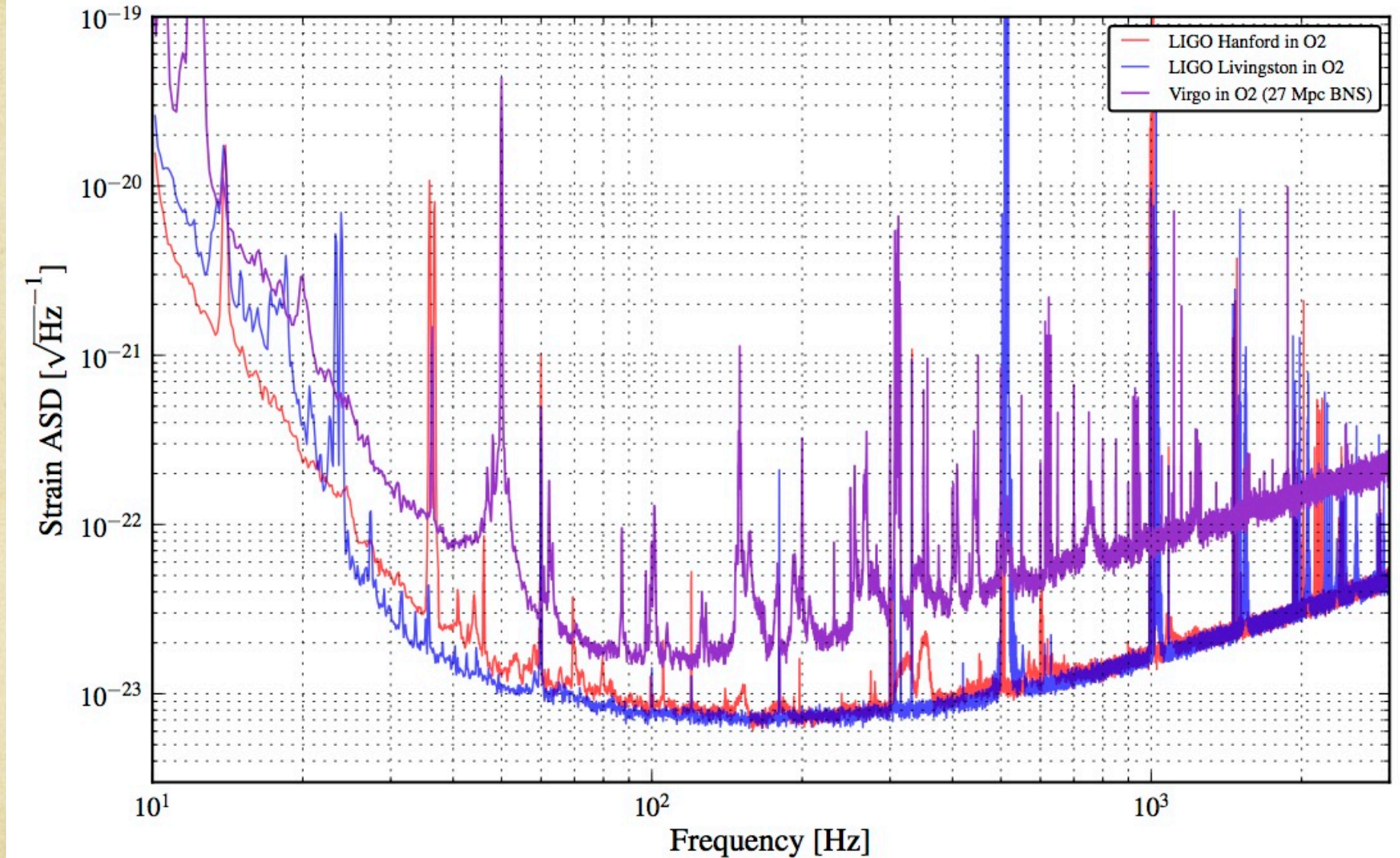
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Gravitational Waves



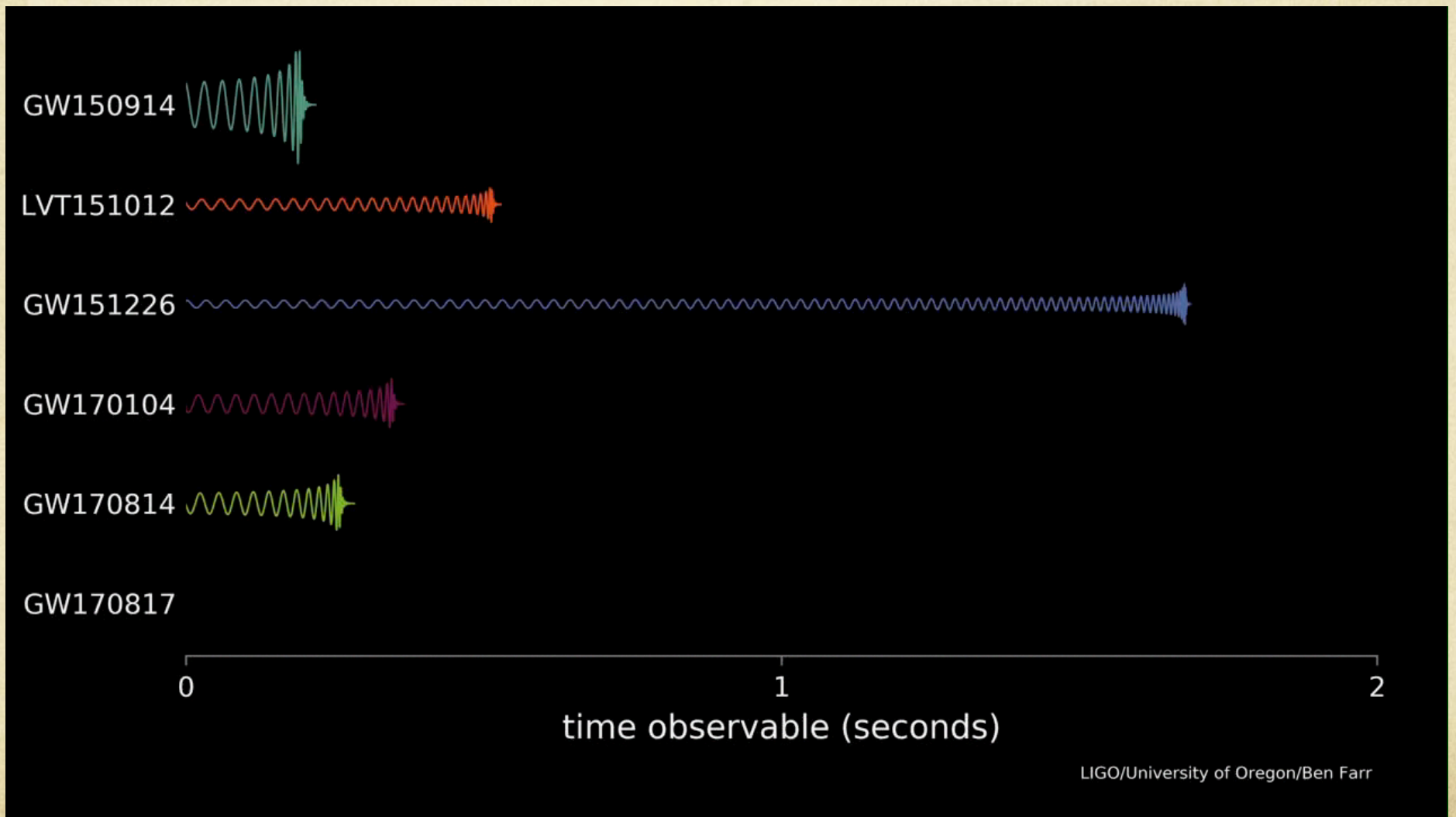
Abbot et al. 2017

LIGO + Virgo



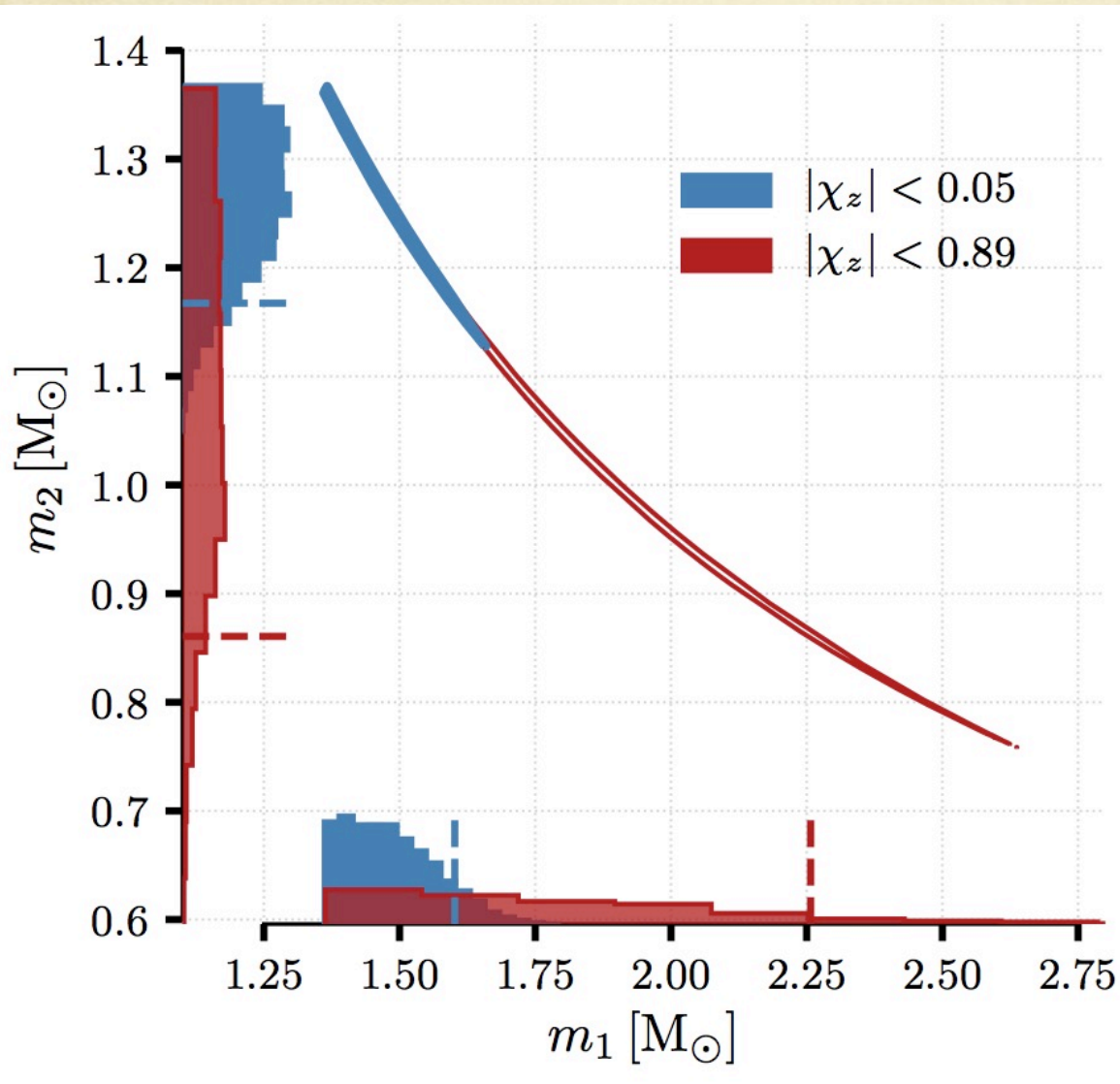
LIGO + Virgo

Gravitational Waves



Gravitational Waves

Component Masses



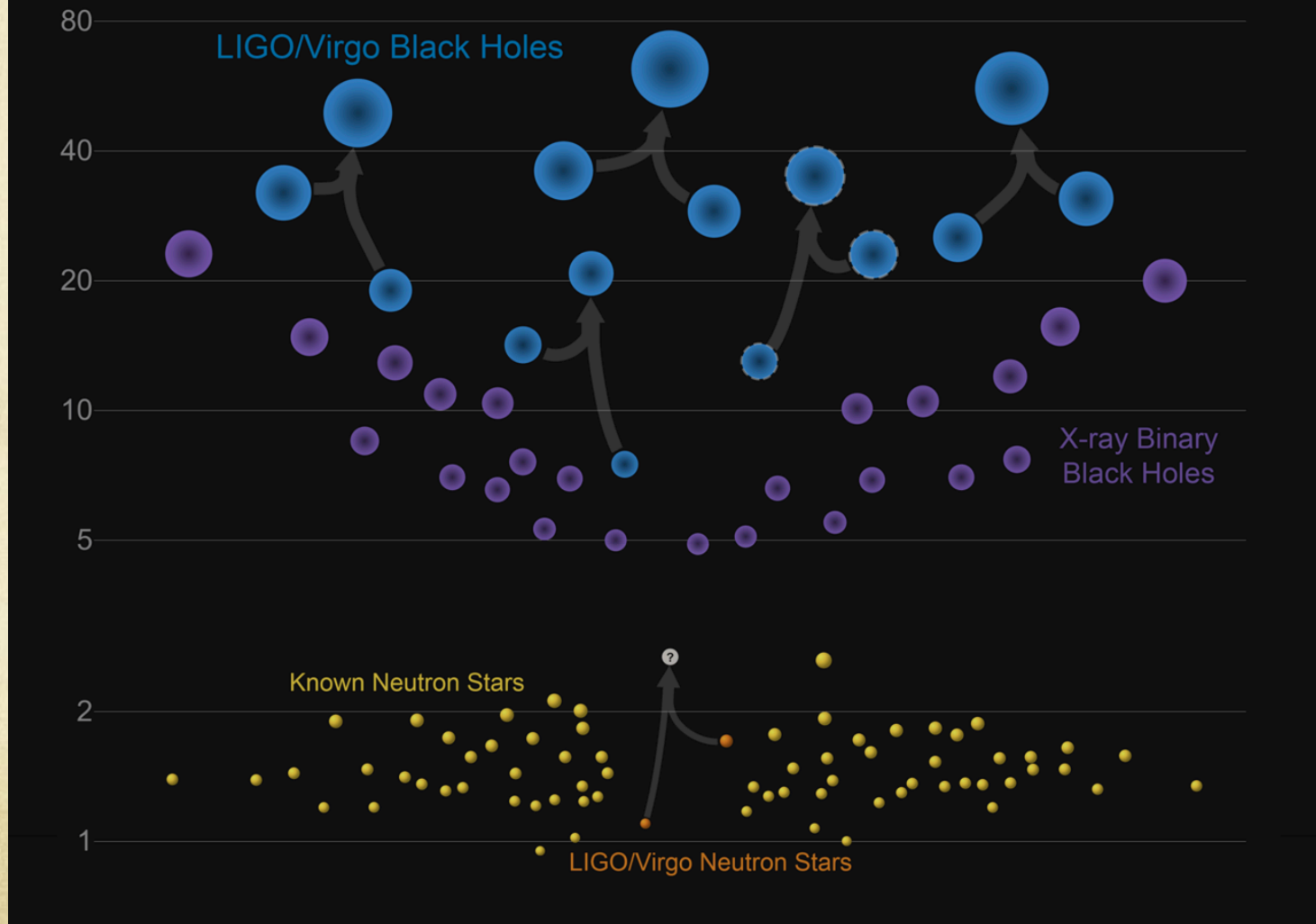
$m_1 = 1.36 - 1.60 M_{\text{sun}}$
 $m_2 = 1.17 - 1.36 M_{\text{sun}}$
 $M_{\text{tot}} = 2.74 M_{\text{sun}}$
Mass ratio = 0.7 - 1.0

Abbott et al (2017)

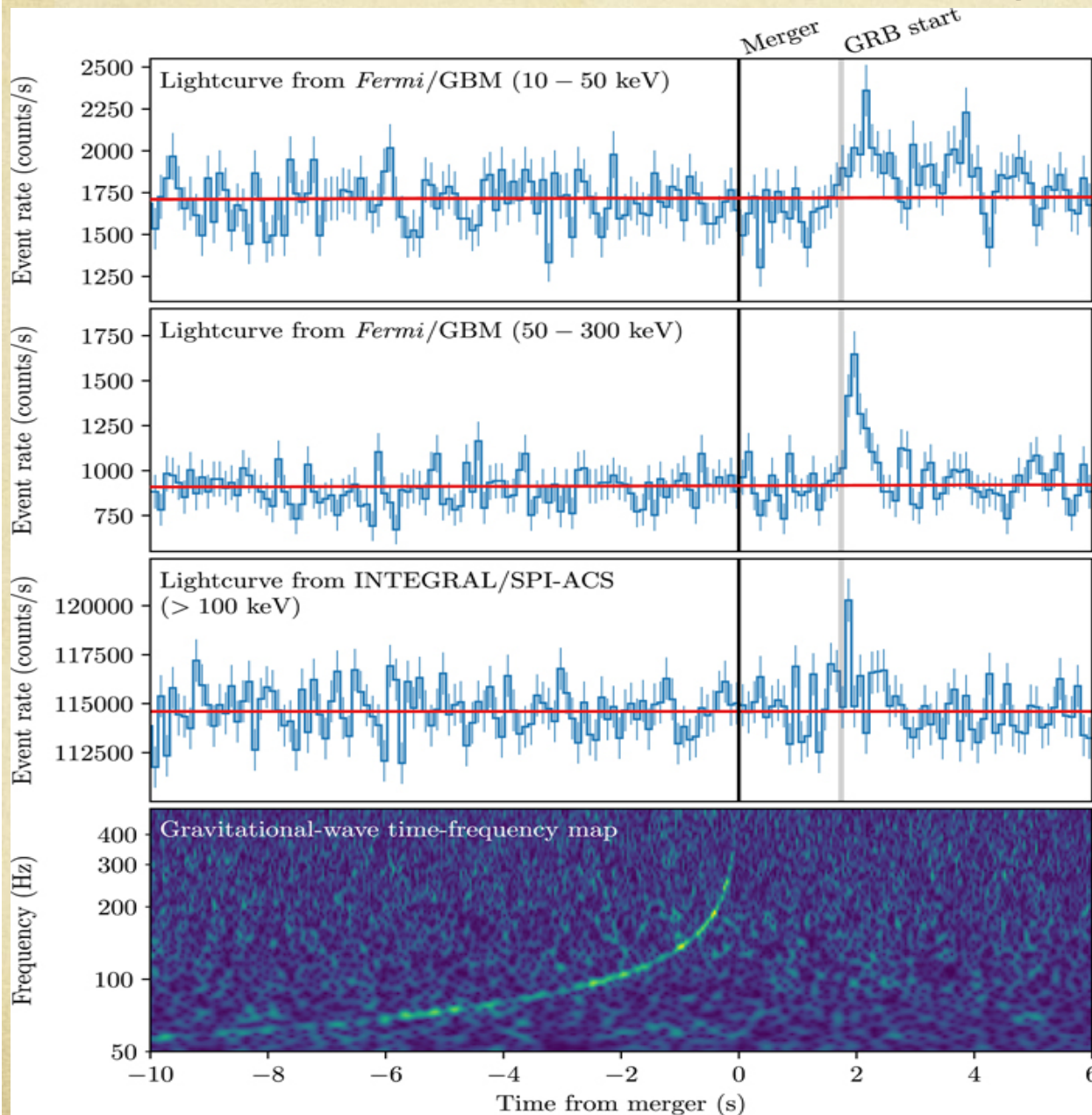
Gravitational Waves

Masses in the Stellar Graveyard

in Solar Masses



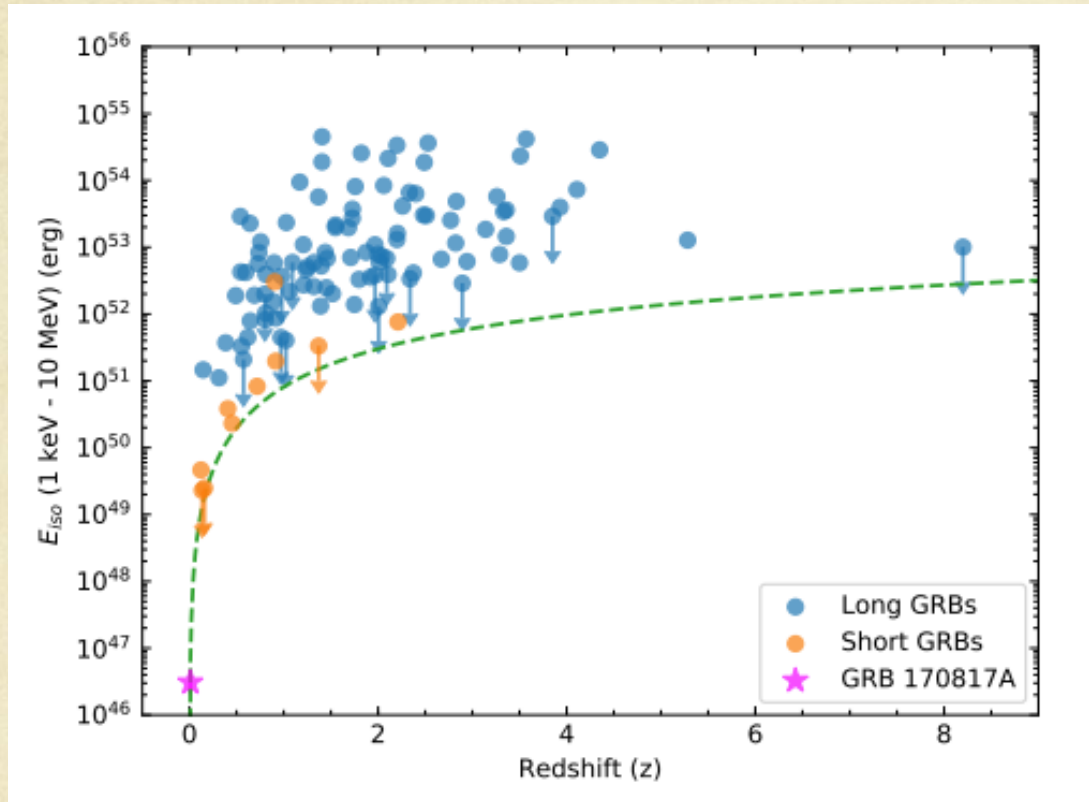
Gamma-rays



- 1.74 ± 0.05 s
- 4-6 orders of mag less than a typical Swift sGRB (Gehrels et al. 2009).
- X-ray emission 6-8 orders of mag less than typical Swift sGRB (Gehrels et al. 2009).

Abbot et al. 2017b

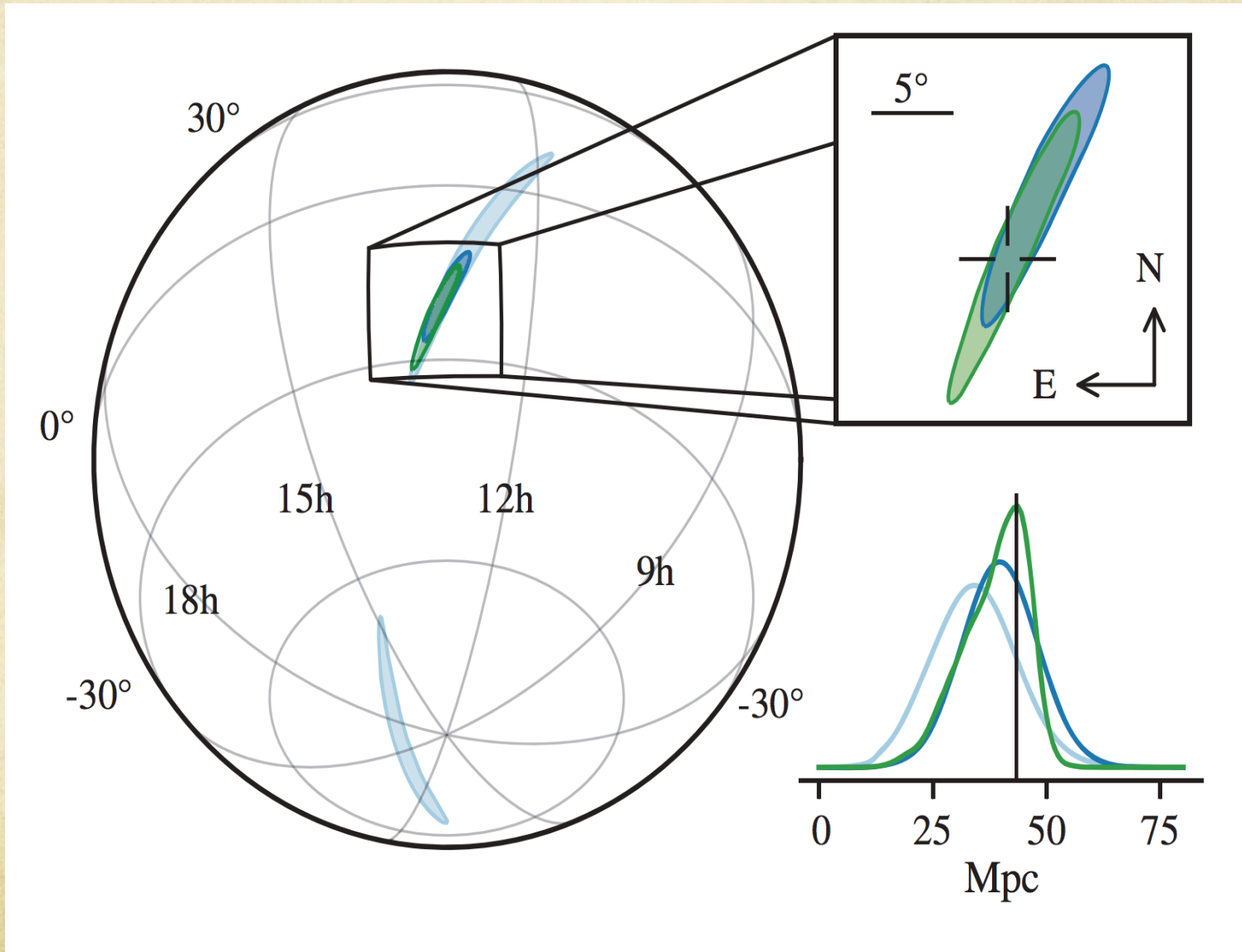
Gamma-rays



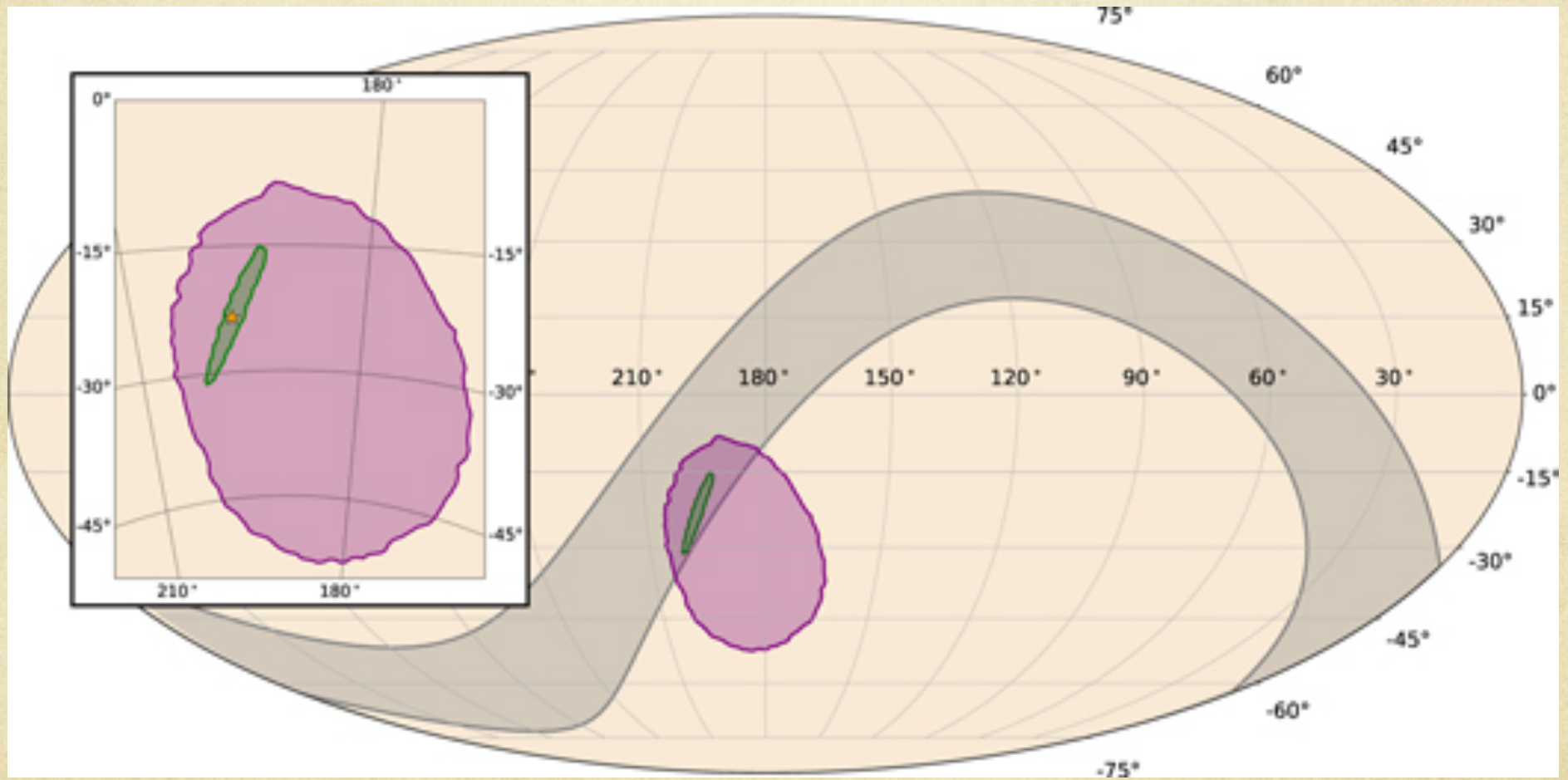
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Abbot et al. 2017b

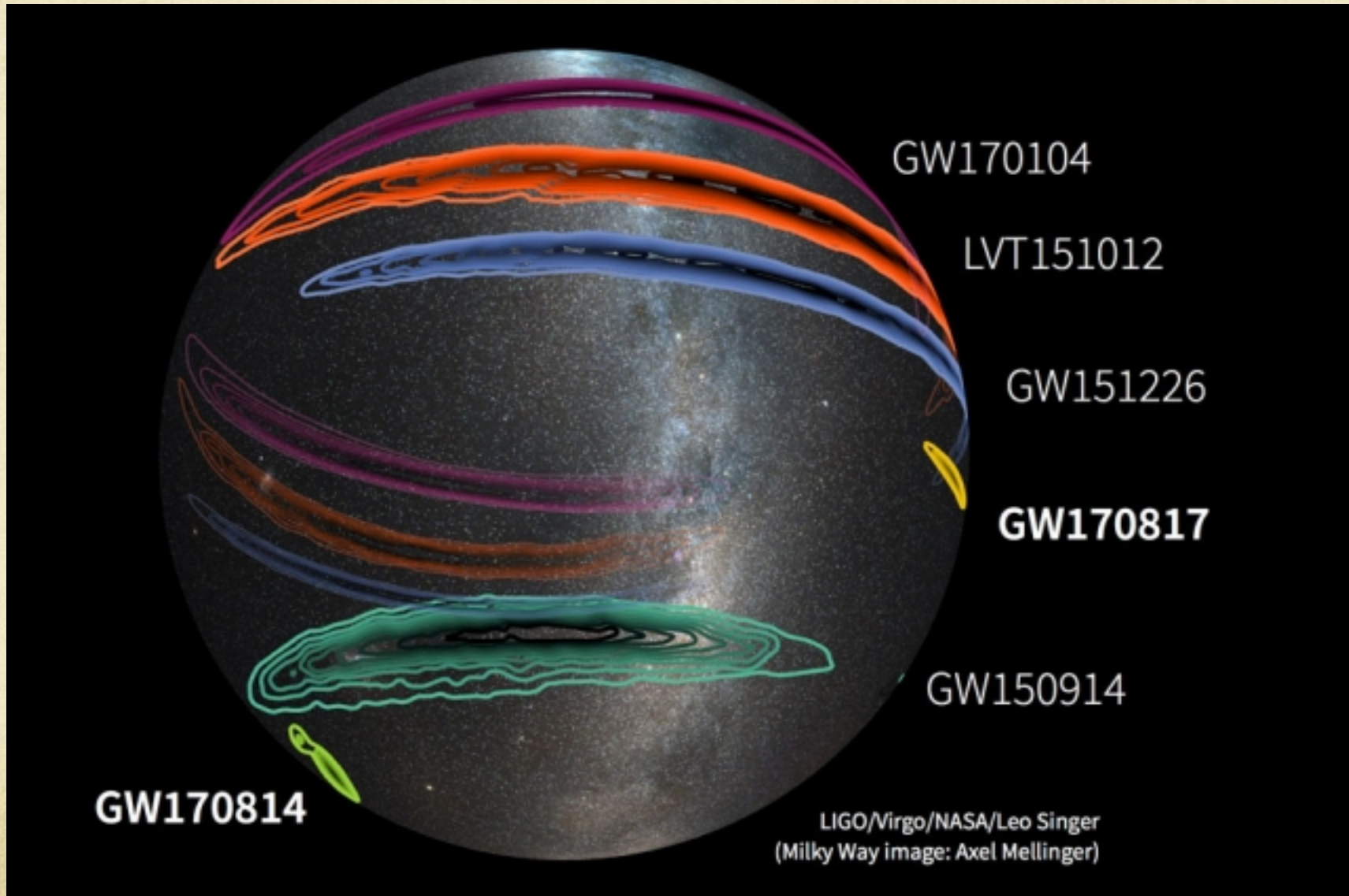
Sky Localization



Not a better localization



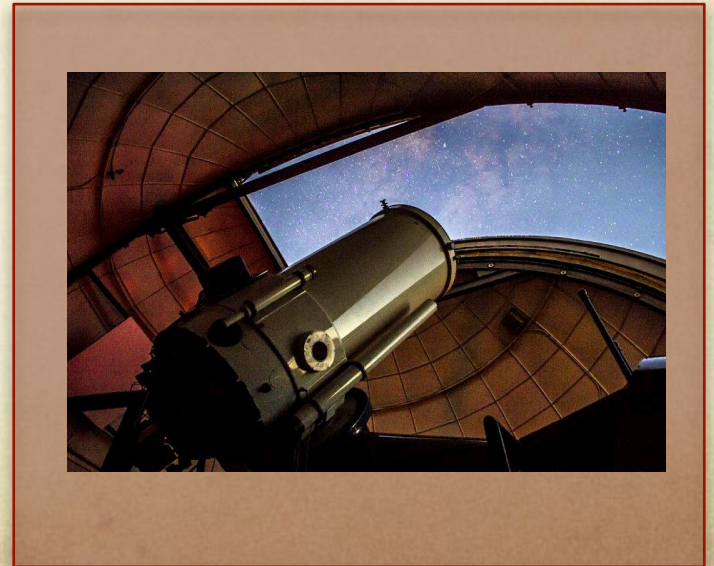
Sky Localization



1M2H



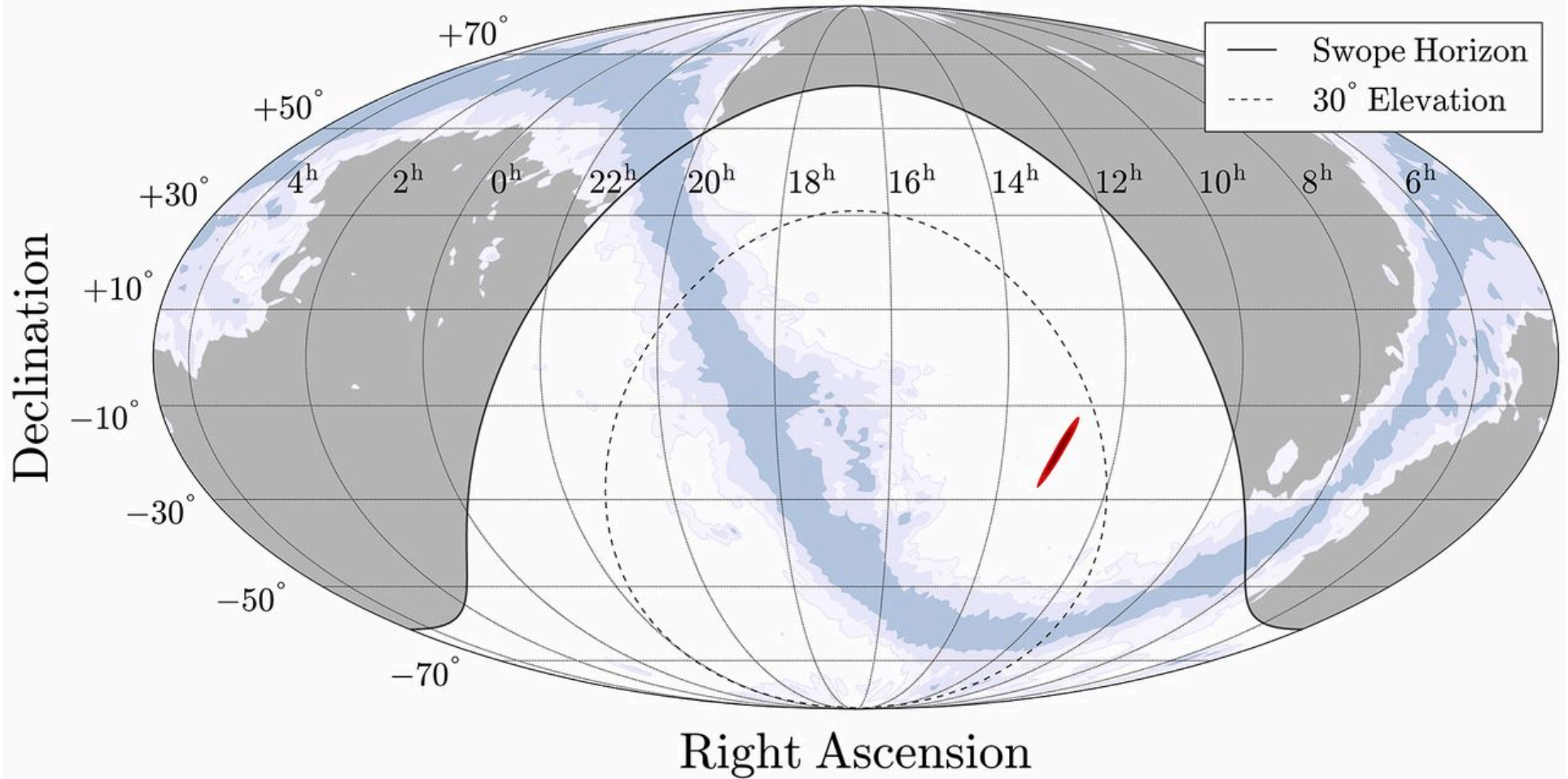
1M2H



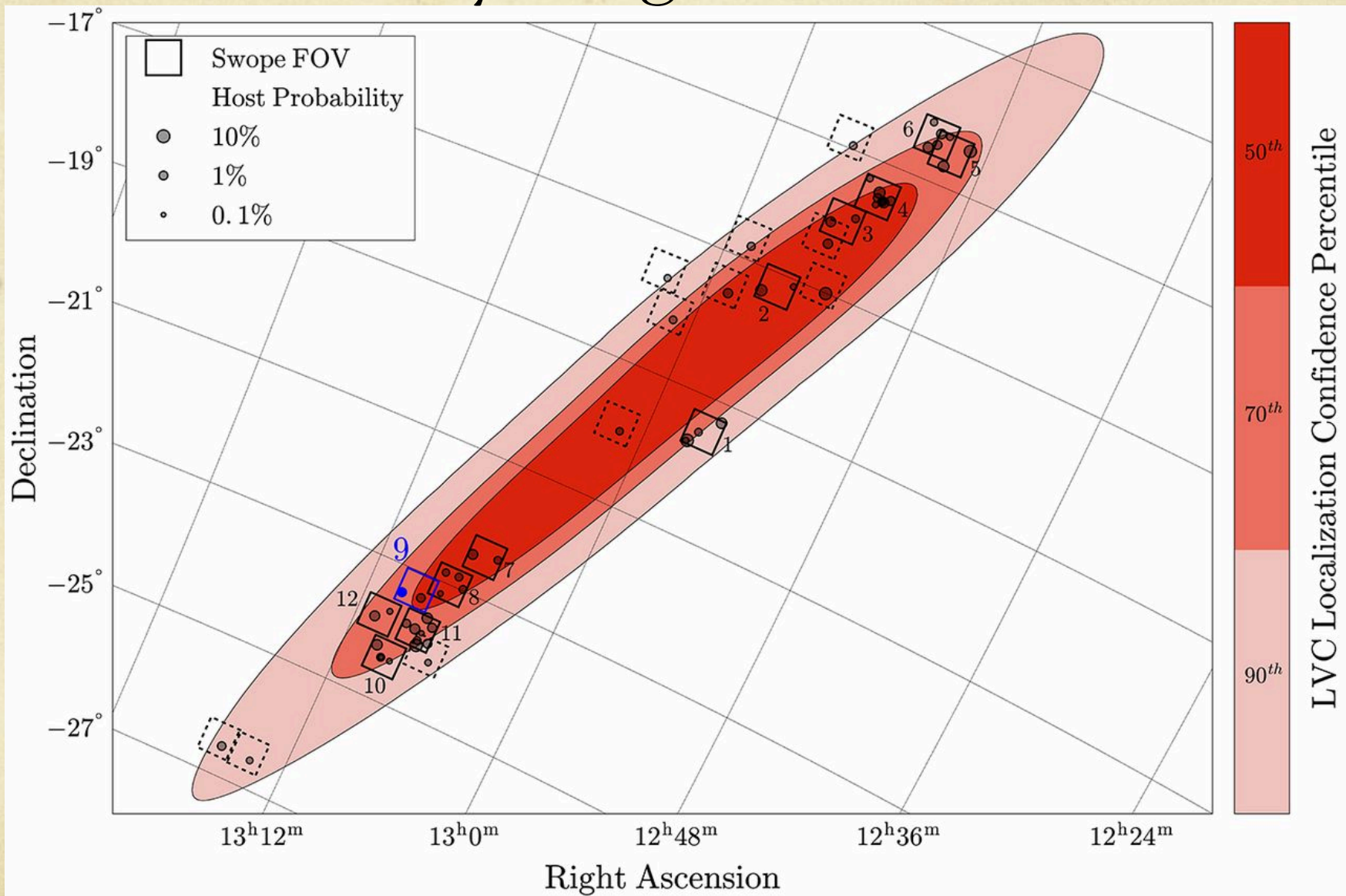
1M2H



1M2H

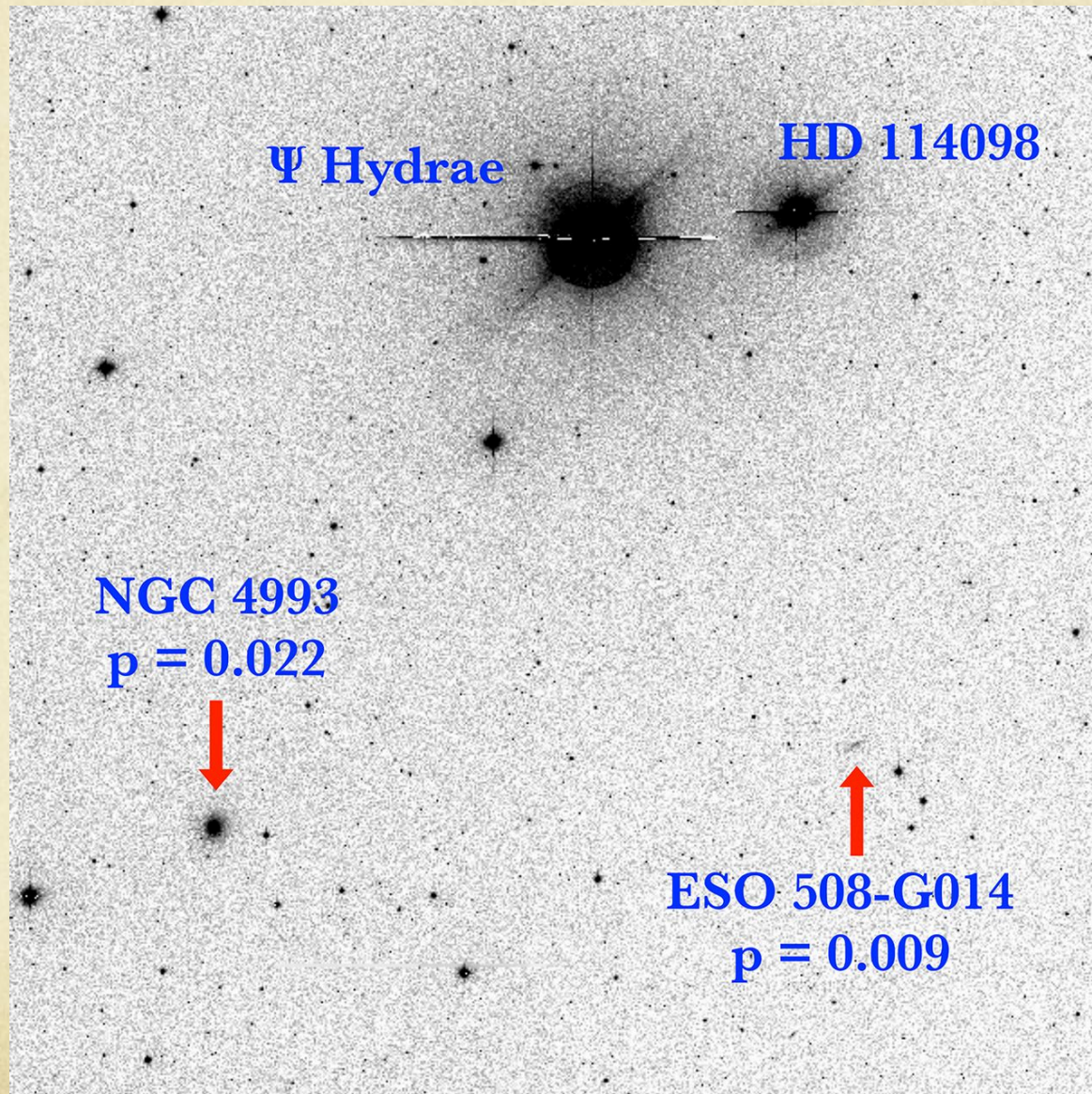


Galaxy-Targeted Search



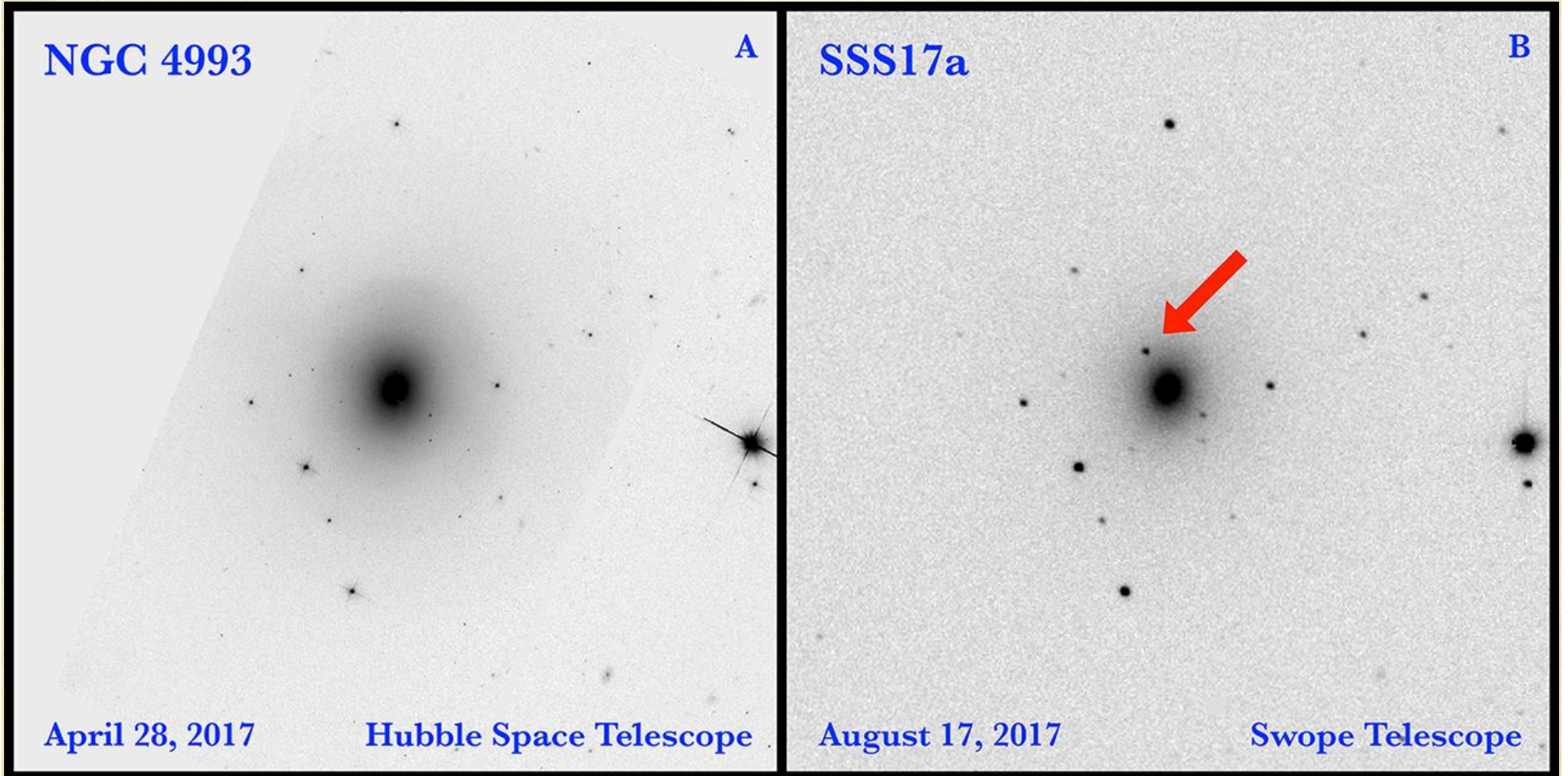
Coulter et al. 2017

The 9th Image

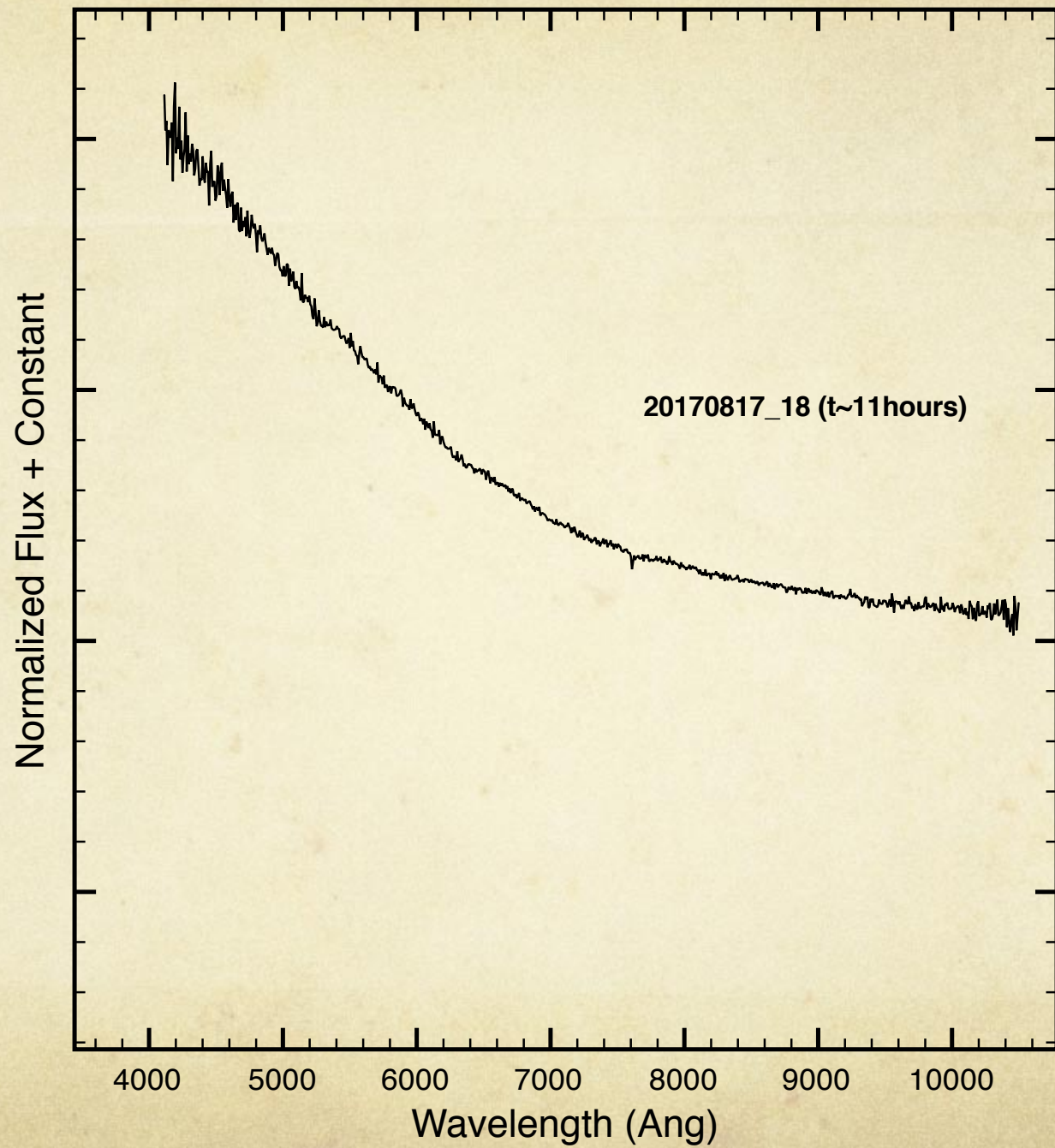


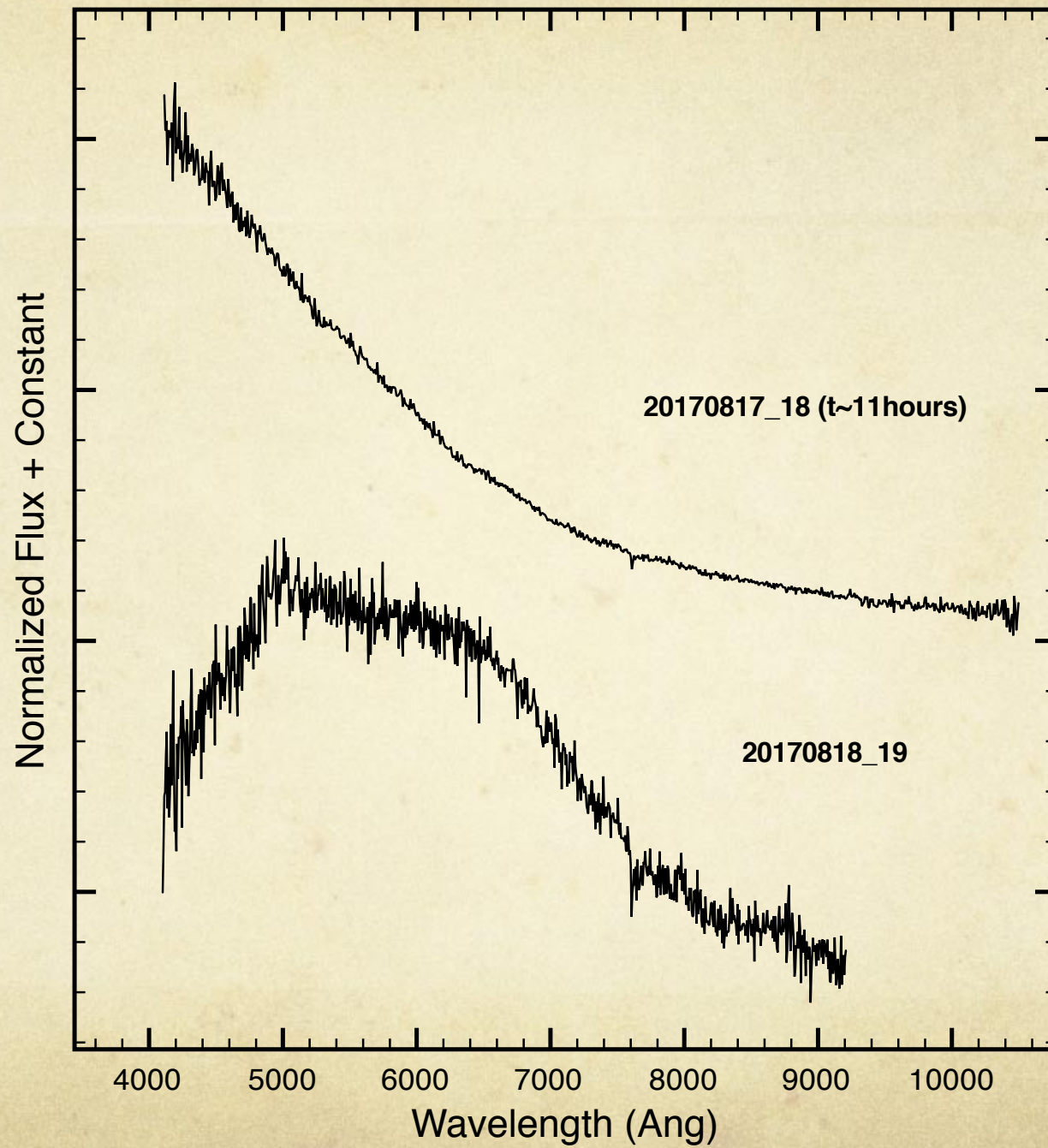
Coulter et al. 2017

SSS17a

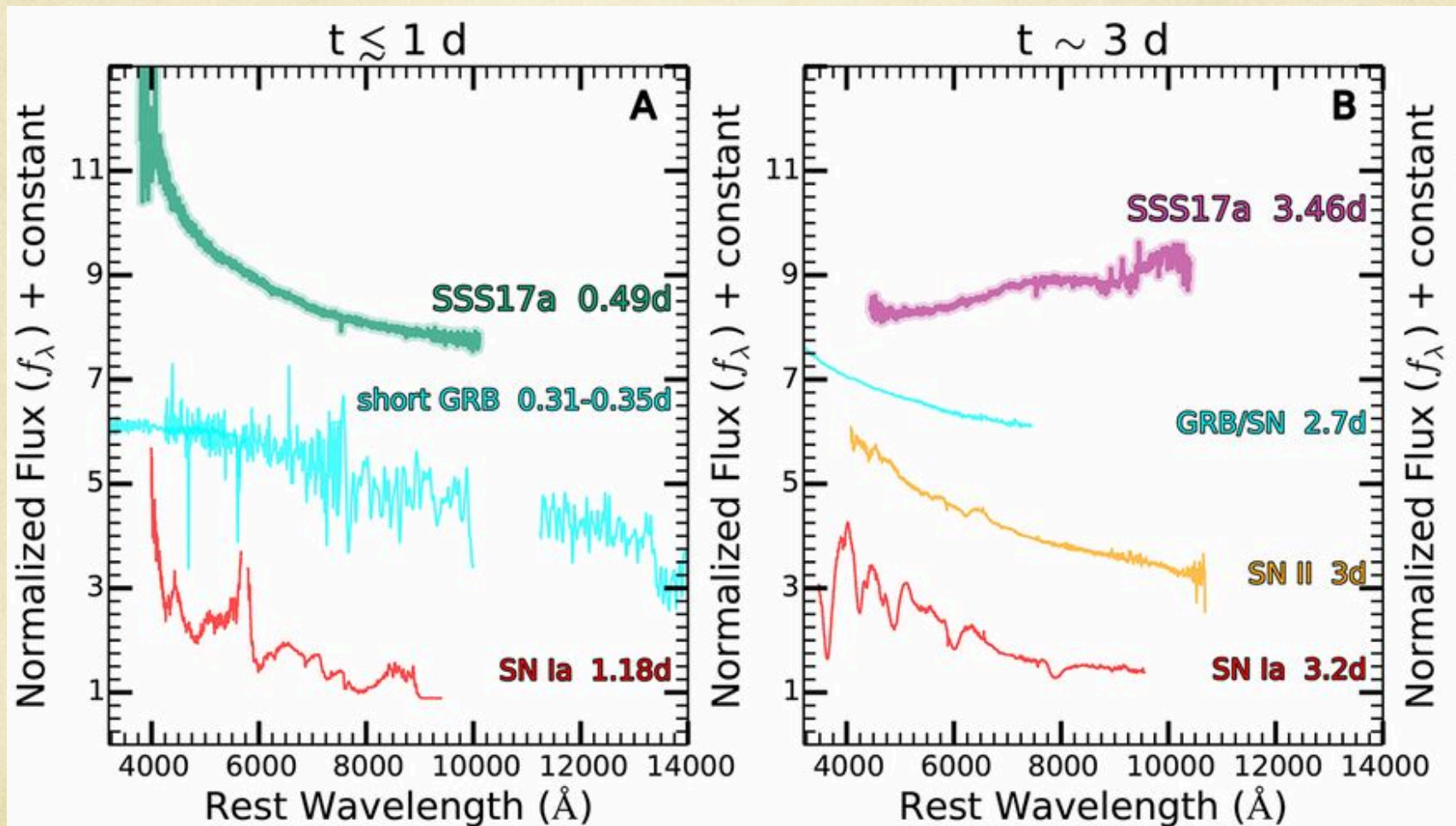


Coulter et al. 2017





SSS17a vs. Other Transients



SSS17a: Follow-Up



Swope 1-m:
Optical
Photometry



du Pont 2.5-m:
Near-IR
Photometry

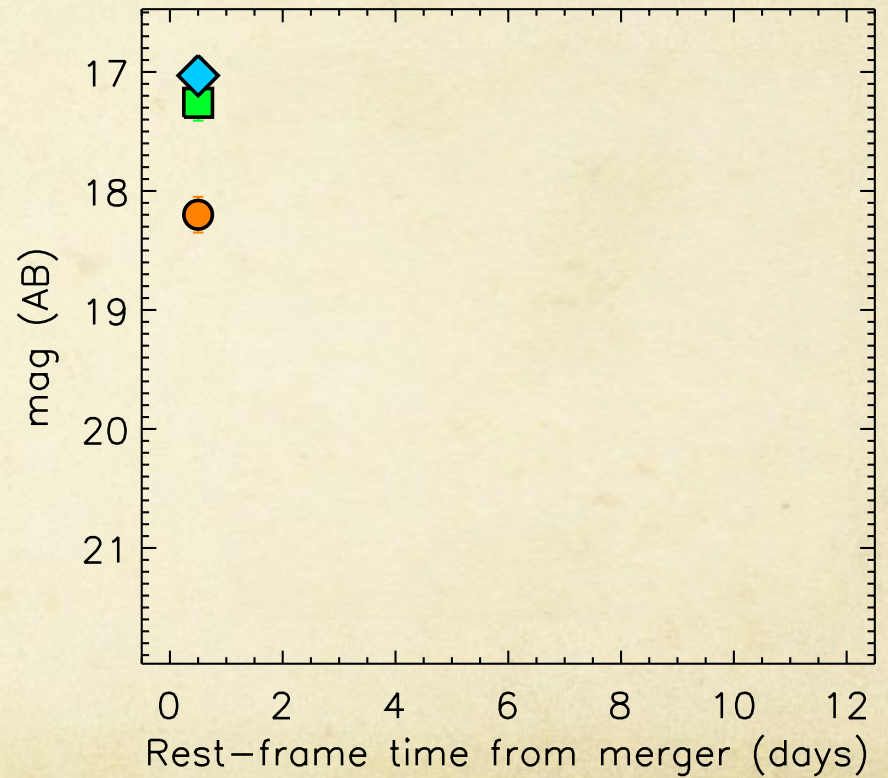
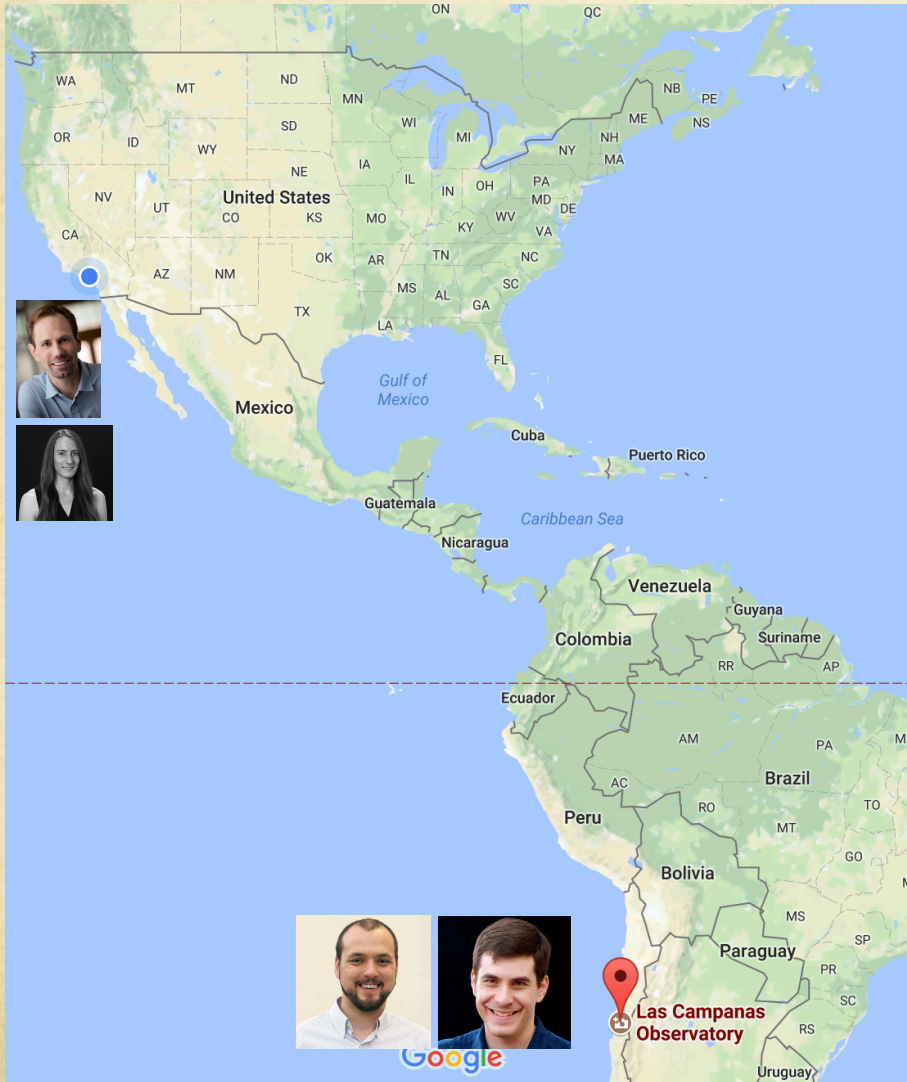
Magellan/Clay and
Magellan/Baade 6.5m:
Optical Photometry
NIR Photometry
Optical Spectroscopy



Where in the World are...
Carnegie Scientists

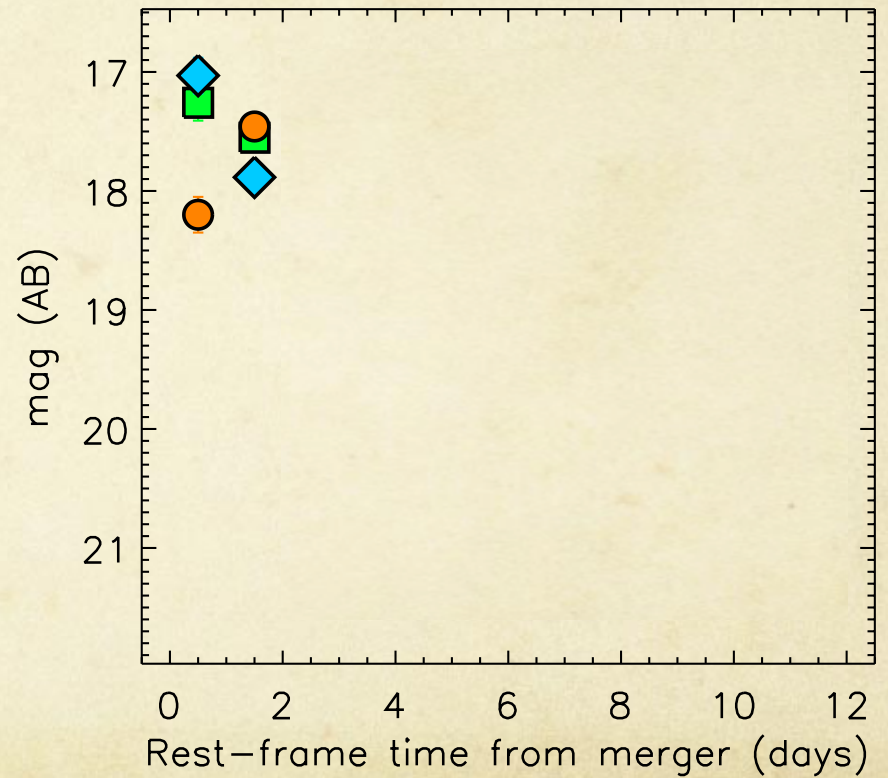
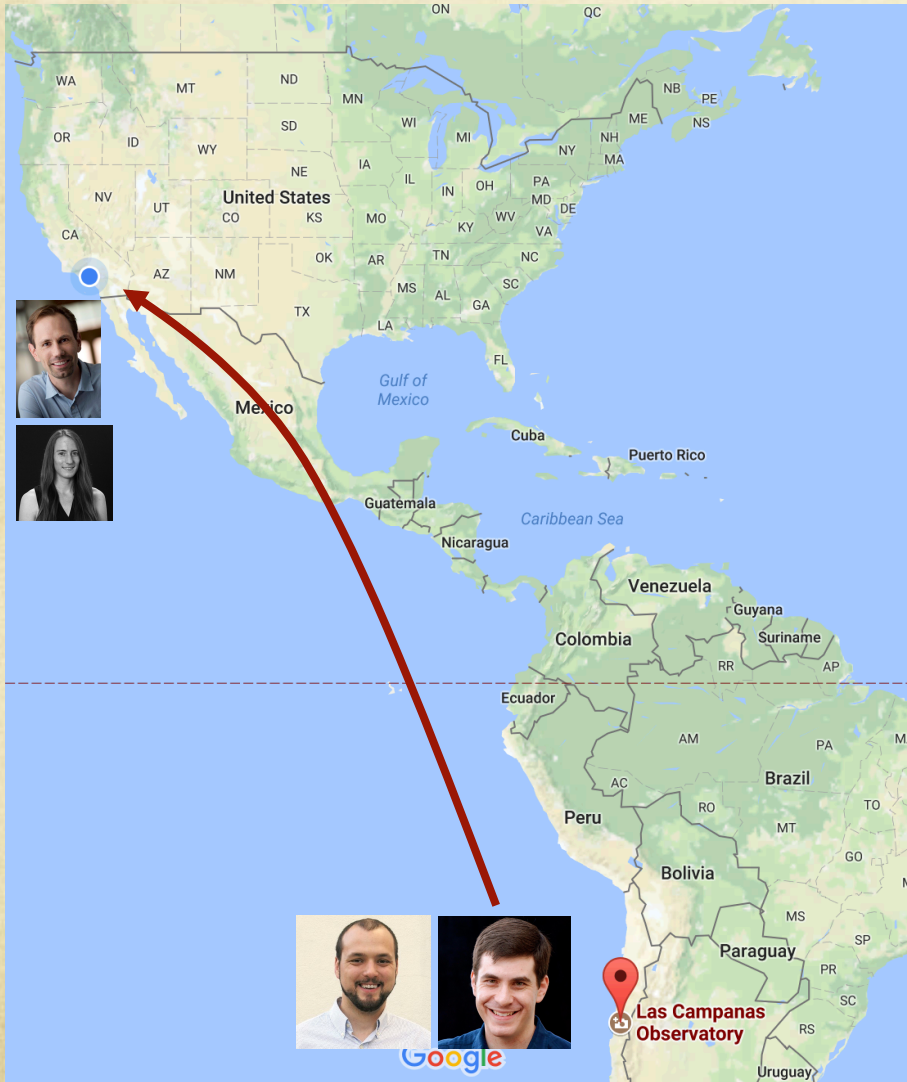
Where in the World are... Carnegie Scientists

+0.5 days (Thurs.)



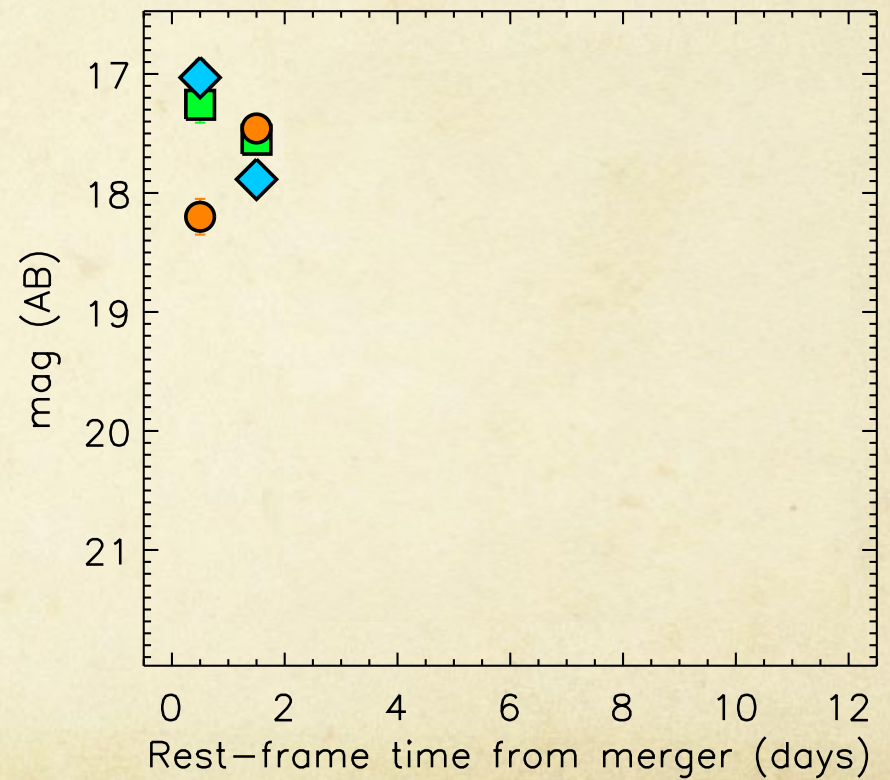
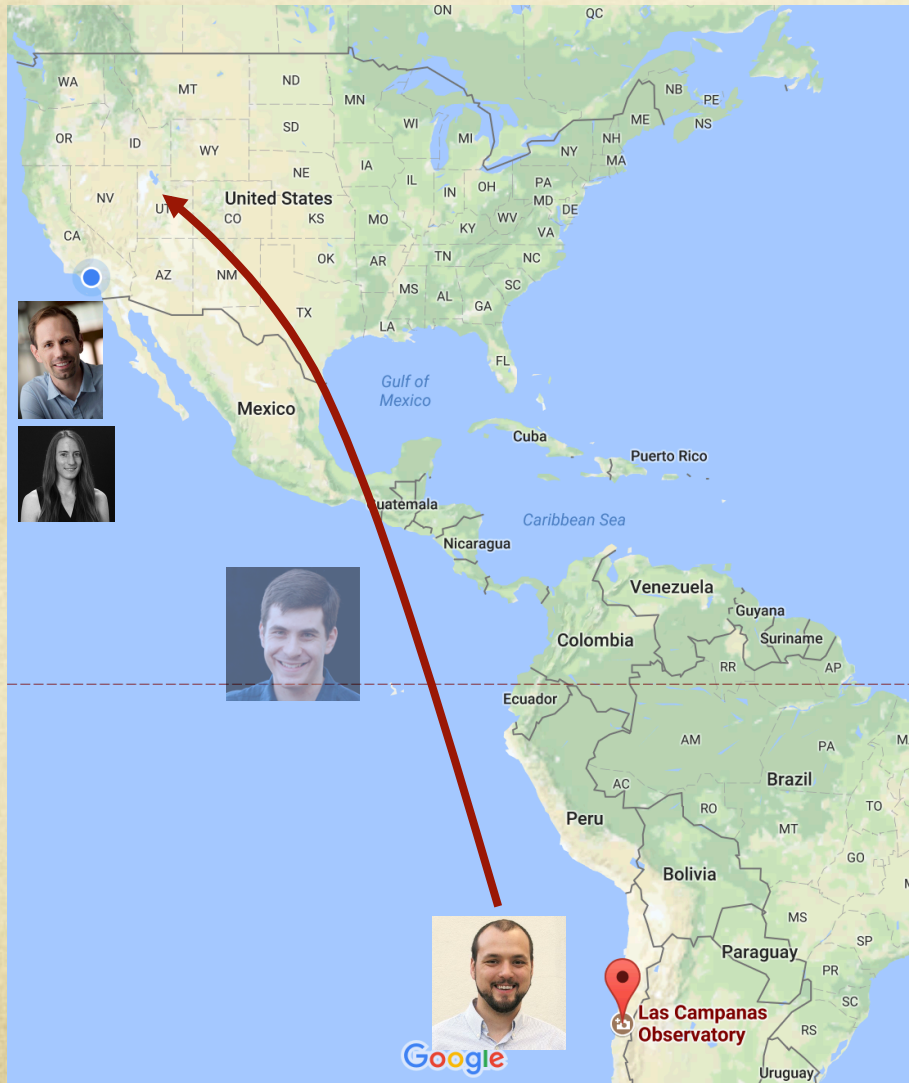
Where in the World are... Carnegie Scientists

+1.5 days (Fri.)



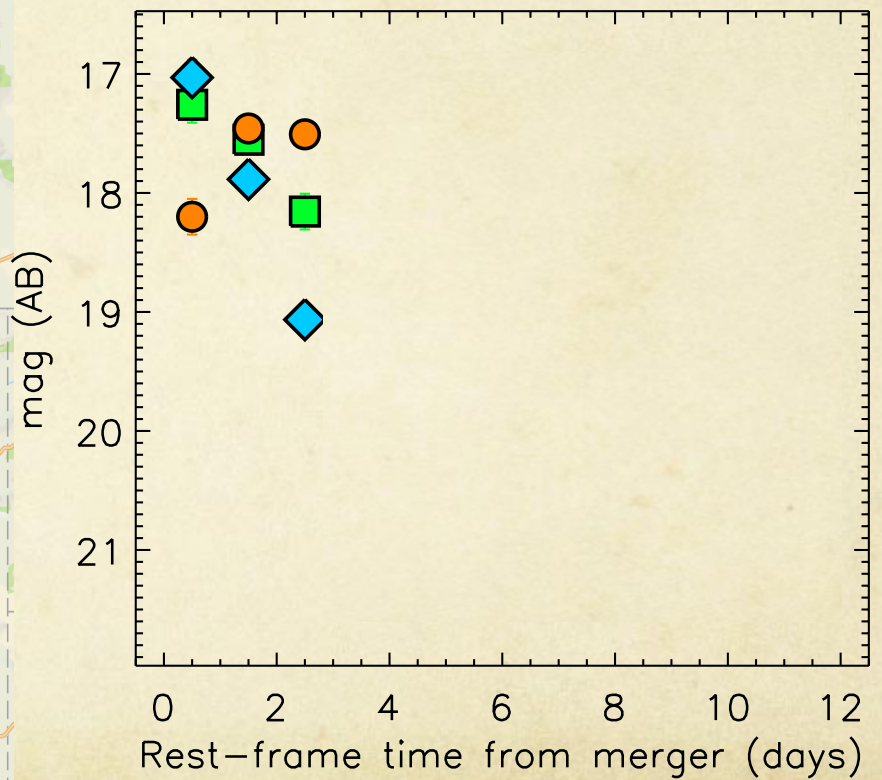
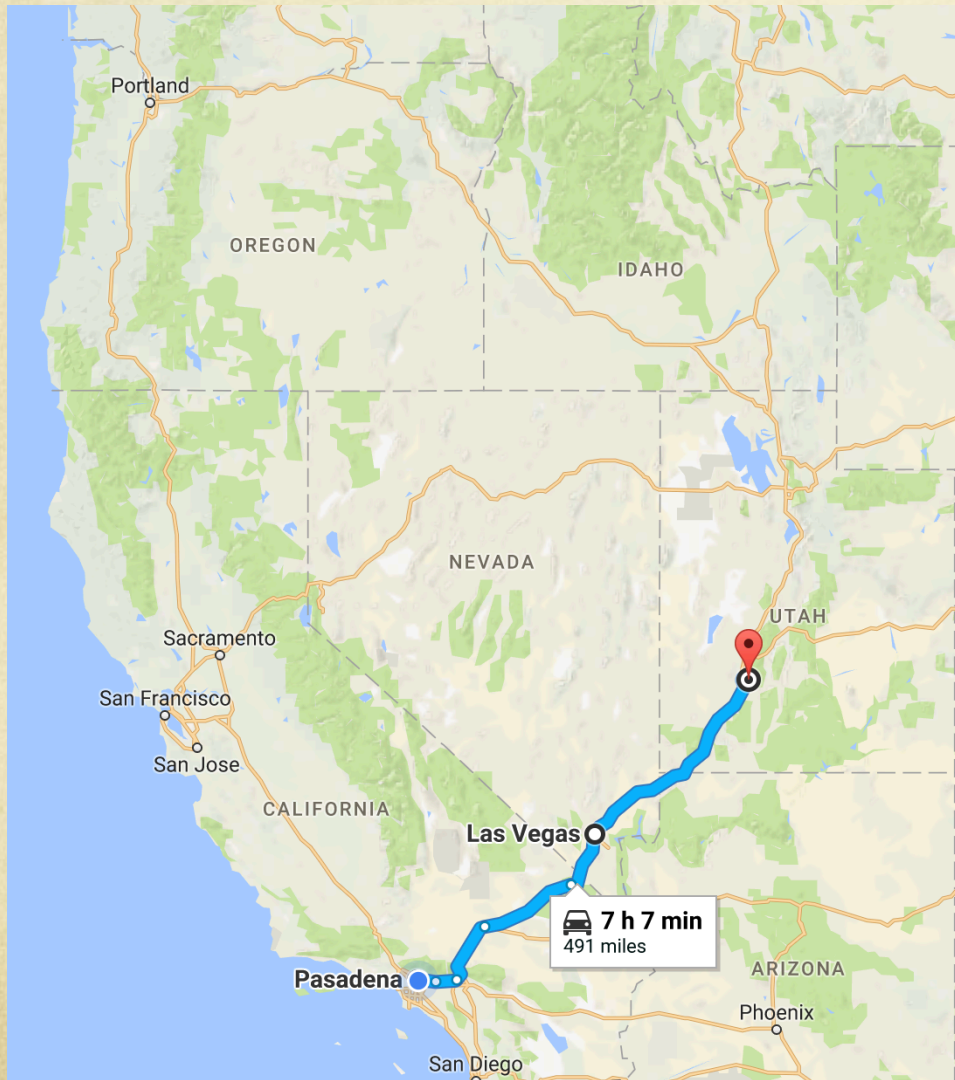
Where in the World are... Carnegie Scientists

+2.5 days (Sat.)



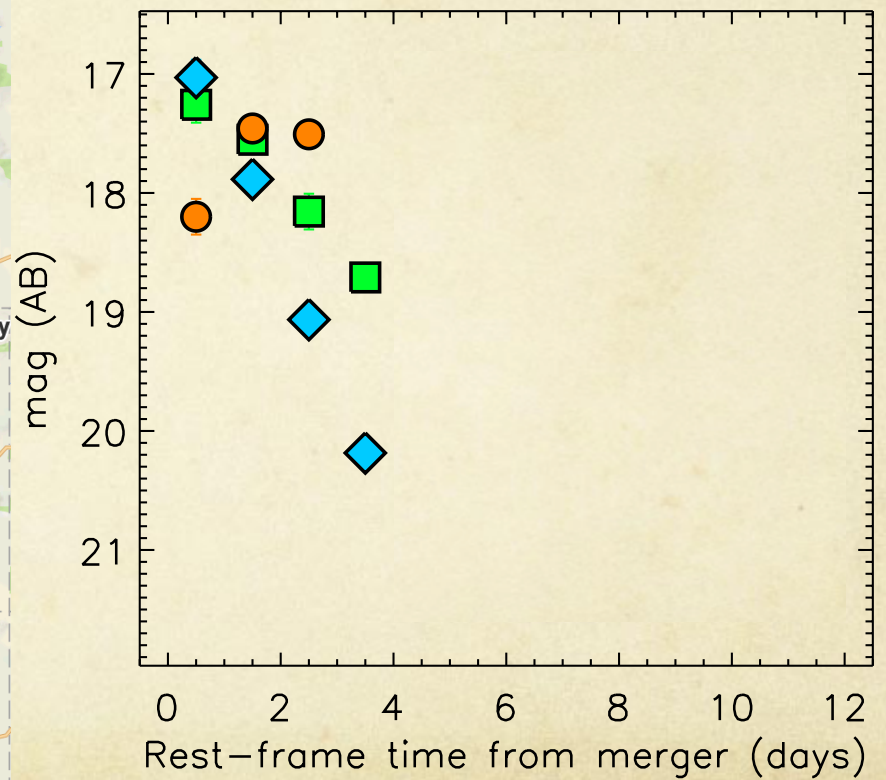
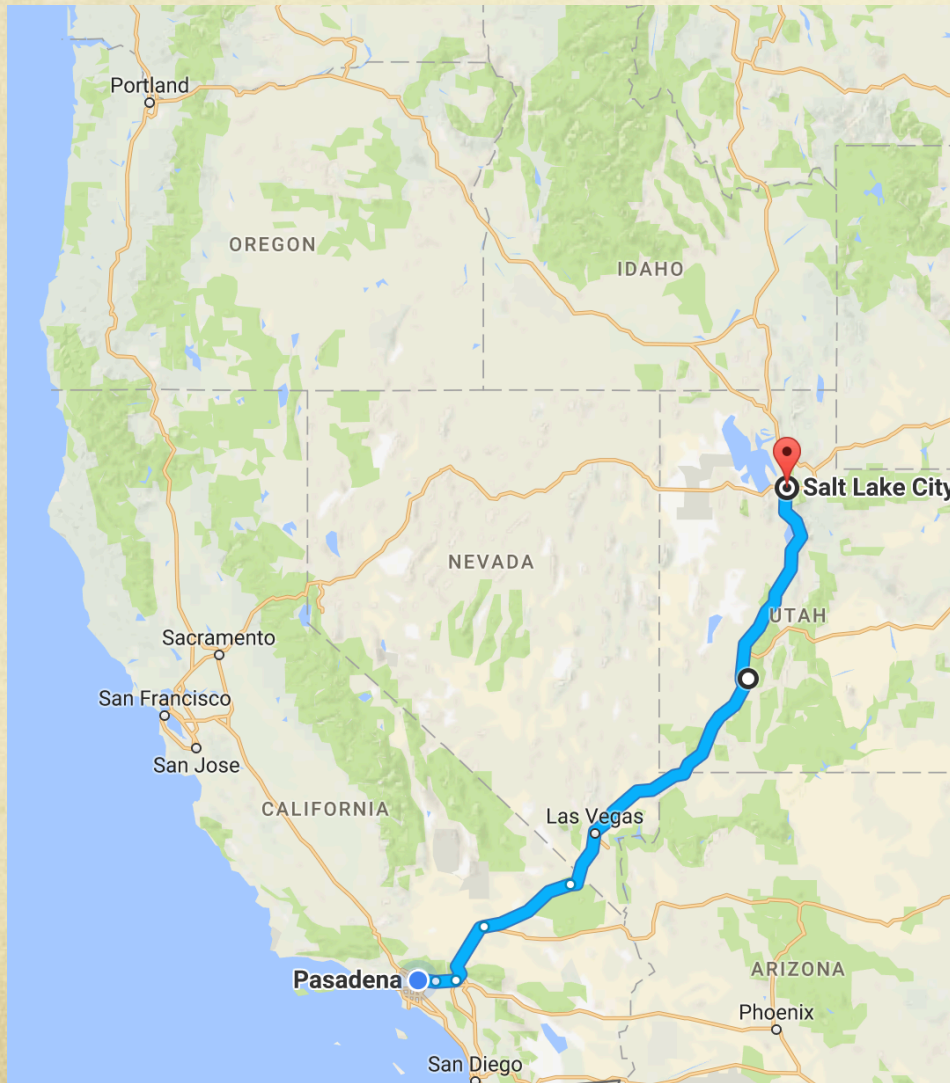
Where in the World are... Carnegie Scientists

+2.5 days (Sat.)



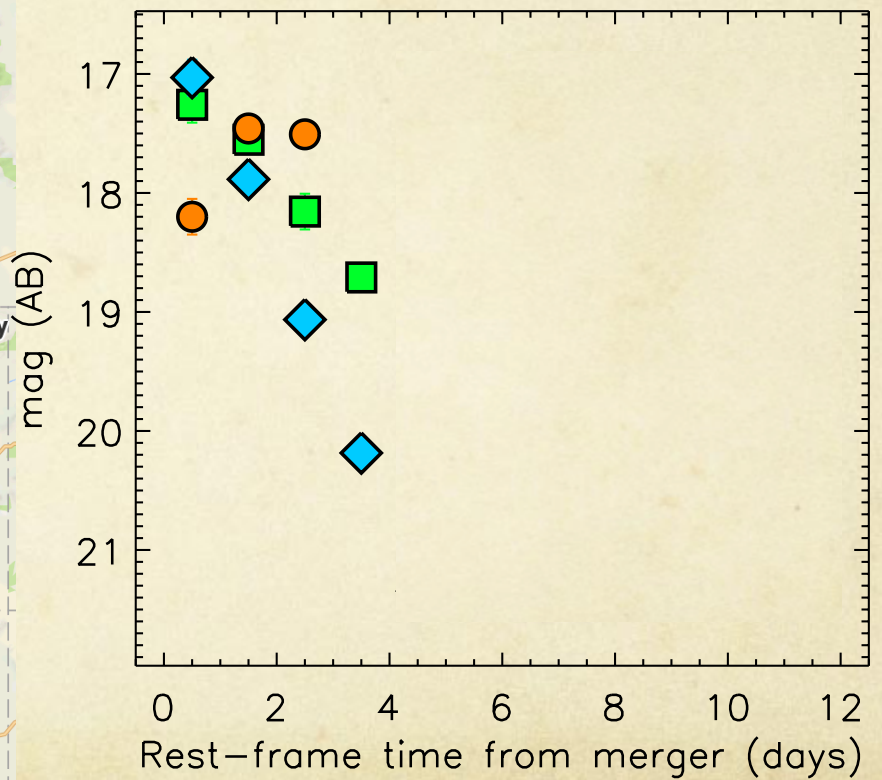
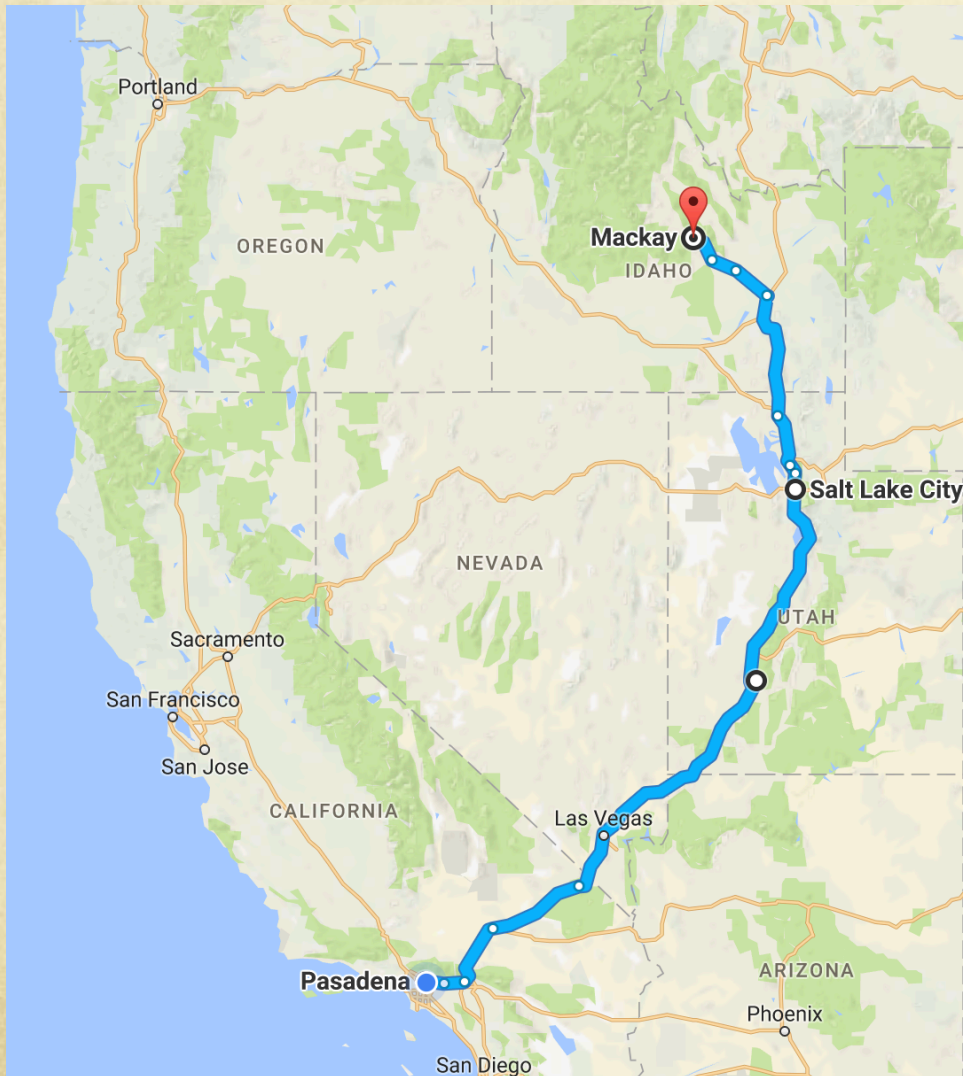
Where in the World are... Carnegie Scientists

+3.5 days (Sun.)



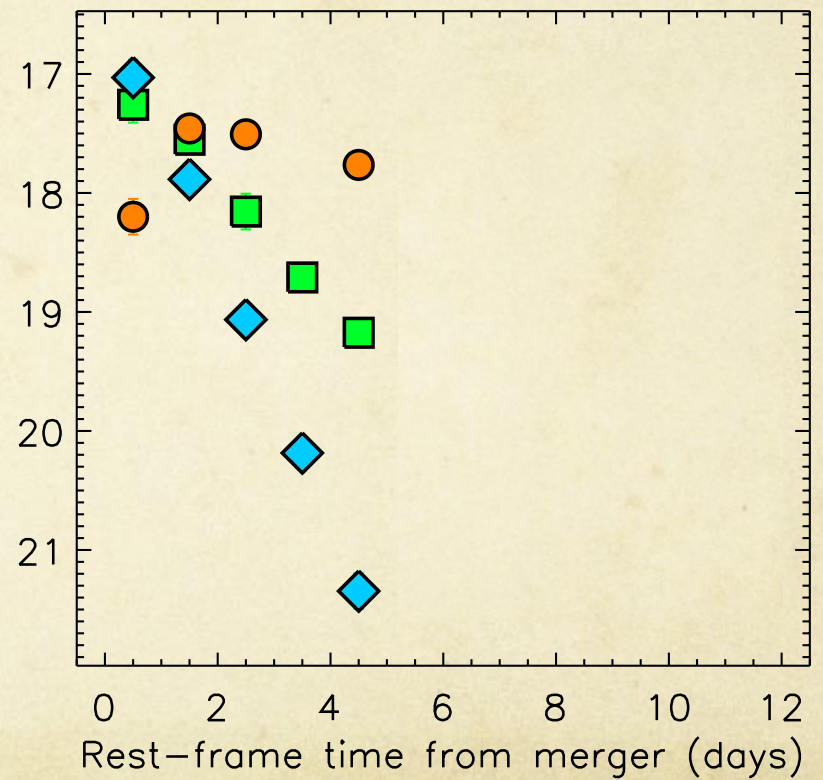
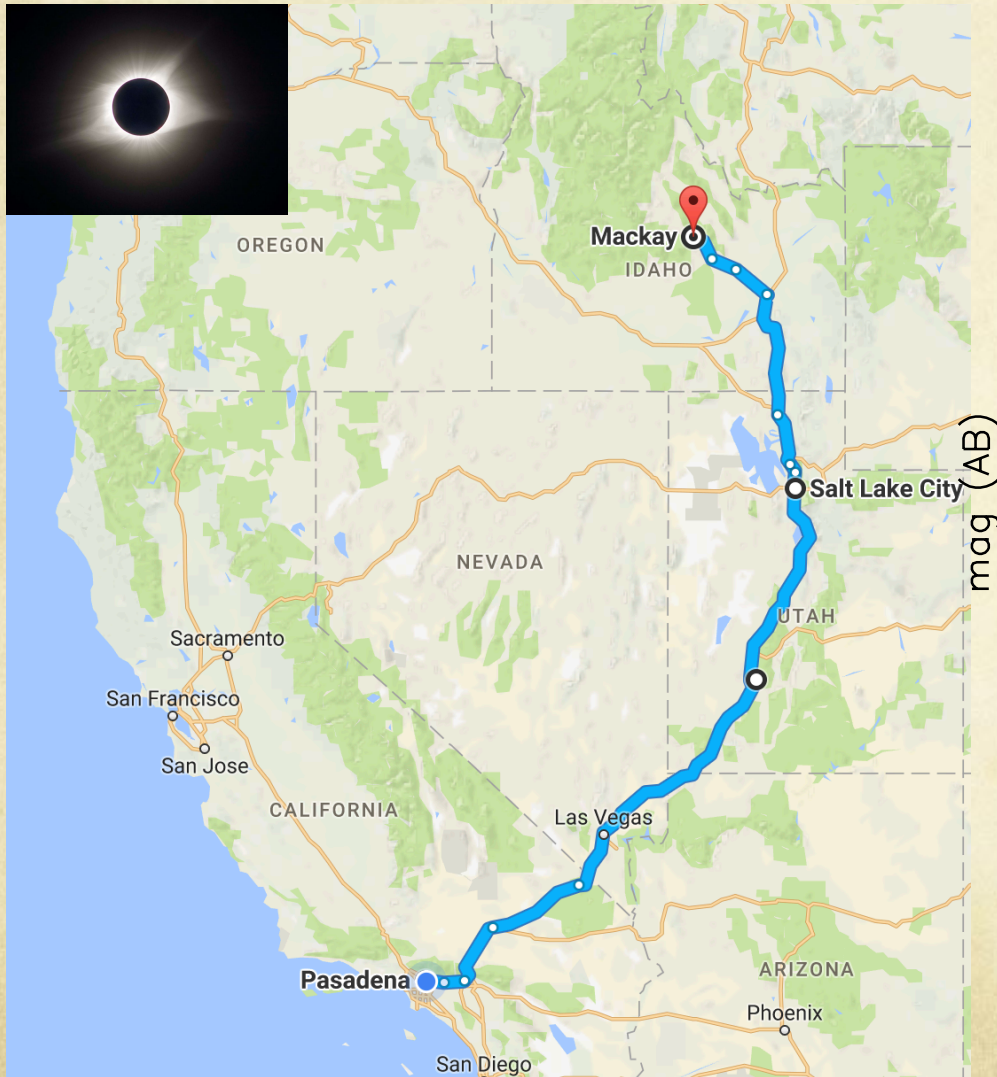
Where in the World are... Carnegie Scientists

+3.5 days (Sun.)



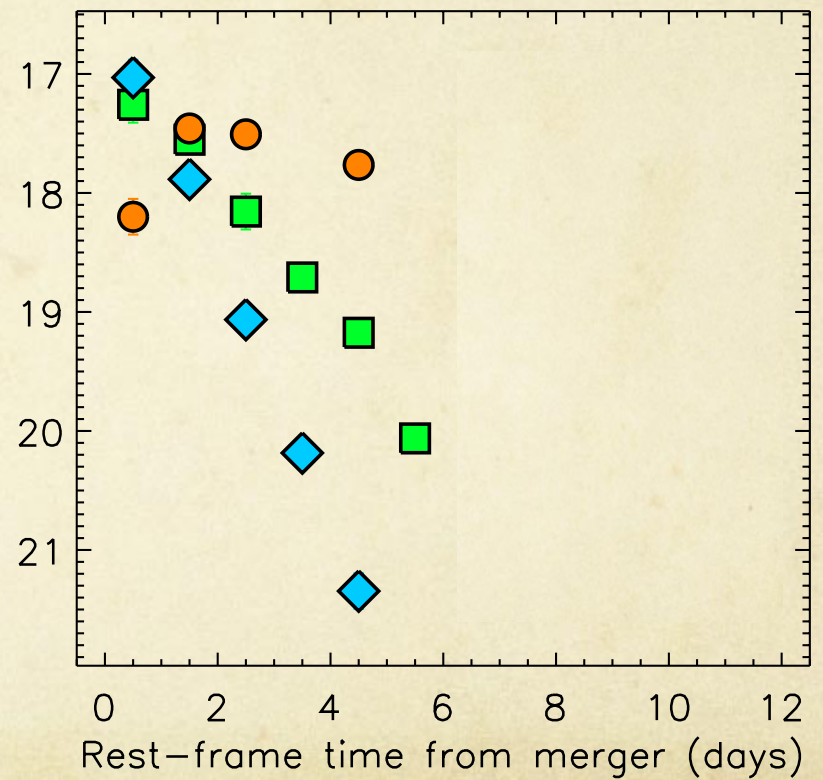
Where in the World are... Carnegie Scientists

+4.5 days (Mon.)



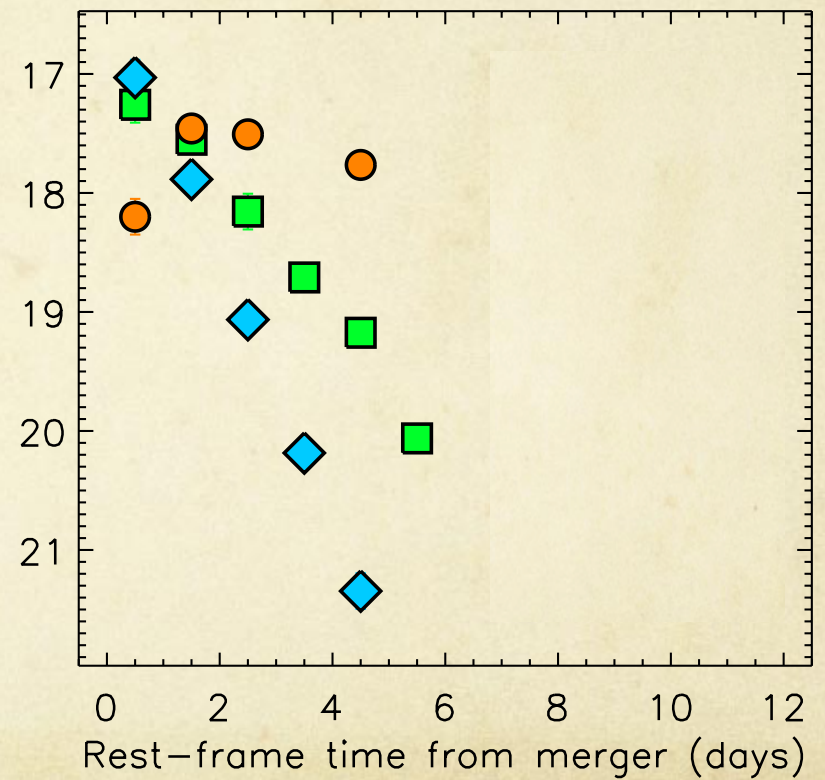
Where in the World are... Carnegie Scientists

+5.5 days (Tues.)



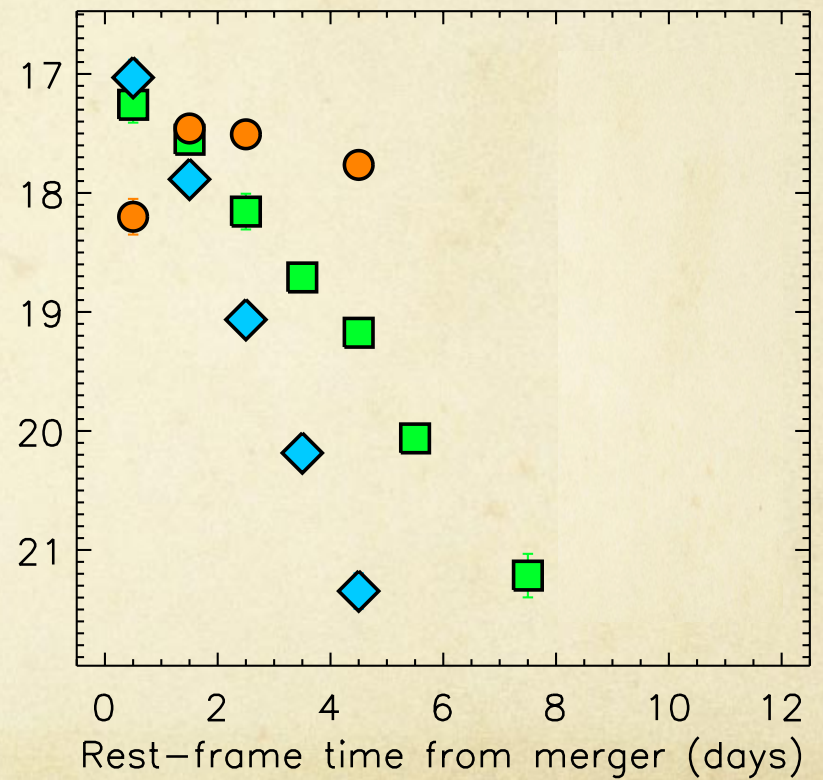
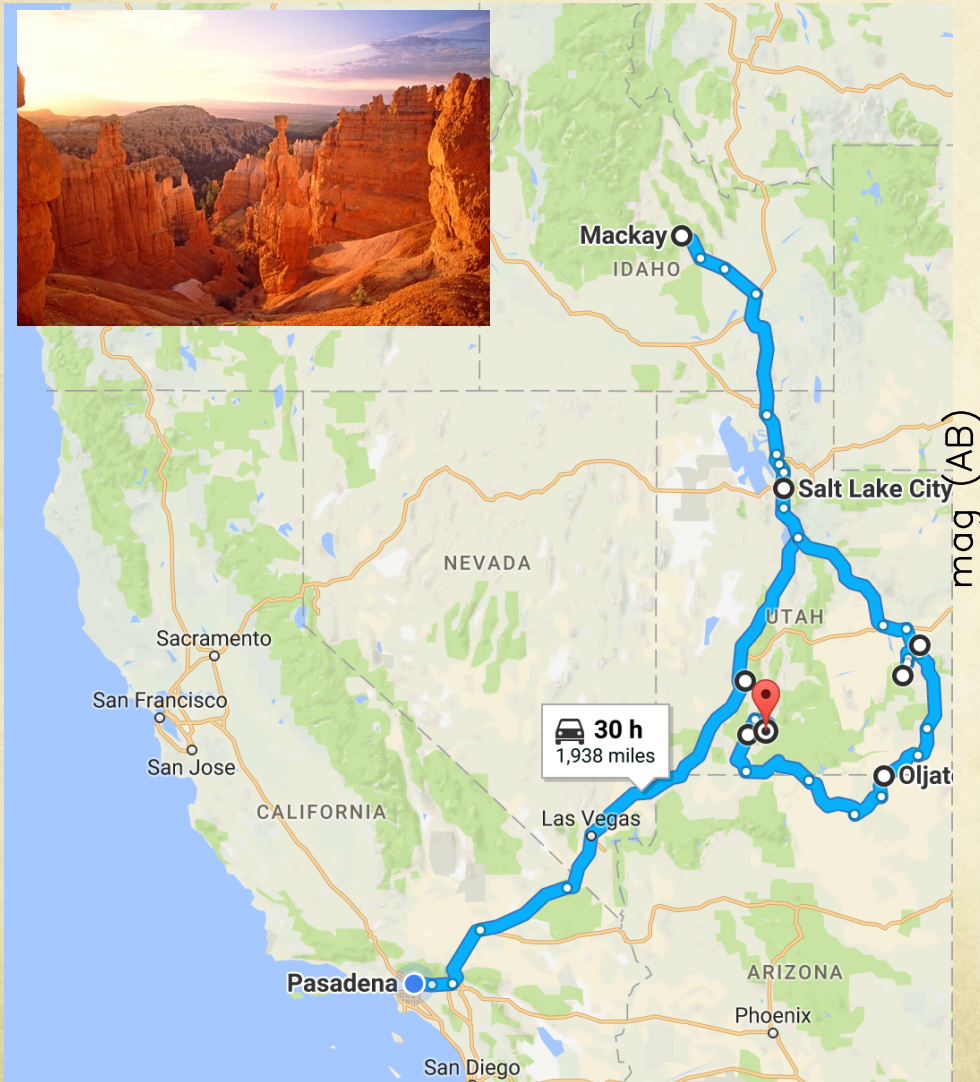
Where in the World are... Carnegie Scientists

+6.5 days (Wed.)



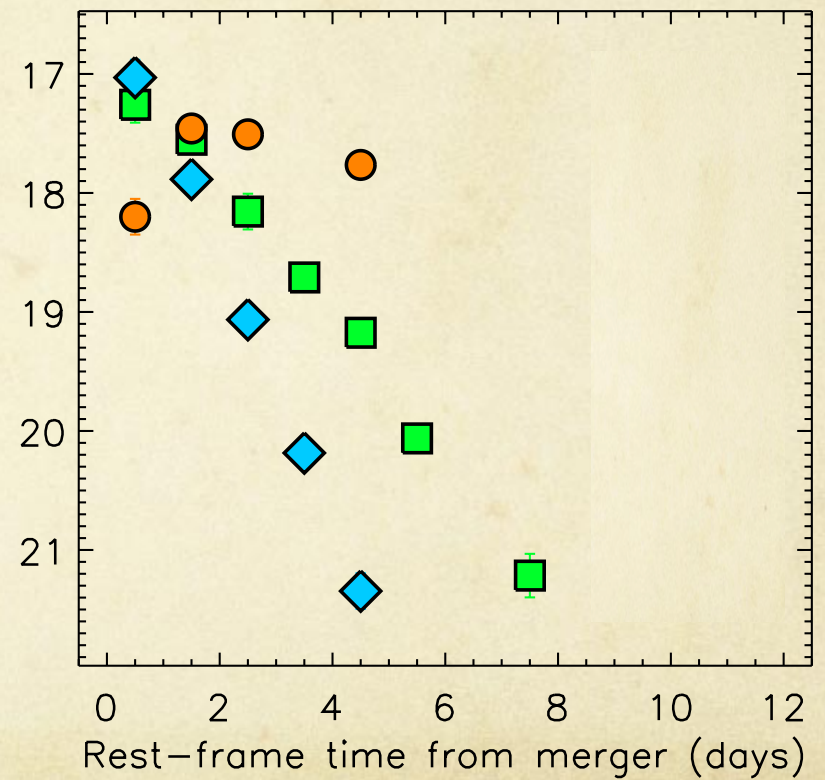
Where in the World are... Carnegie Scientists

+7.5 days (Thurs.)



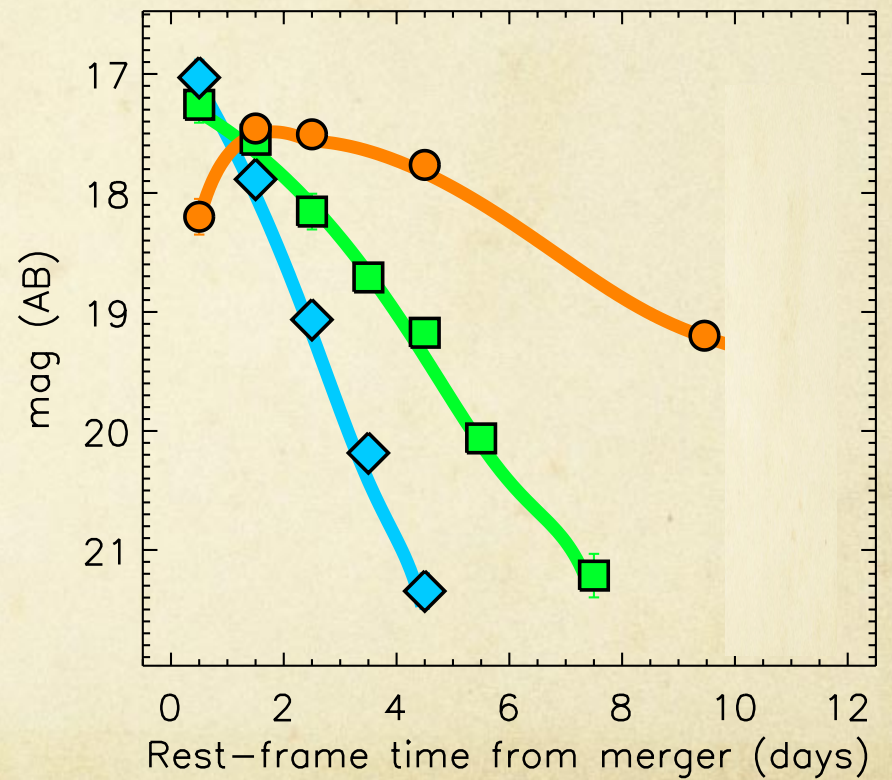
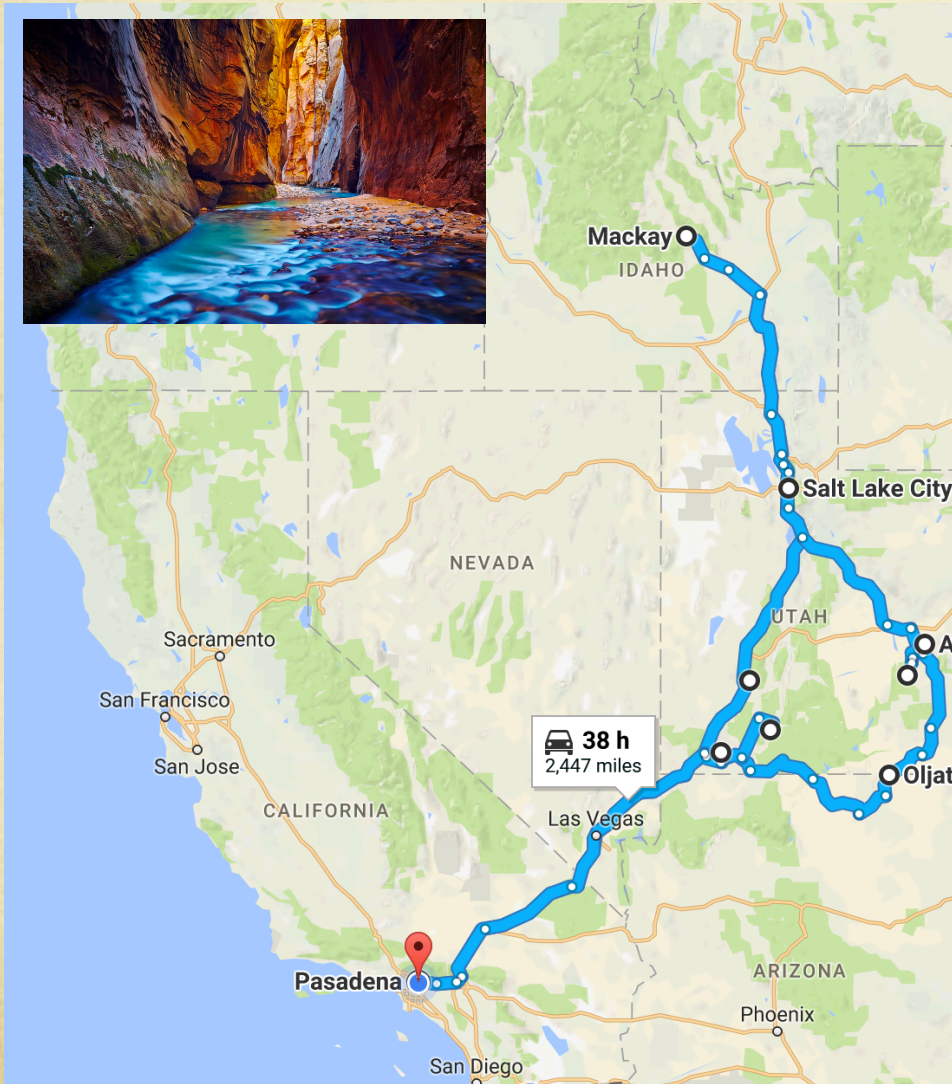
Where in the World are... Carnegie Scientists

+8.5 days (Fri.)

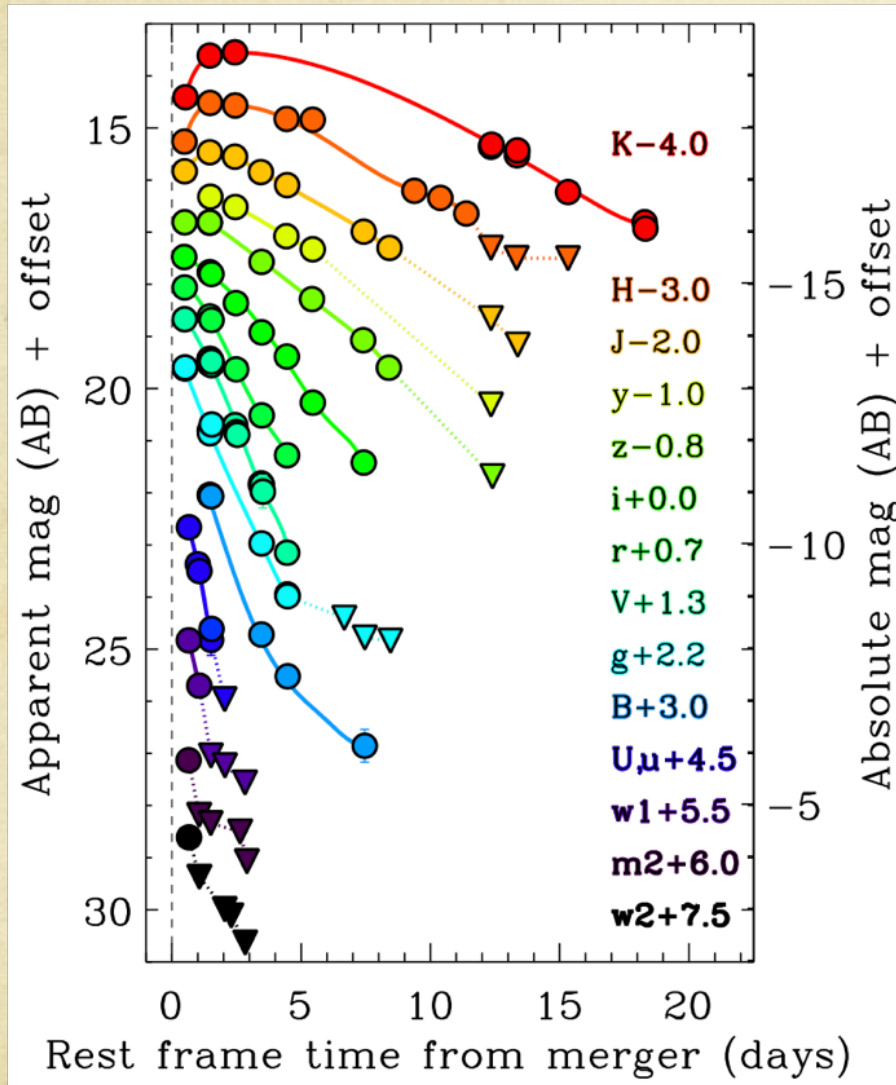


Where in the World are... Carnegie Scientists

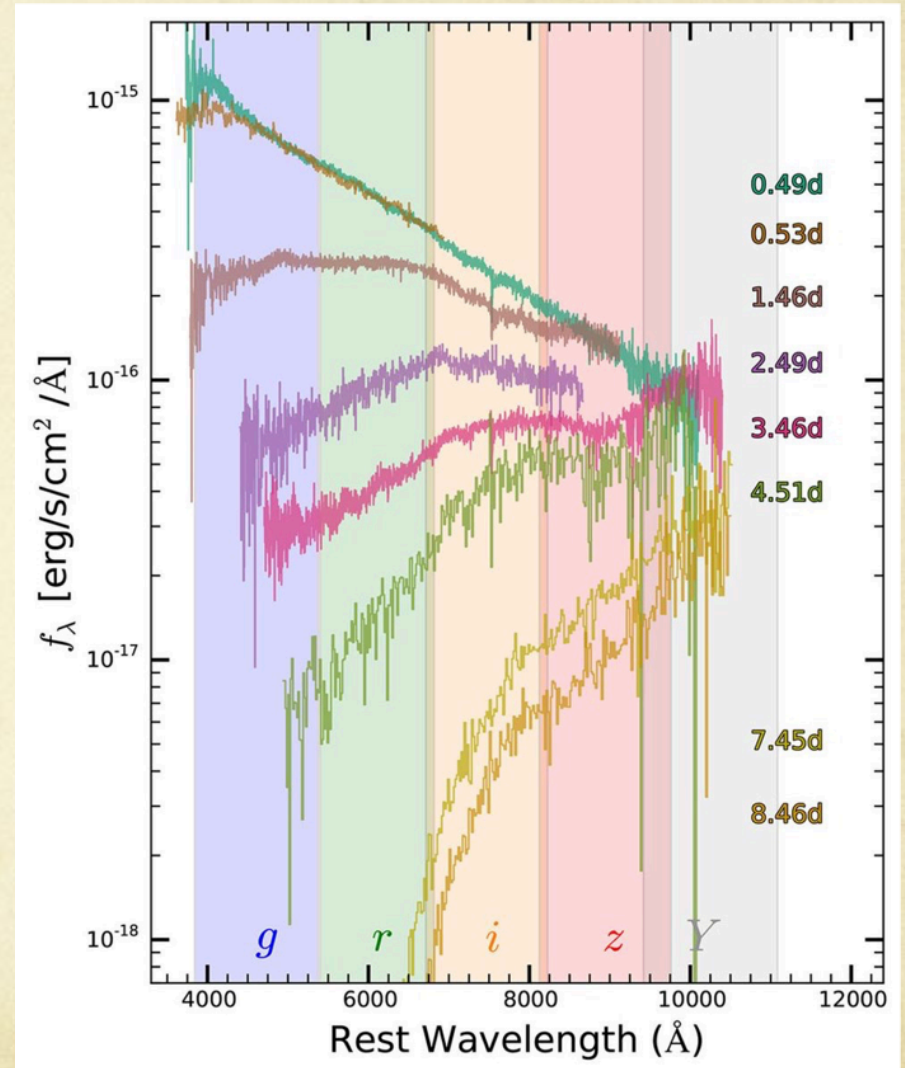
+9.5 days (Sat.)



SSS17a: Observations

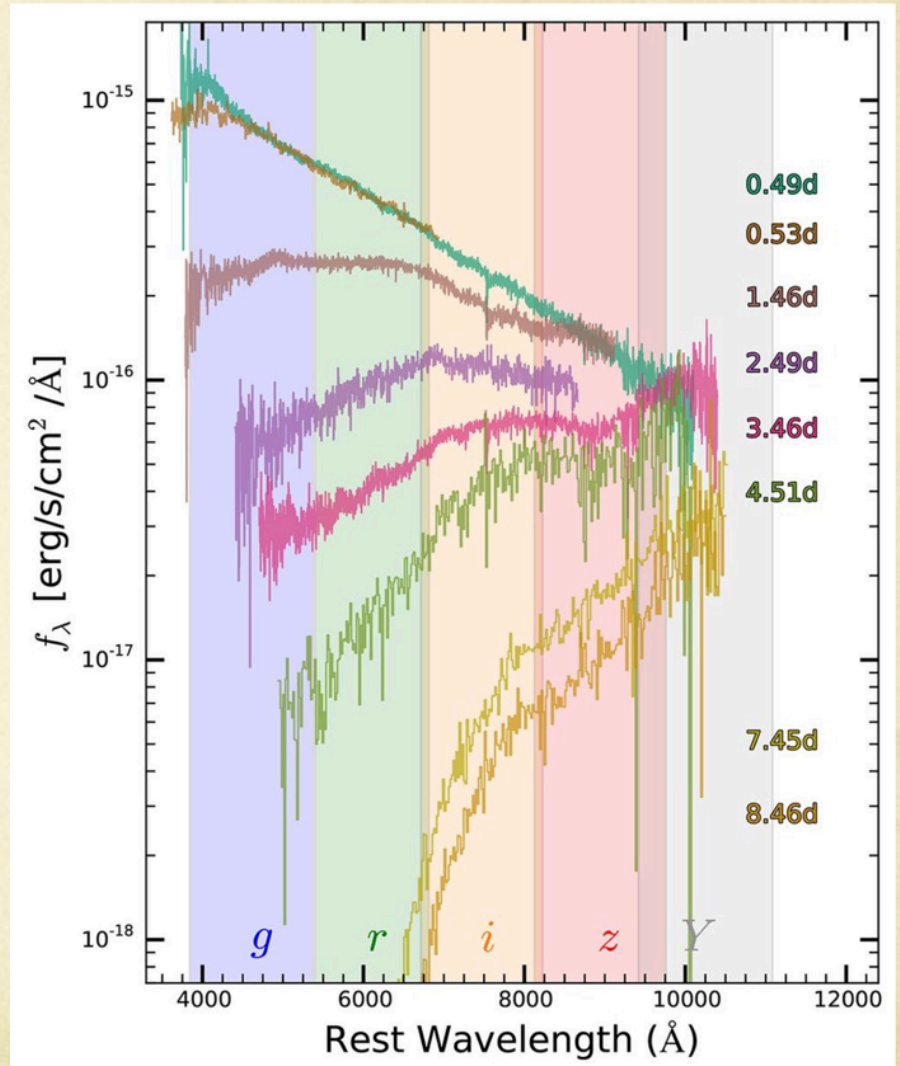
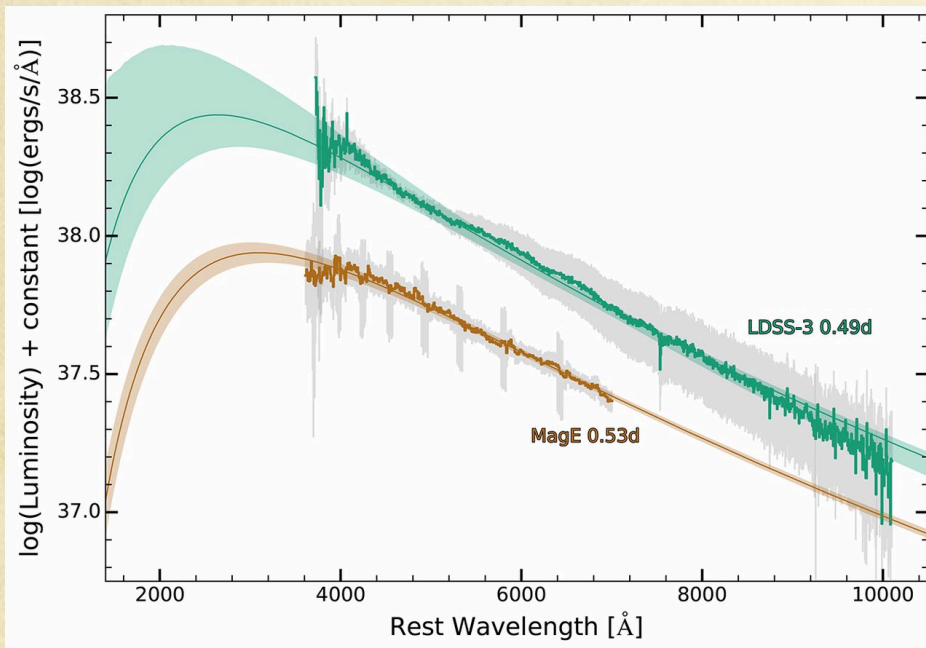


Drout et al. (2017)



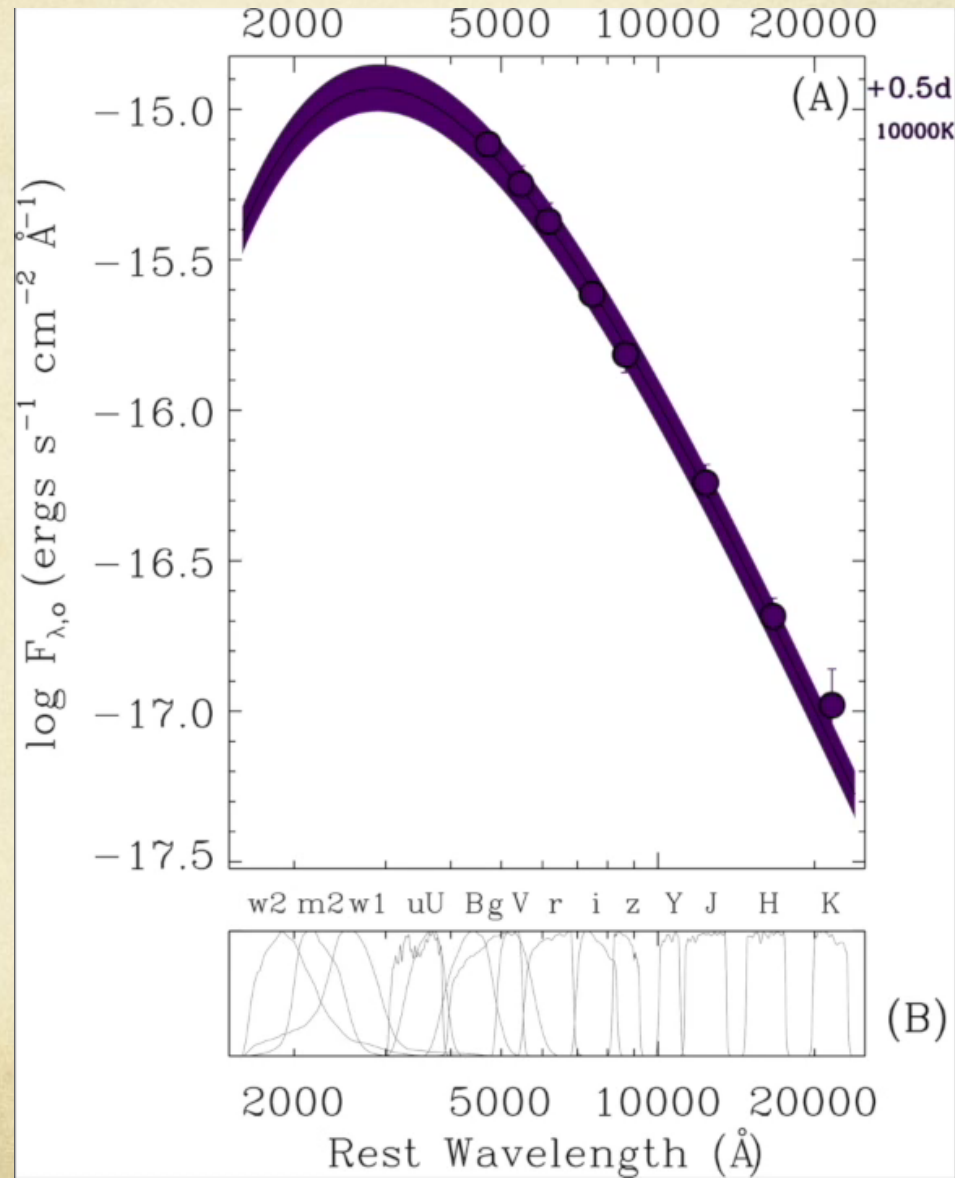
Shappee et al. (2017)

SSS17a Evolution



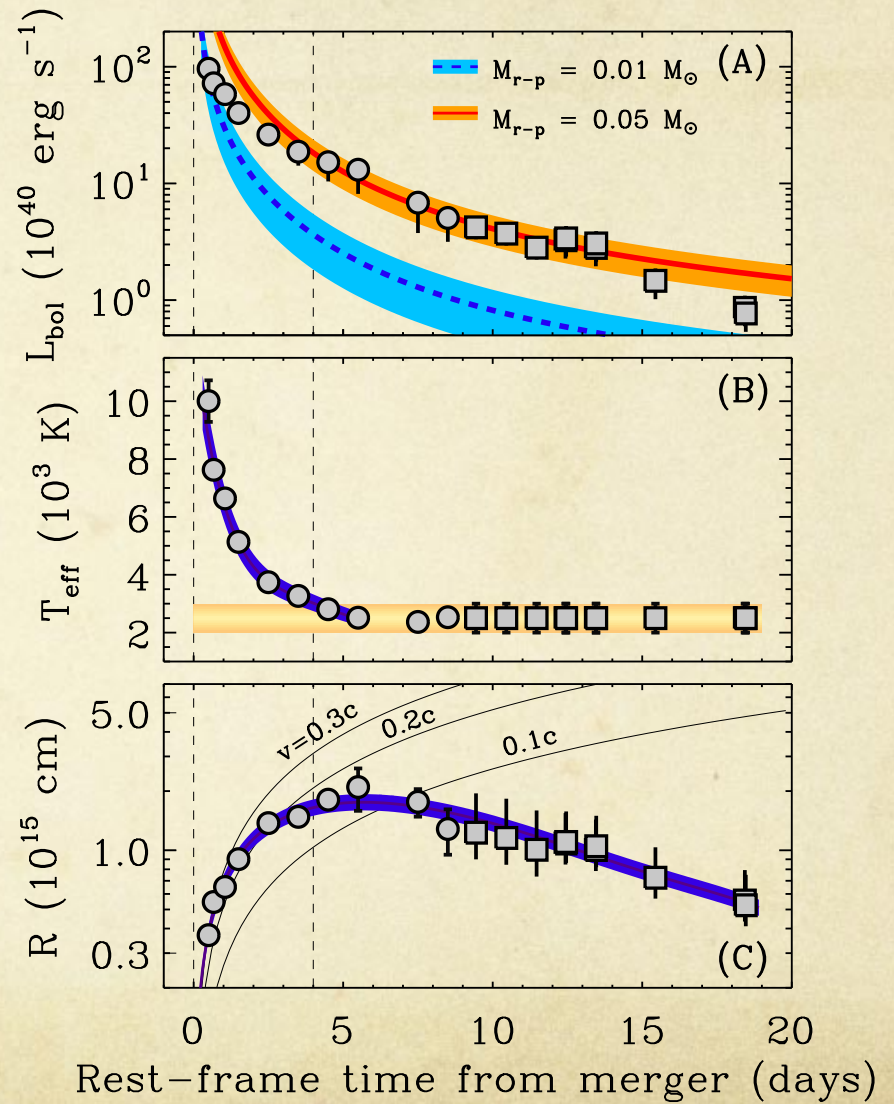
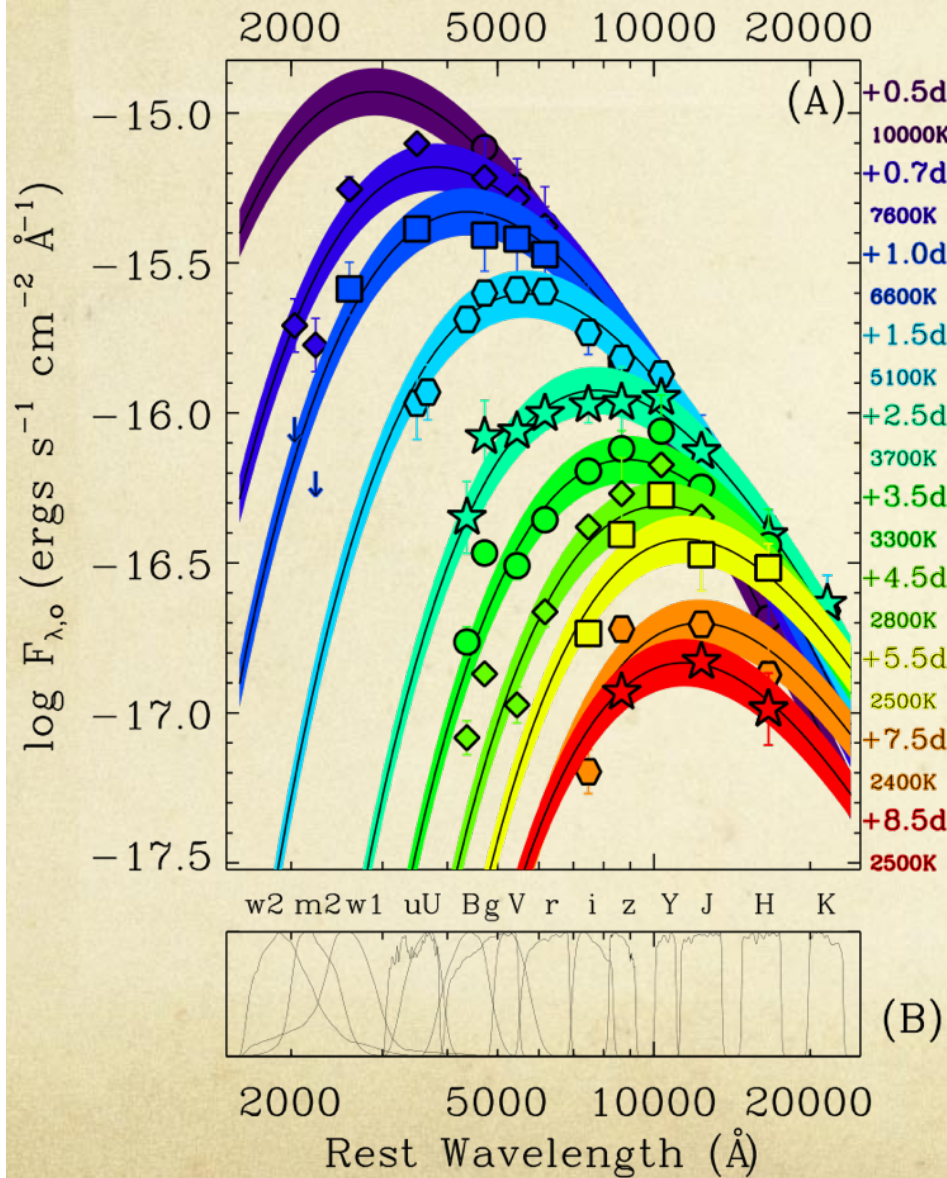
Shappee et al. (2017)

SSS17a: Observations



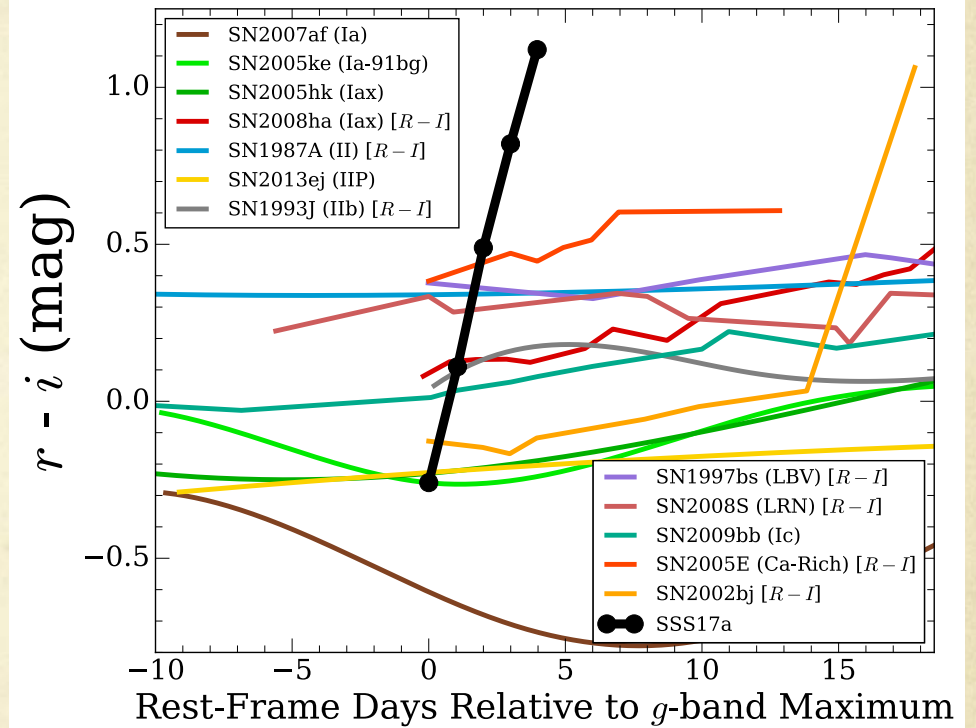
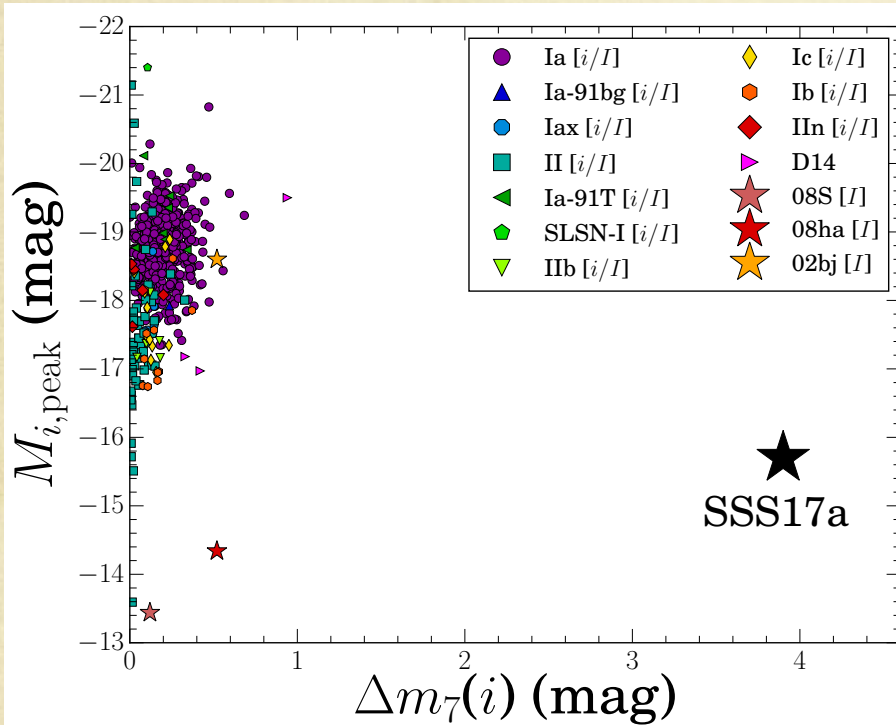
Drout et al. (2017)

SSS17a: Observations

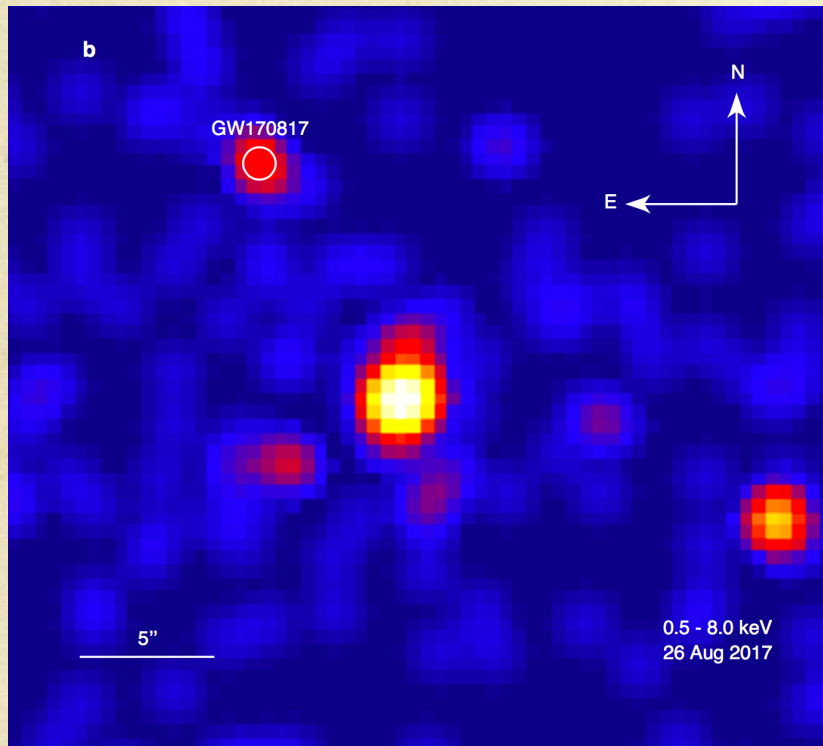


Drout et al. (2017)

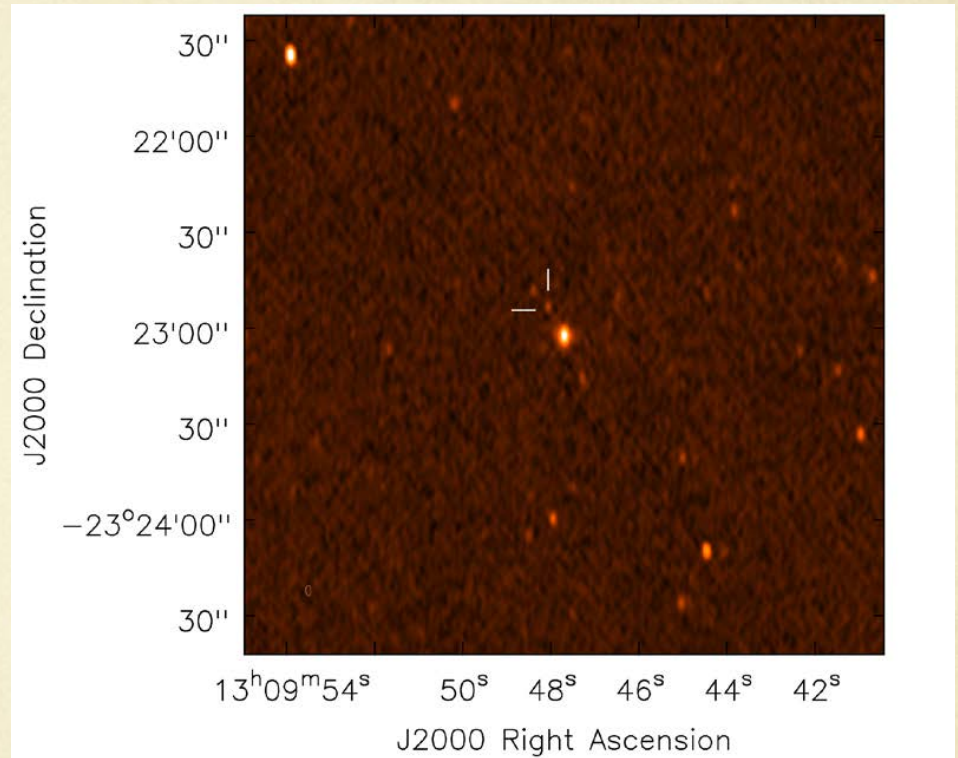
SSS17a is Unique



X-ray and Radio Emission



Chandra +9d; (Troja et al. 2017)



JVLA +16.4d; (Hallinan et al. 2017)

All Together:

Earth

Space



LIGO

What does it all mean?

Multi-messenger Astronomy

Motivation

Gravitational Wave data only provides specific information

1. Precise Location/Distance Determination
 - Provides environmental information
 - Breaks degeneracy with inclination
2. Measure Influence/Test Merger Models
 - Origin of short-duration GRBs
 - Origin of r-process elements
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 - “Standard Siren” Hubble diagram

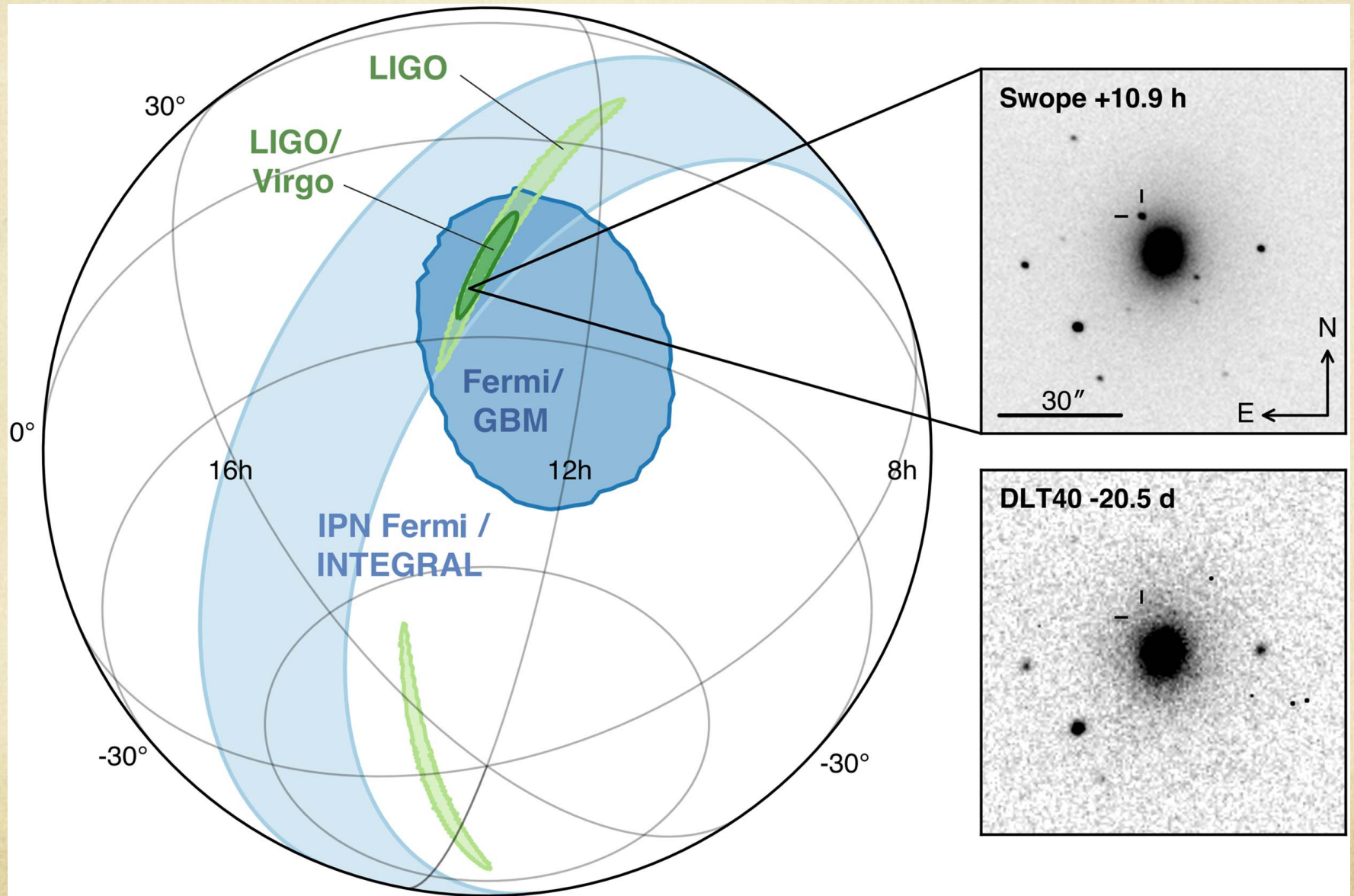
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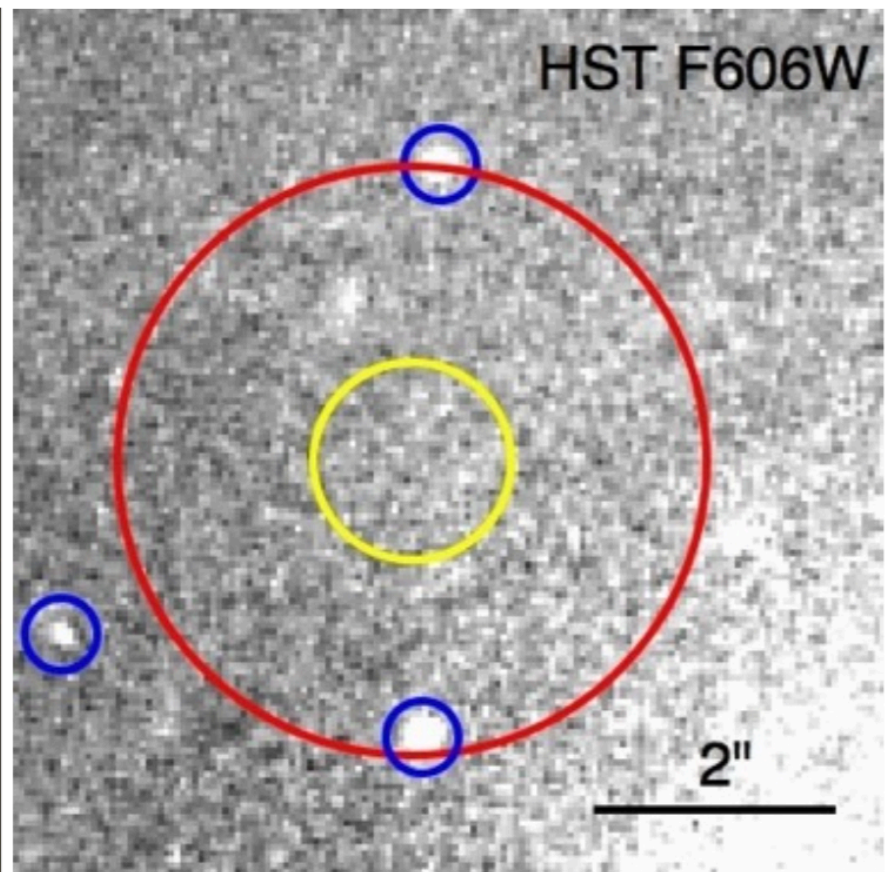
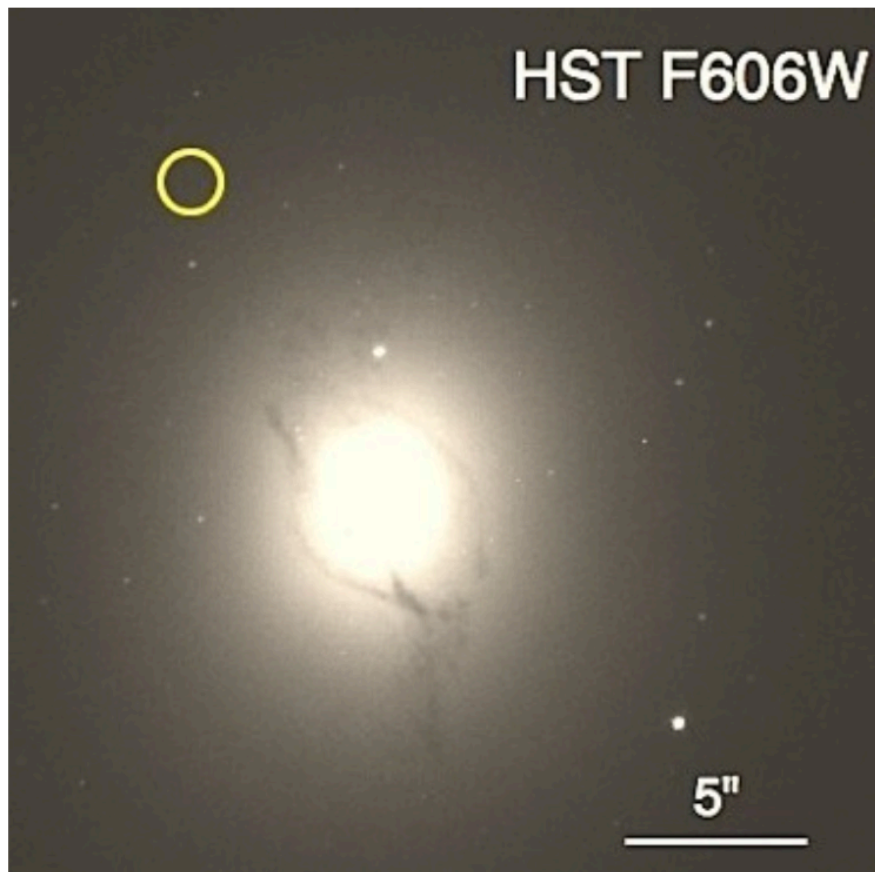
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Precise Localization



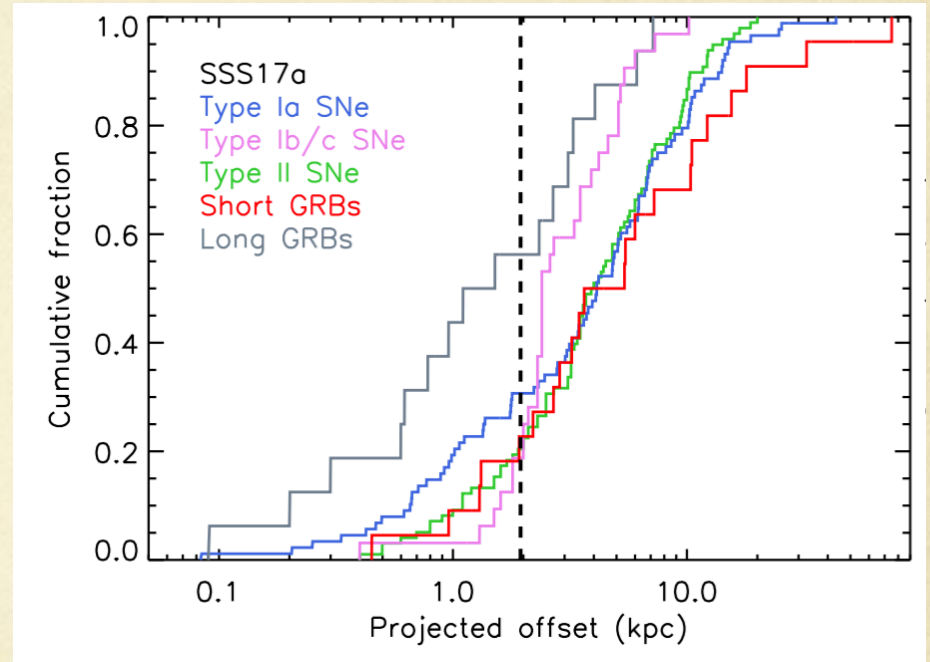
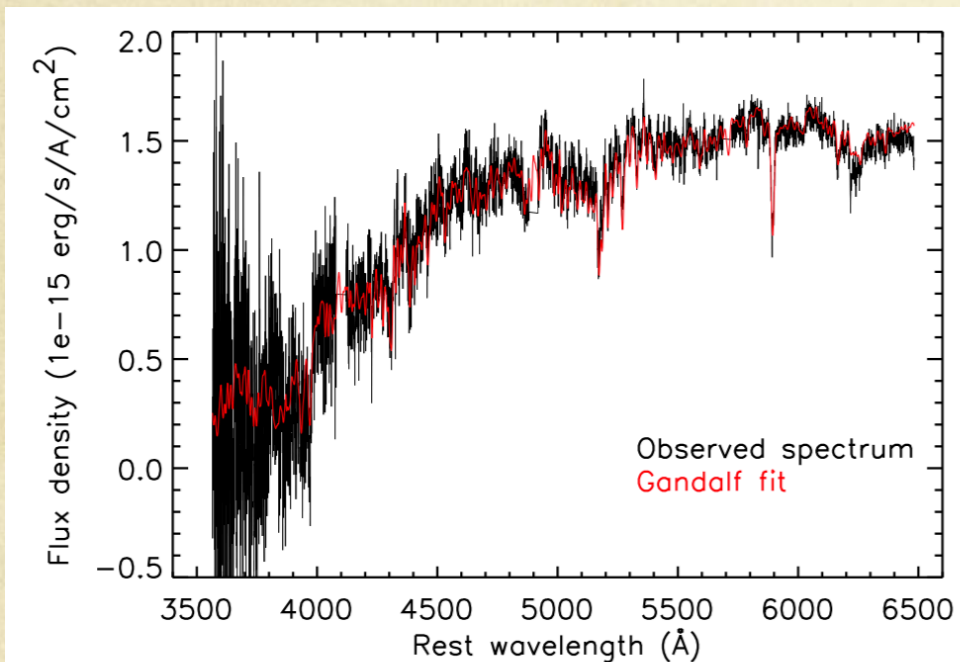
Host Environment

No Progenitor



Host Environment

Old Population, Similar to SGRB Hosts



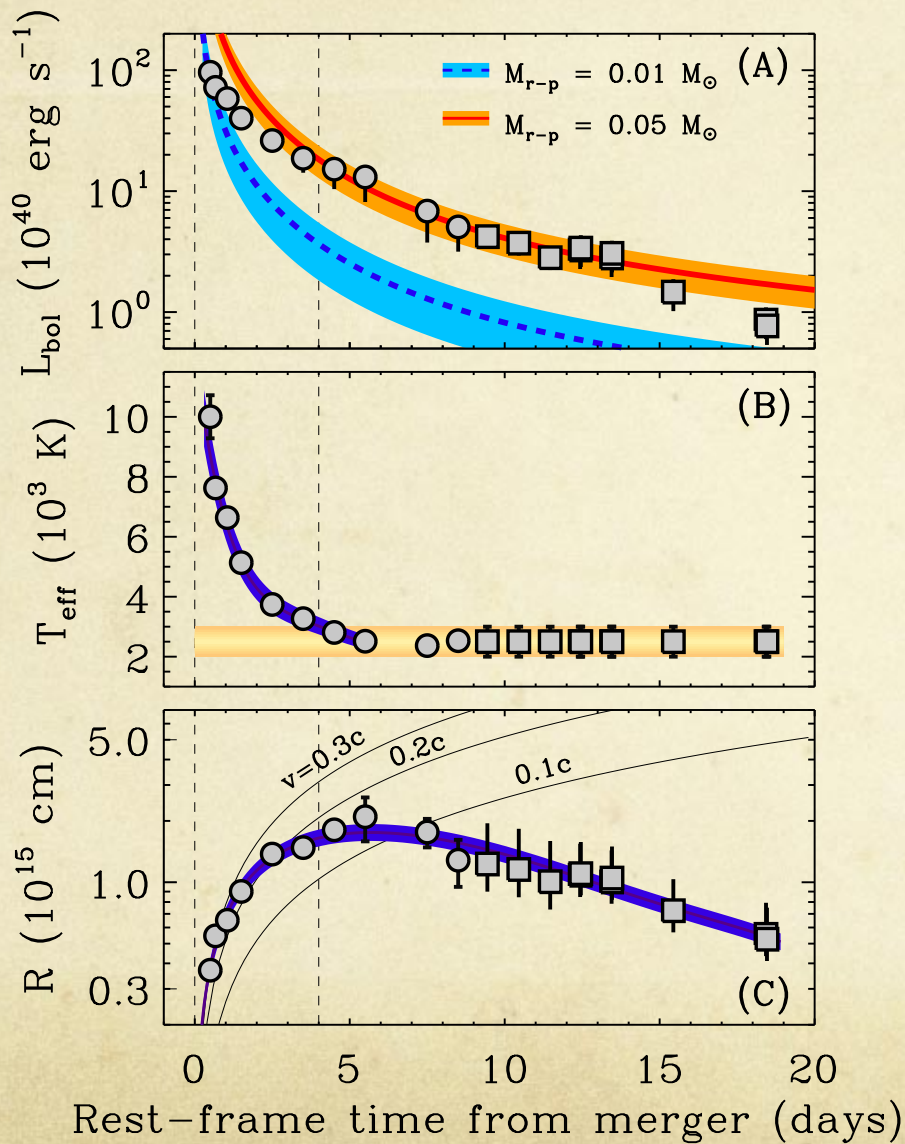
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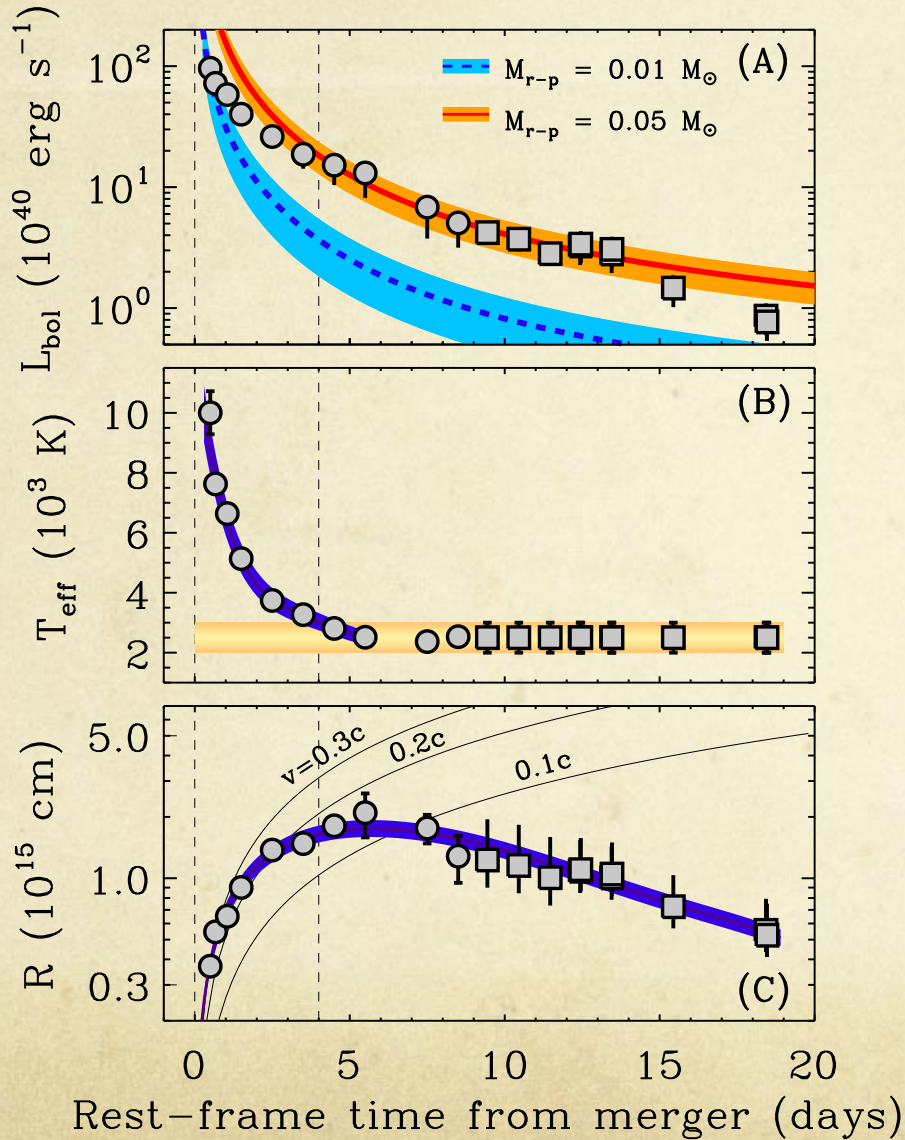
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SSS17a: Physical Parameters

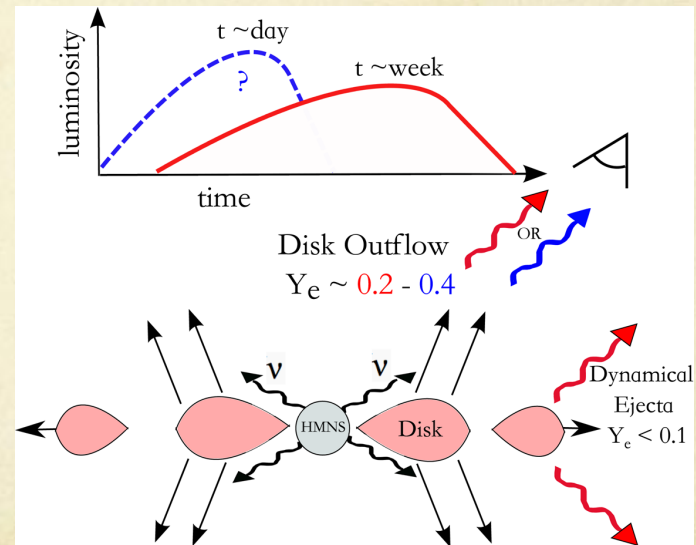


- r-process heating goes as $\tau^{1.3}$

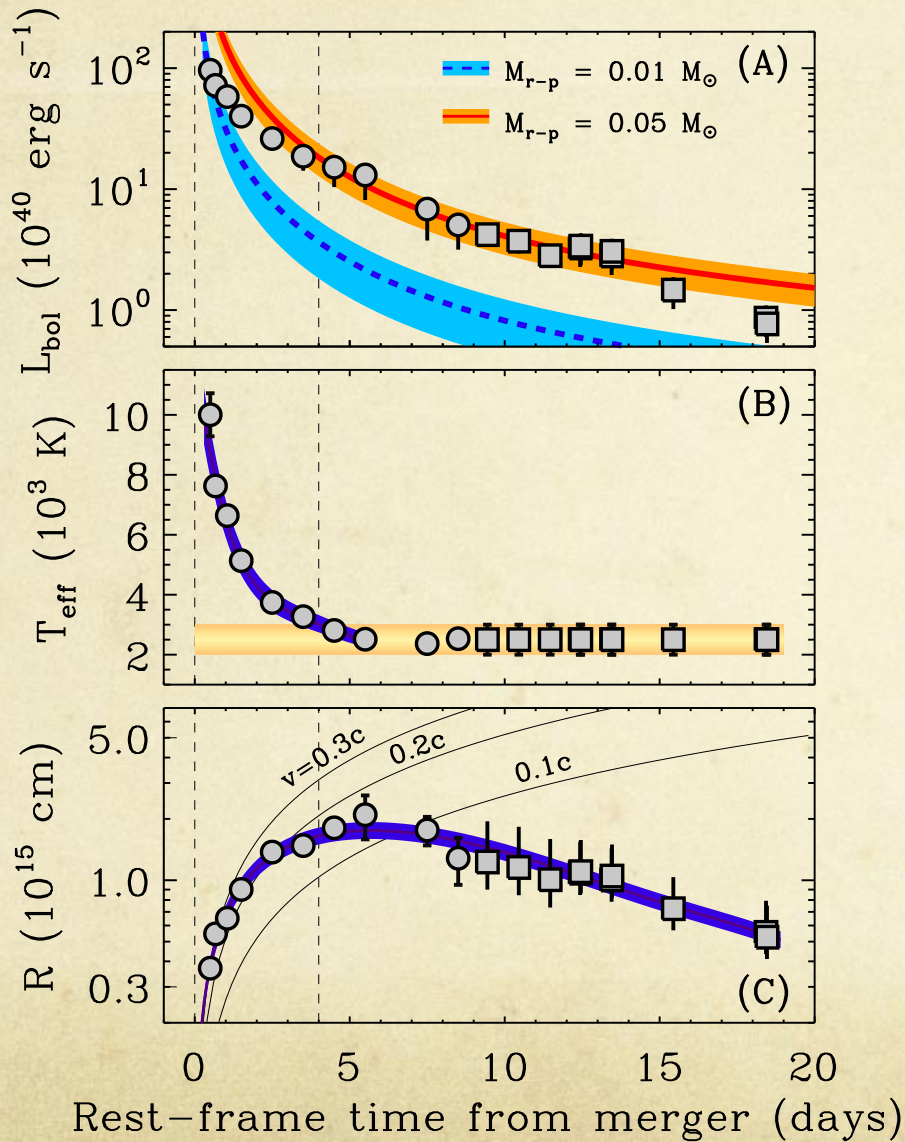
SSS17a: Physical Parameters



- r-process heating goes as $\tau^{1.3}$
- two components are required

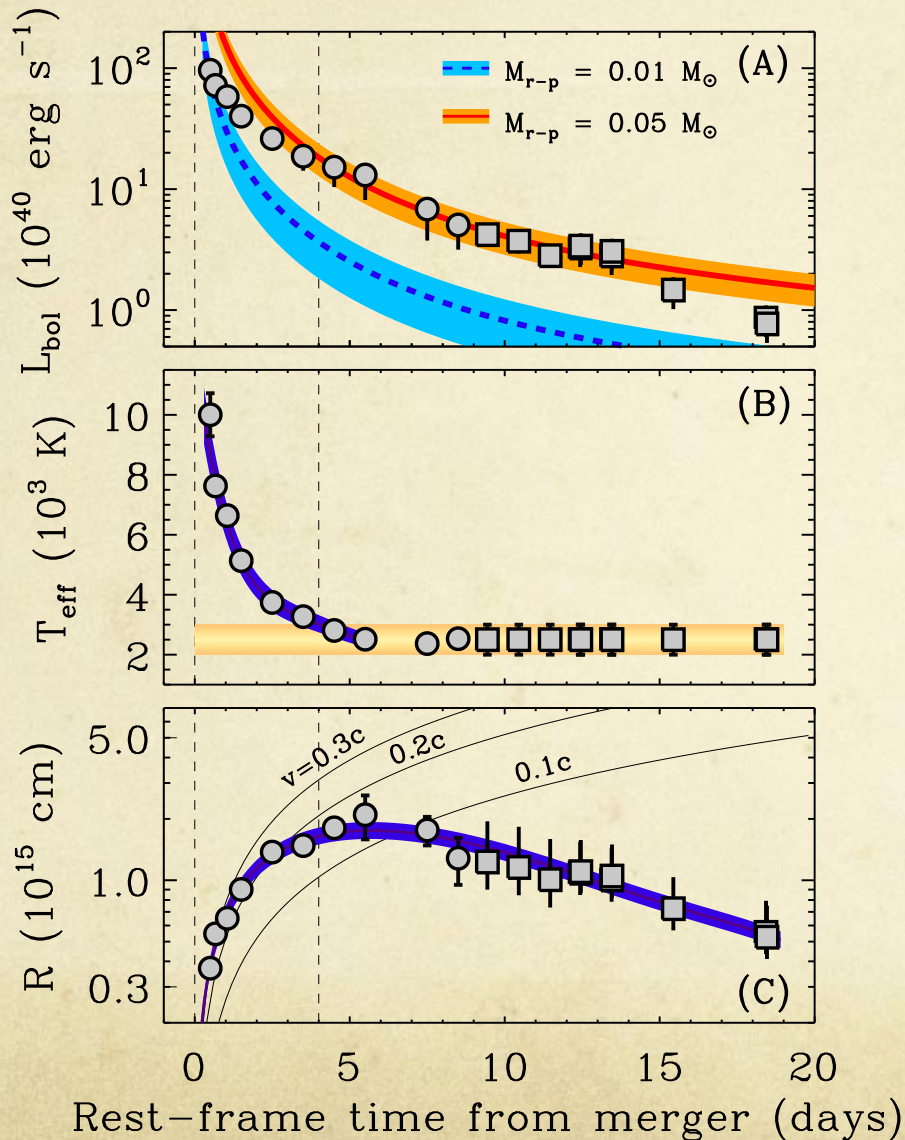


SSS17a: Physical Parameters



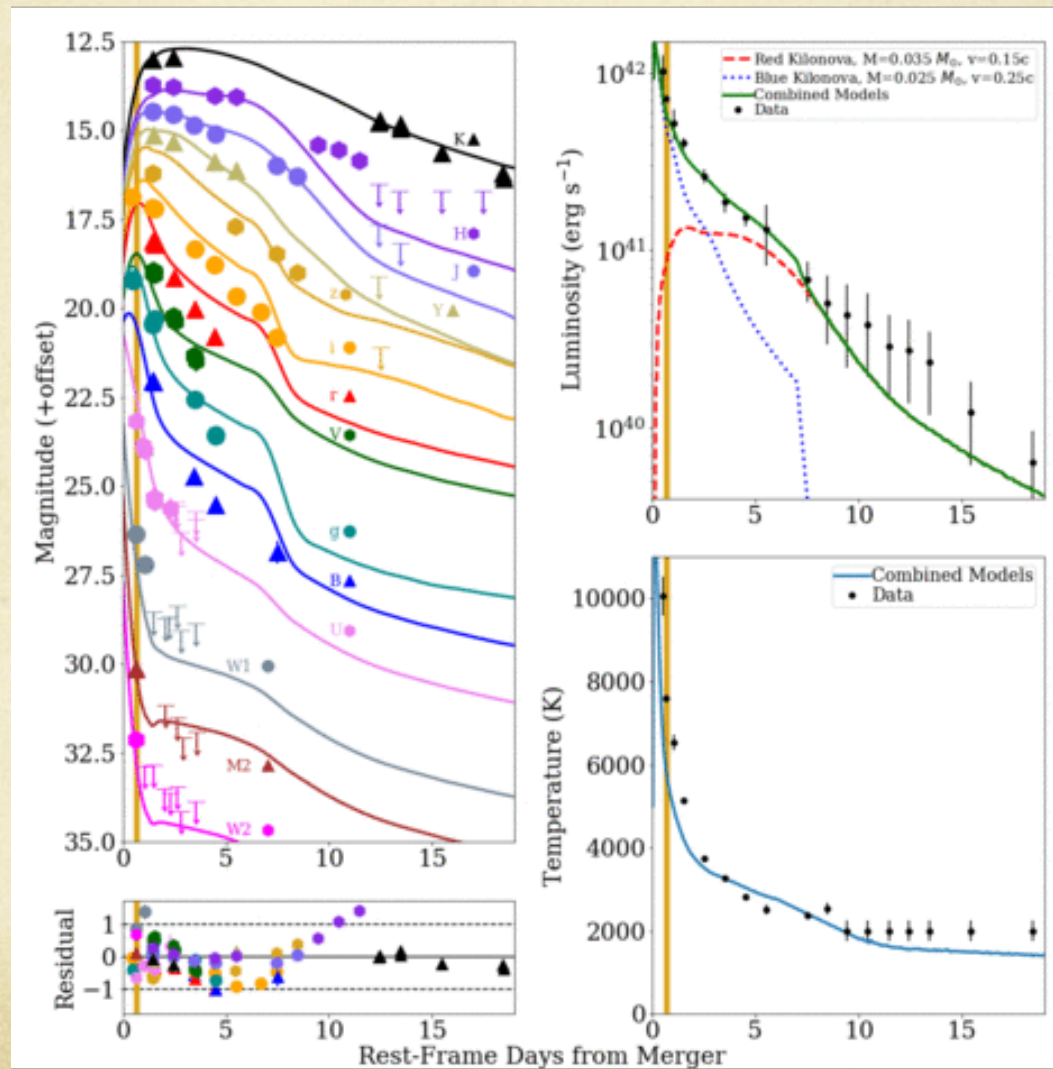
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SSS17a: Physical Parameters



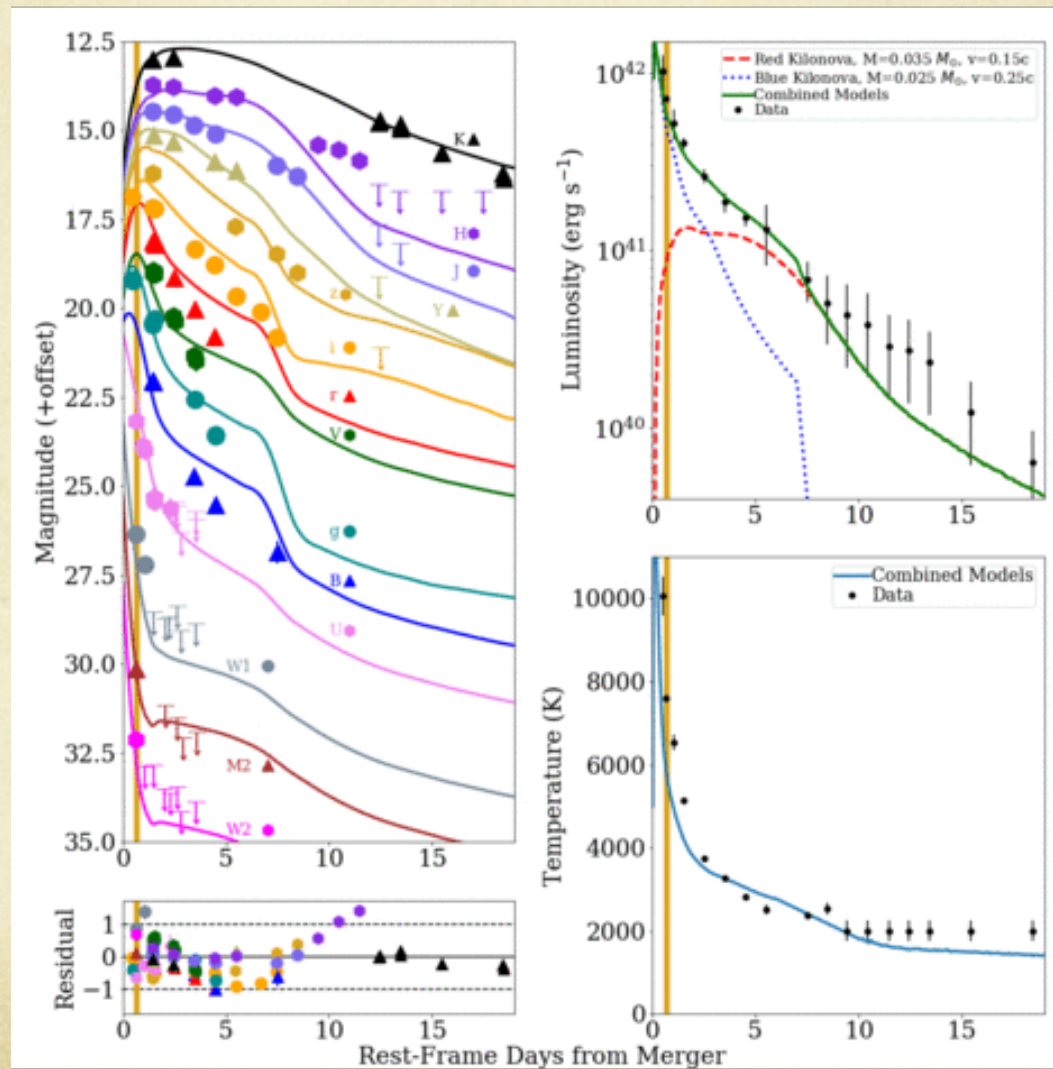
- r-process heating goes as $\tau^{1.3}$
- two components are required
- temperature evolution consistent with lanthanide recombination

Kilonova Modeling



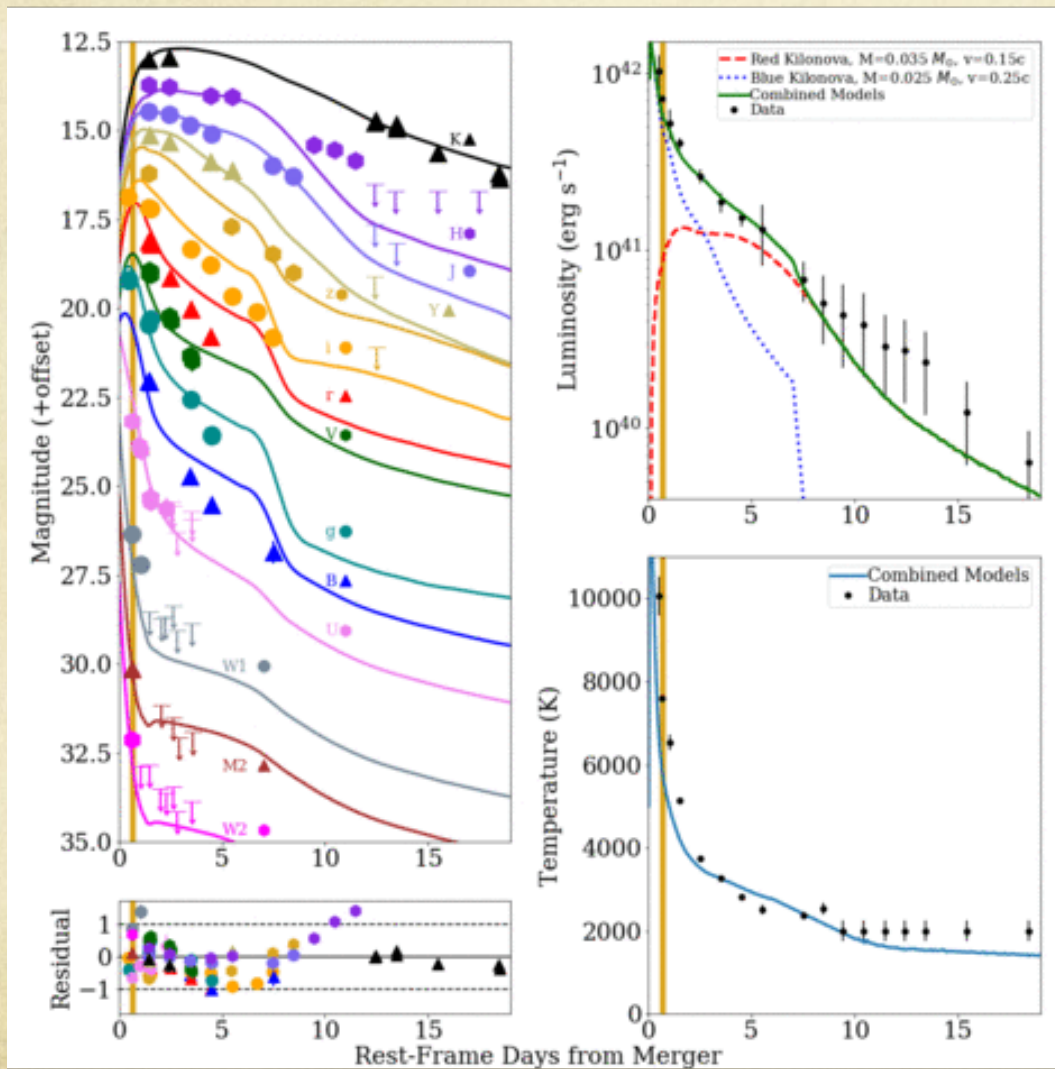
Kilpatrick et al. (2017), Kasen et al. (2017)

Kilonova Modeling



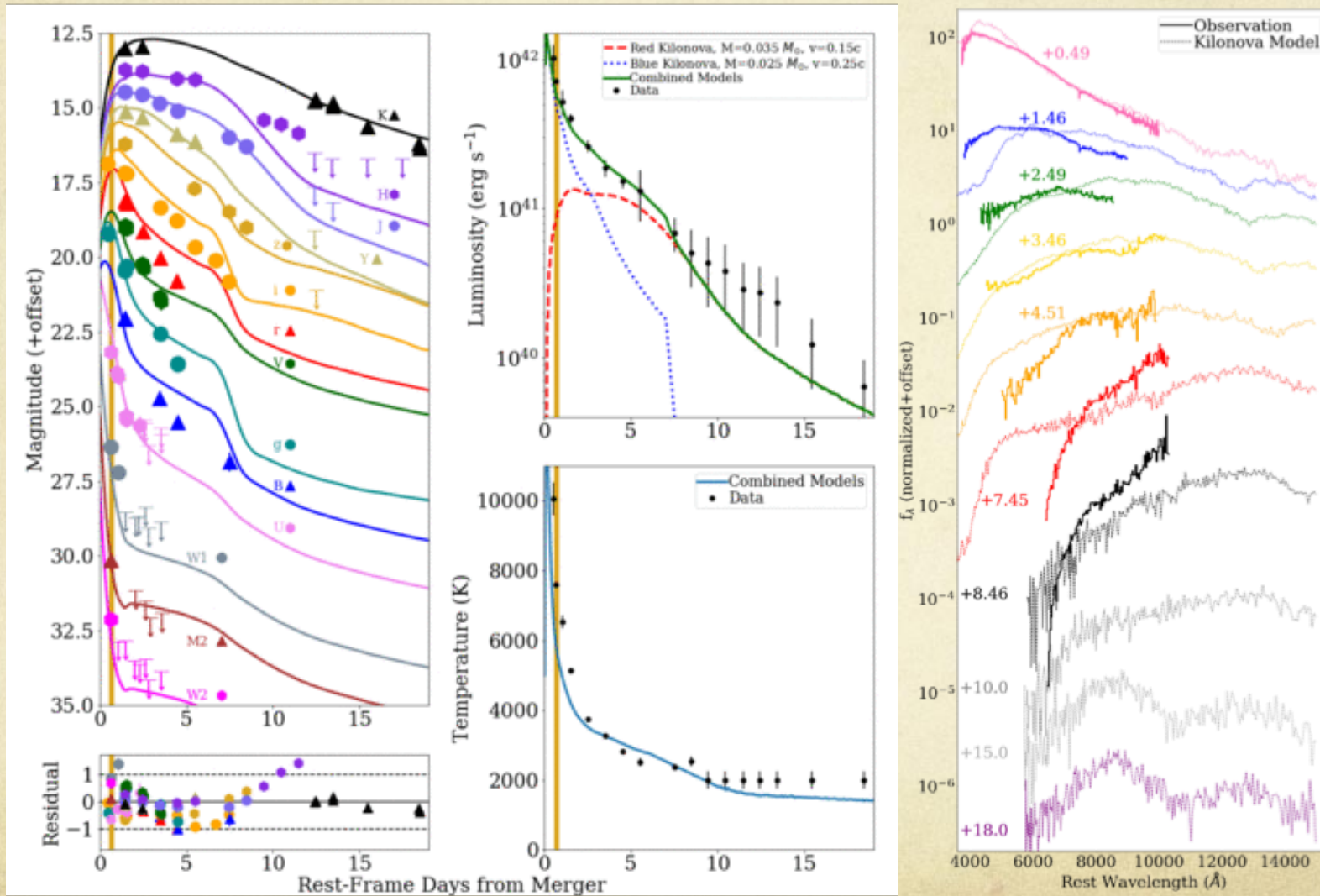
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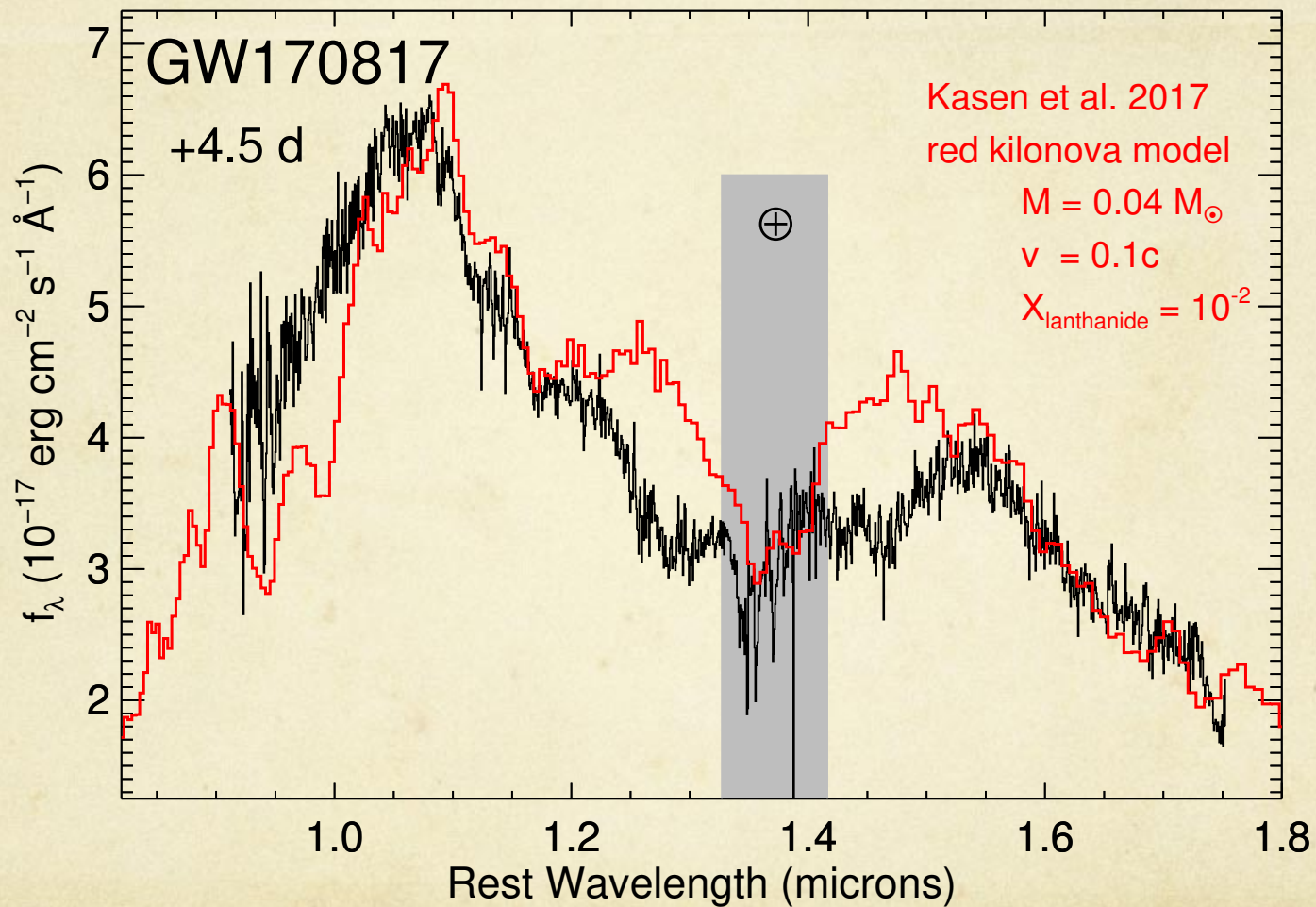
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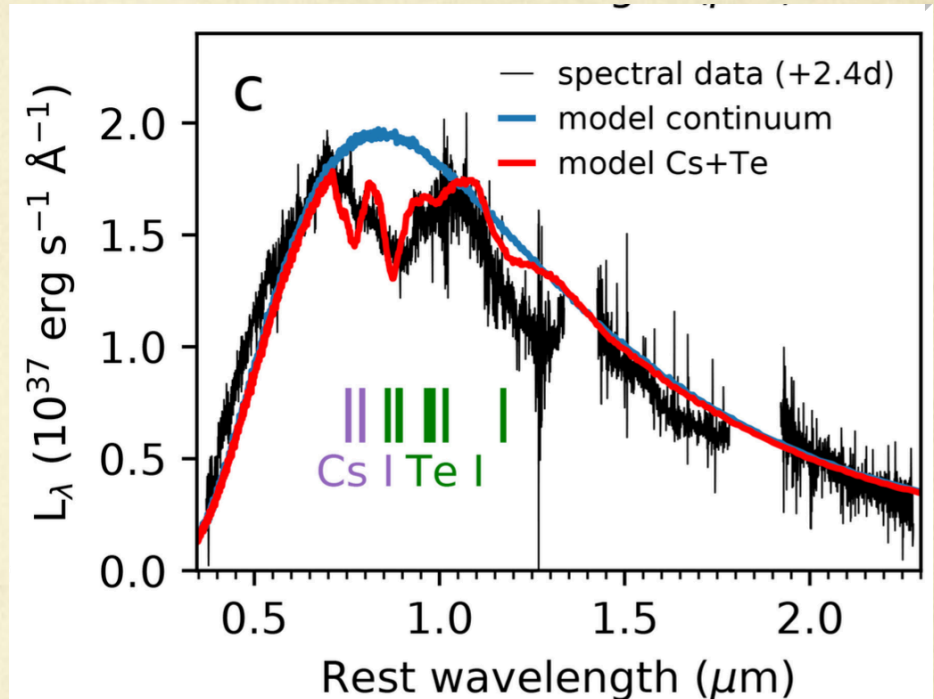
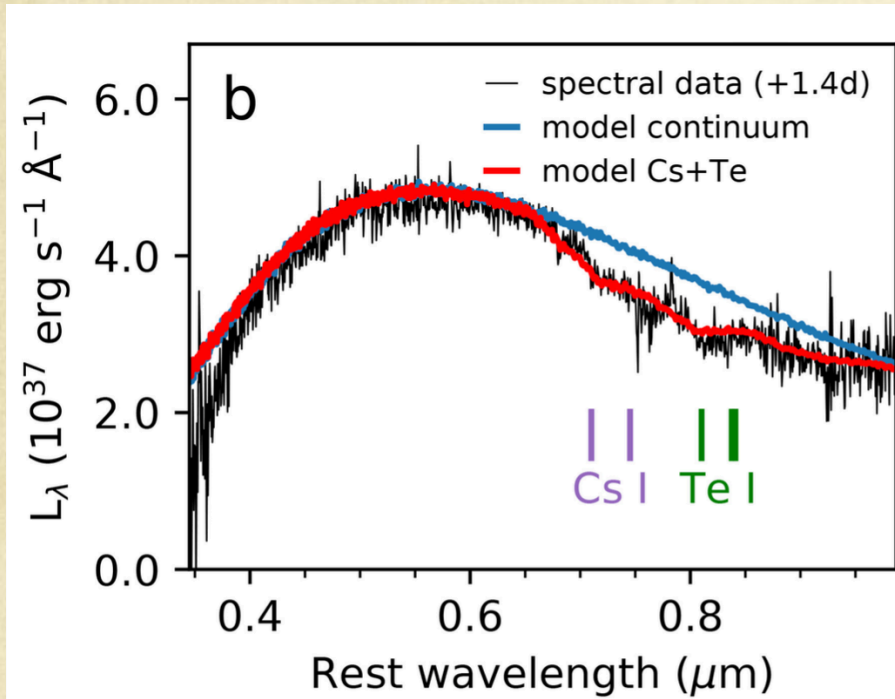


Kilpatrick et al. (2017), Kasen et al. (2017)

1 Micron Spectral Feature



Elemental Spectral Feature?

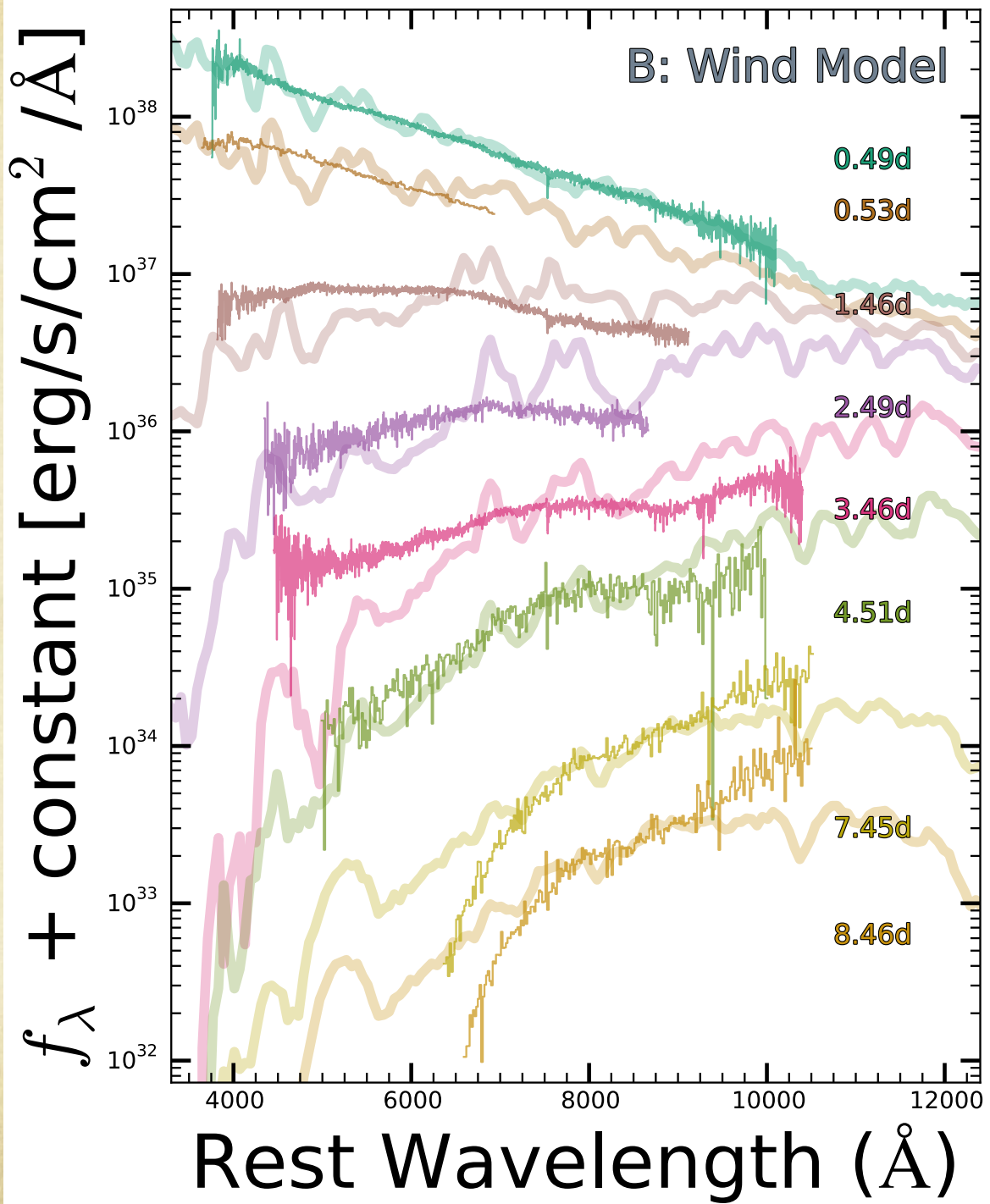


Consistent with an r-process powered transient in which approximately 0.05 solar masses of lanthanide-rich material is ejected

Consistent with an r-process powered transient in which approximately 0.05 solar masses of lanthanide-rich material is ejected

Nature of the blue component

(See Piro & Kollmeier 2017)

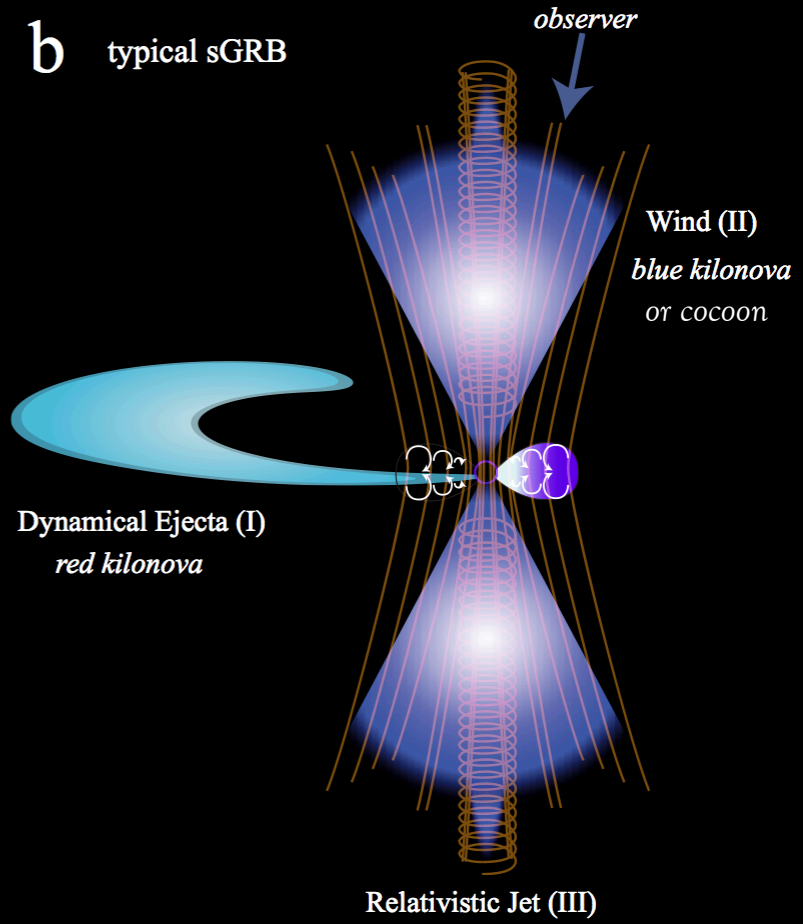
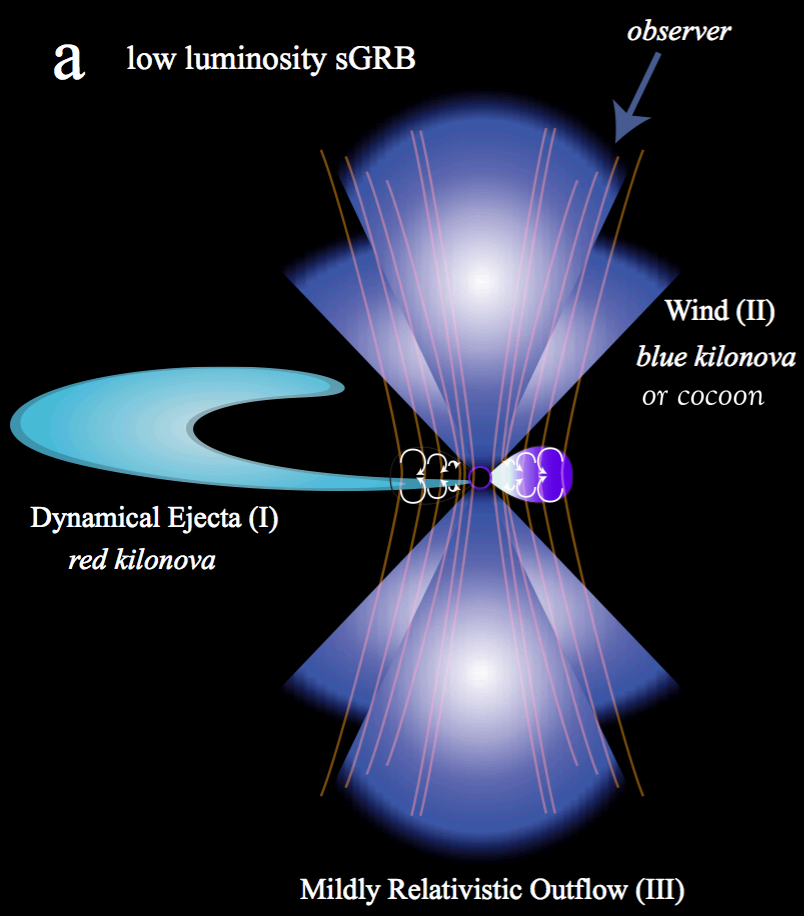


Consistent with an r-process powered transient in which approximately 0.05 solar masses of lanthanide-rich material is ejected

Nature of the blue component

(See Piro & Kollmeier 2017)

Nature of the GRB/X-rays/radio



Murguia-Berthier et al. 2017

Multi-messenger Astronomy

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Difference is
between -3×10^{-15}
and $+7 \times 10^{-16}$
times the speed
of light

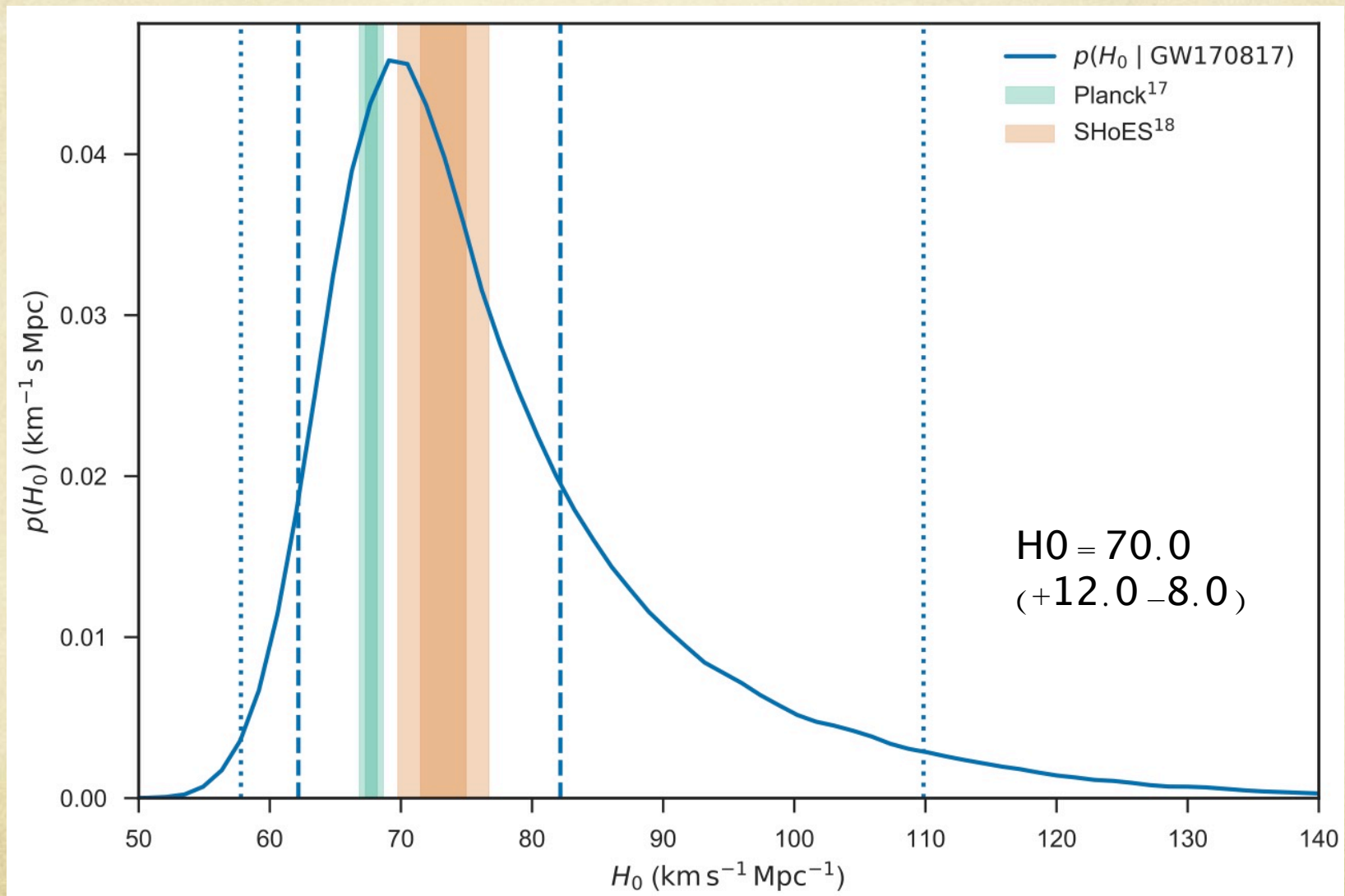
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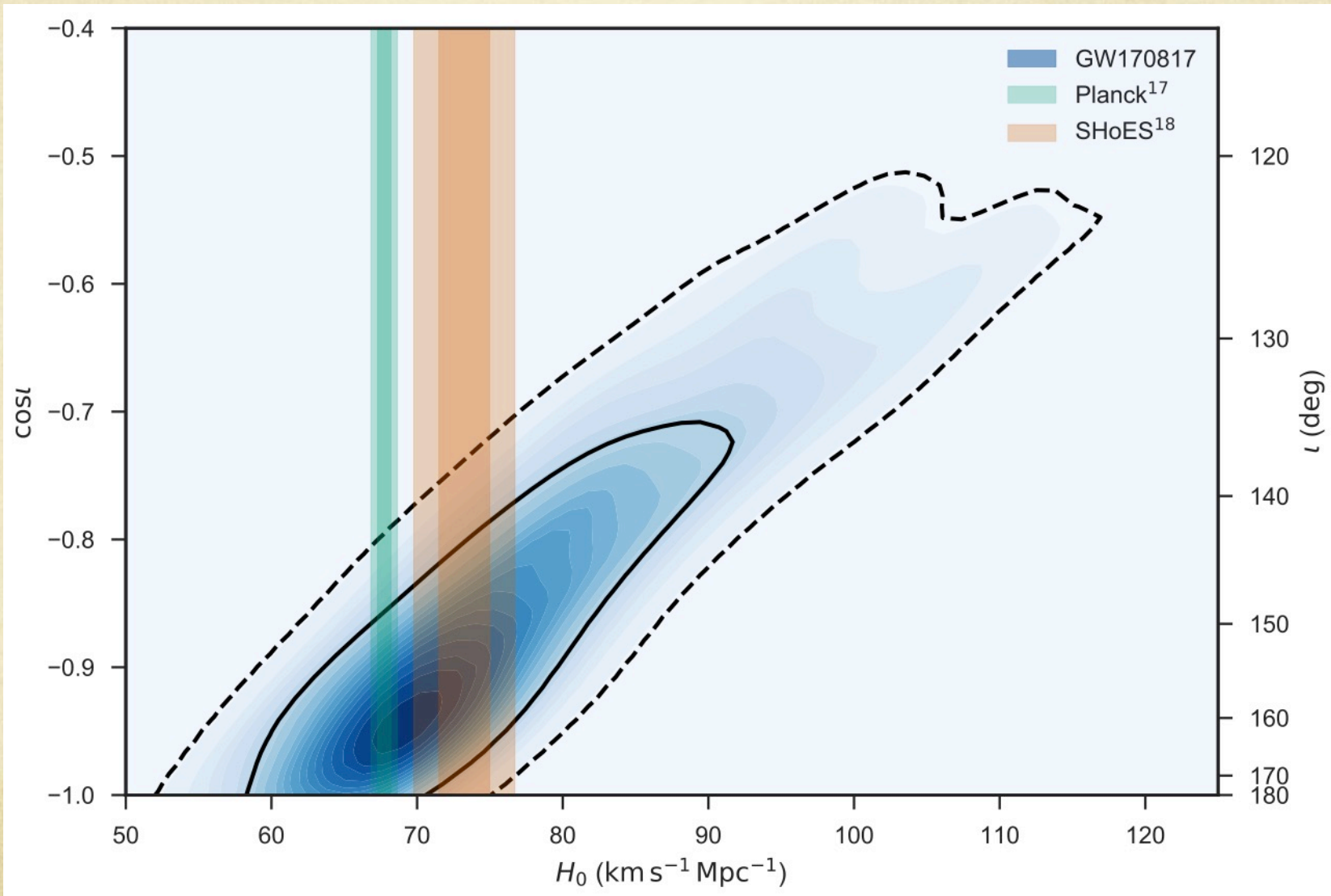
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“Standard Siren” Hubble Constant



“Standard Siren” Hubble Constant



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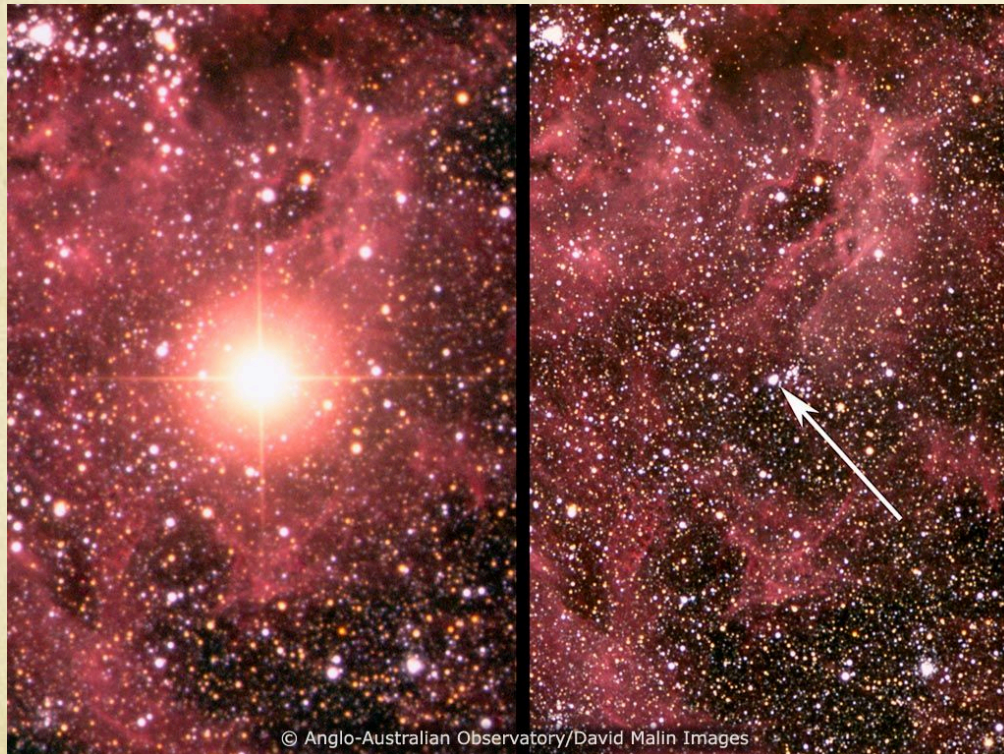
Summary

We discovered the first EM counterpart to a gravitational wave source

This begins the era of multi-messenger gravitational-waves astronomy, with implications across many areas of astronomy

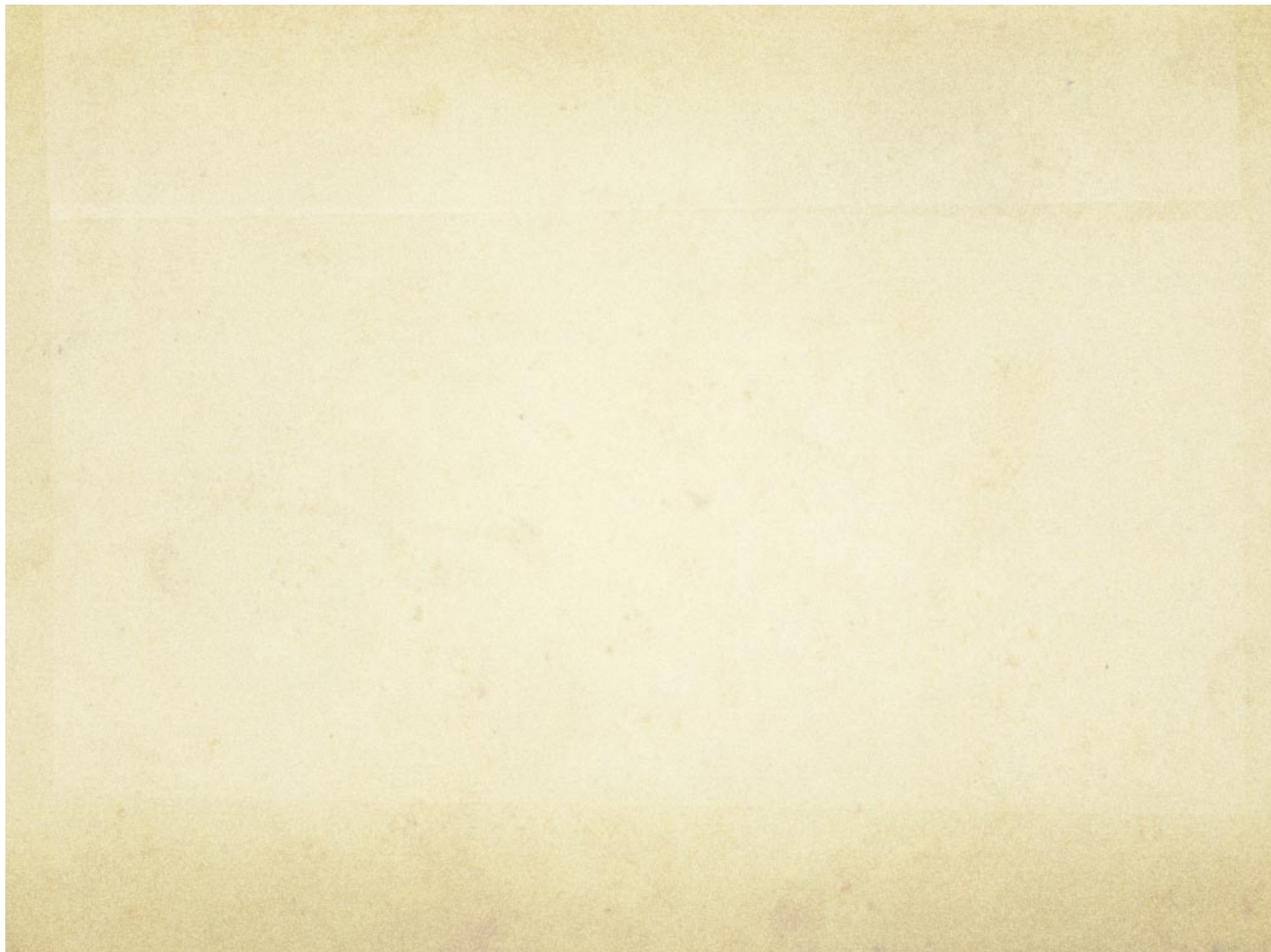
Multi-messenger Astronomy Success Stories

SN1987A:
photons + neutrinos



SSS17A:
photons + GWs





Multi-messenger Astronomy

Motivation

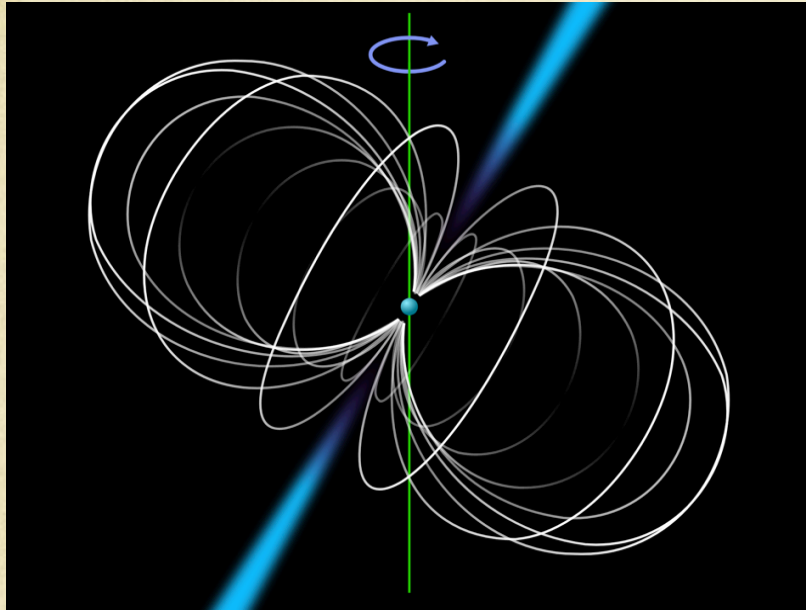
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SSS17a: Follow-Up

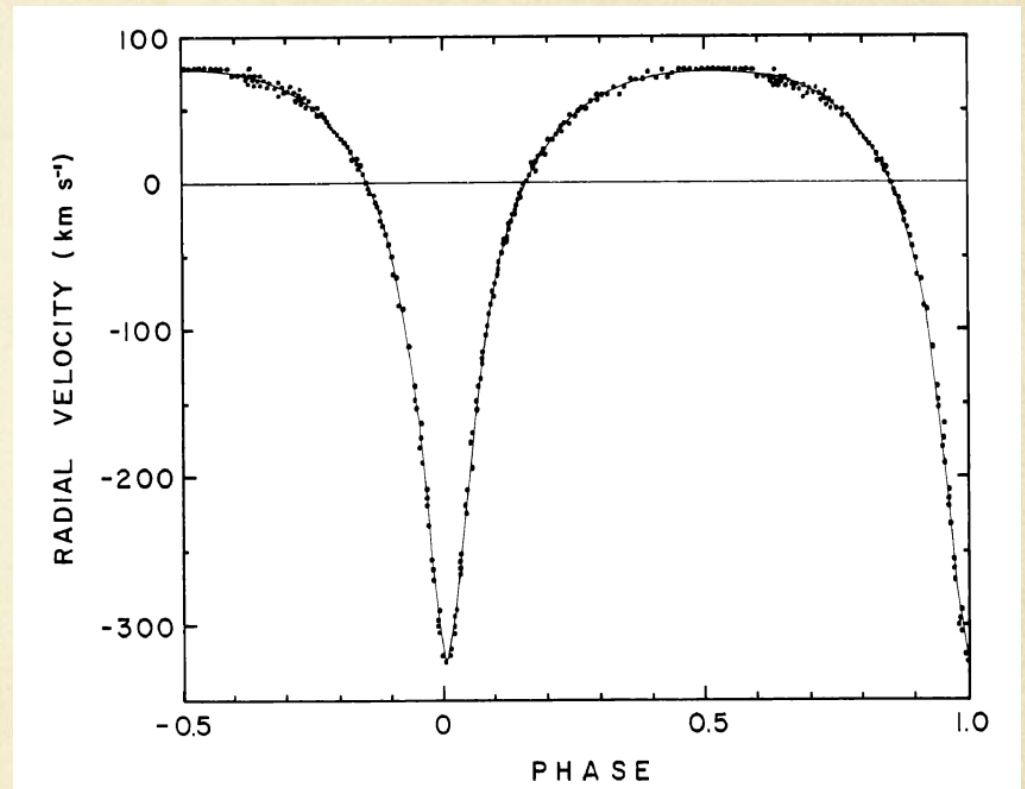


Hulse-Taylor binary



Measured Orbital Parameters for PSR B1913+16

Fitted Parameter	Value
$a_p \sin i$ (s)	2.3417725 (8)
e	0.6171338 (4)
T_0 (MJD)	52144.90097844 (5)
P_b (d)	0.322997448930 (4)
ω_0 (deg)	292.54487 (8)
$\langle \dot{\omega} \rangle$ (deg/yr)	4.226595 (5)
γ (s)	0.0042919 (8)
\dot{P}_b (10^{-12} s/s)	-2.4184 (9)



Multi-messenger Astronomy

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Gravitational Wave data only provides specific information

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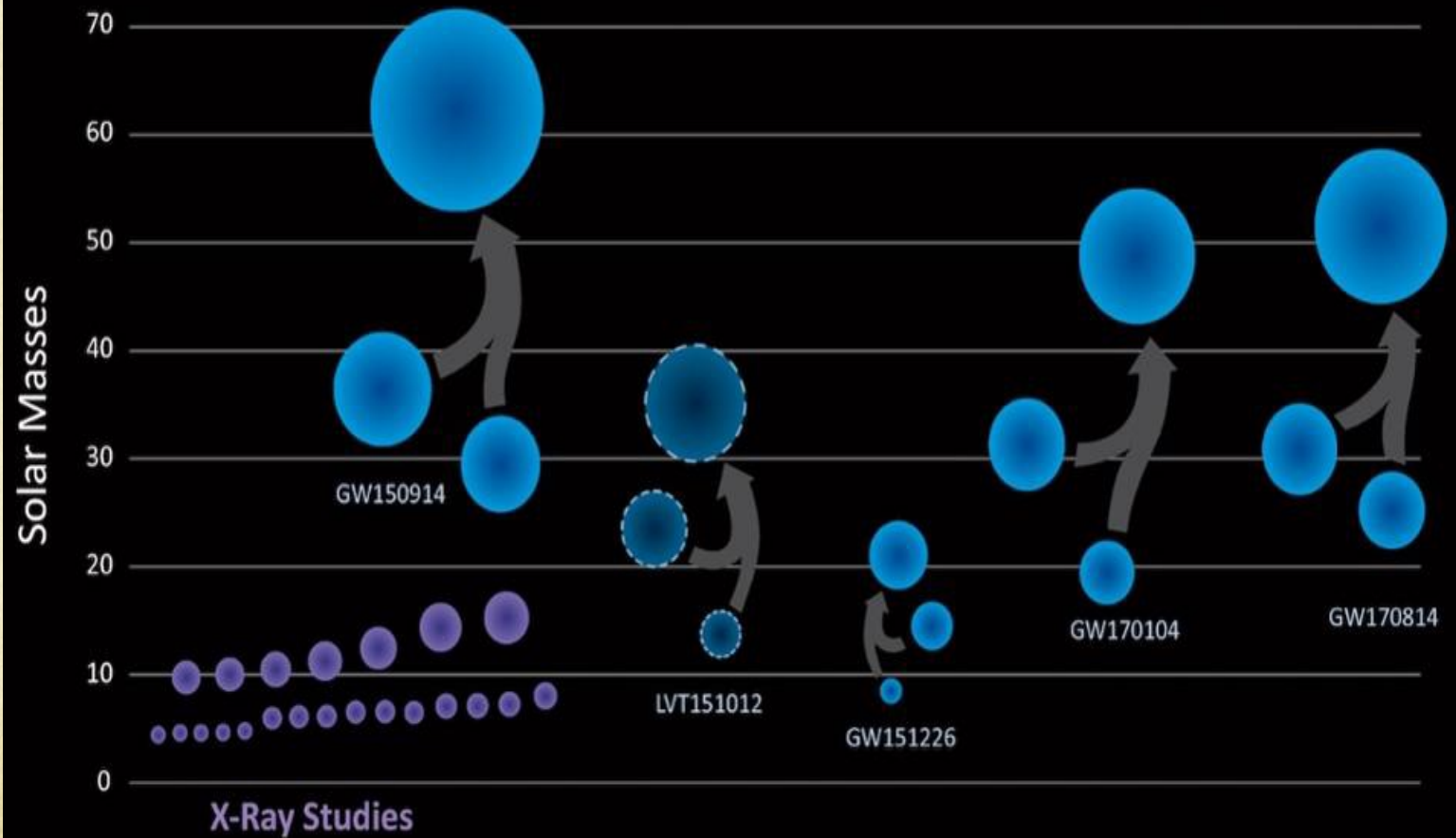
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GW150814

Black Holes of Known Mass



Precise Localization