

The Hubble Constant

It only took 13 years!

It's been a long long road...



DR3 Is DONE!!!!

v3x



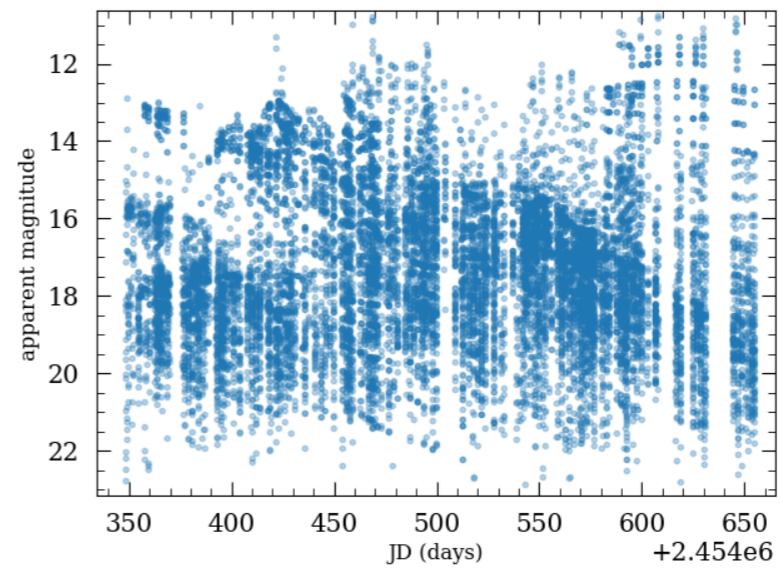
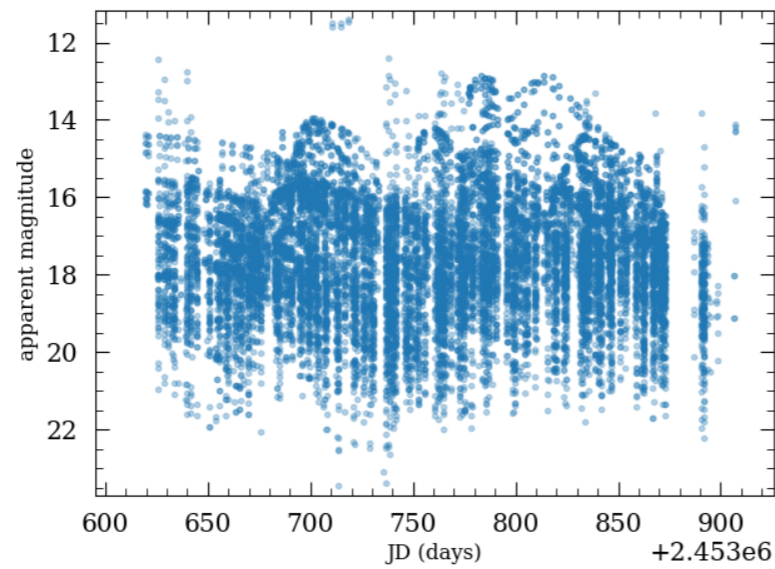
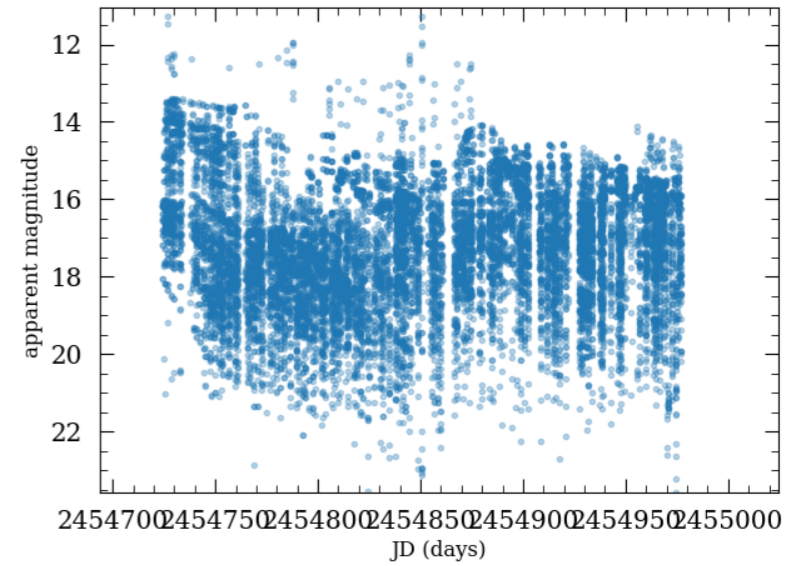
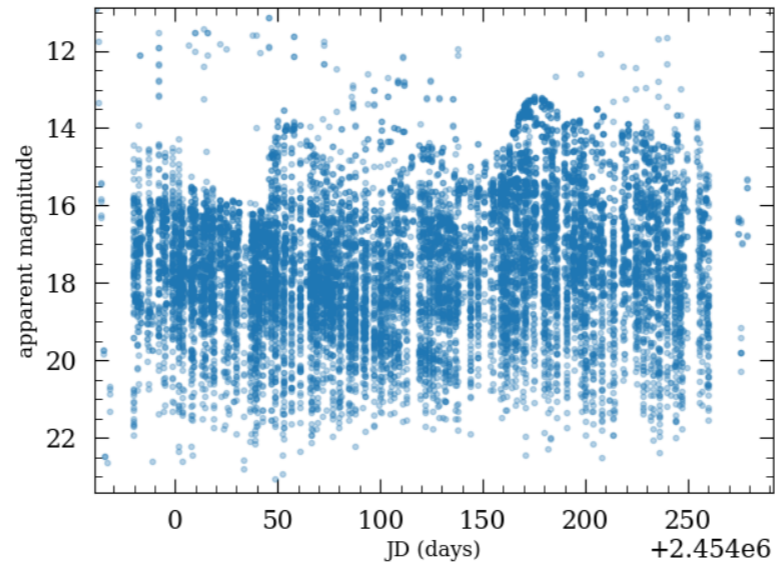
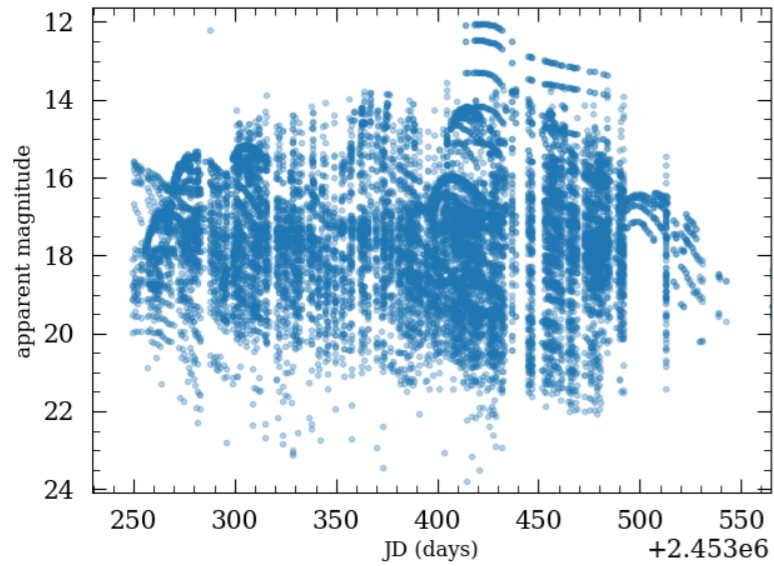
The Carnegie Supernova Project: Third Photometry Data Release of Low-Redshift Type Ia Supernovae and Related Subtypes

Kevin Krisciunas¹, Carlos Contreras², Mark M. Phillips², Abdo Campillay², Nidia Morrell², Maximilian D. Stritzinger^{2,3}, Luis Boldt², Christopher R. Burns⁴, Gastón Folatelli^{2,5}, Consuelo Gonzalez², Sergio Castellón², Wojtek Krzeminski², Miguel Roth², Francisco Salgado², Darren L. Depoy¹, Mario Hamuy⁶, Wendy L. Freedman^{3,7}, Eric Y. Hsiao^{2,5}, Barry F. Madore^{3,8}, Jennifer L. Marshall¹, Sven Eric Persson³, Jean-Philippe Rheault¹, Nicholas B. Suntzeff¹, Steven Villanueva¹, Weidong Li^{9,10}, and Alexei V. Filippenko⁹

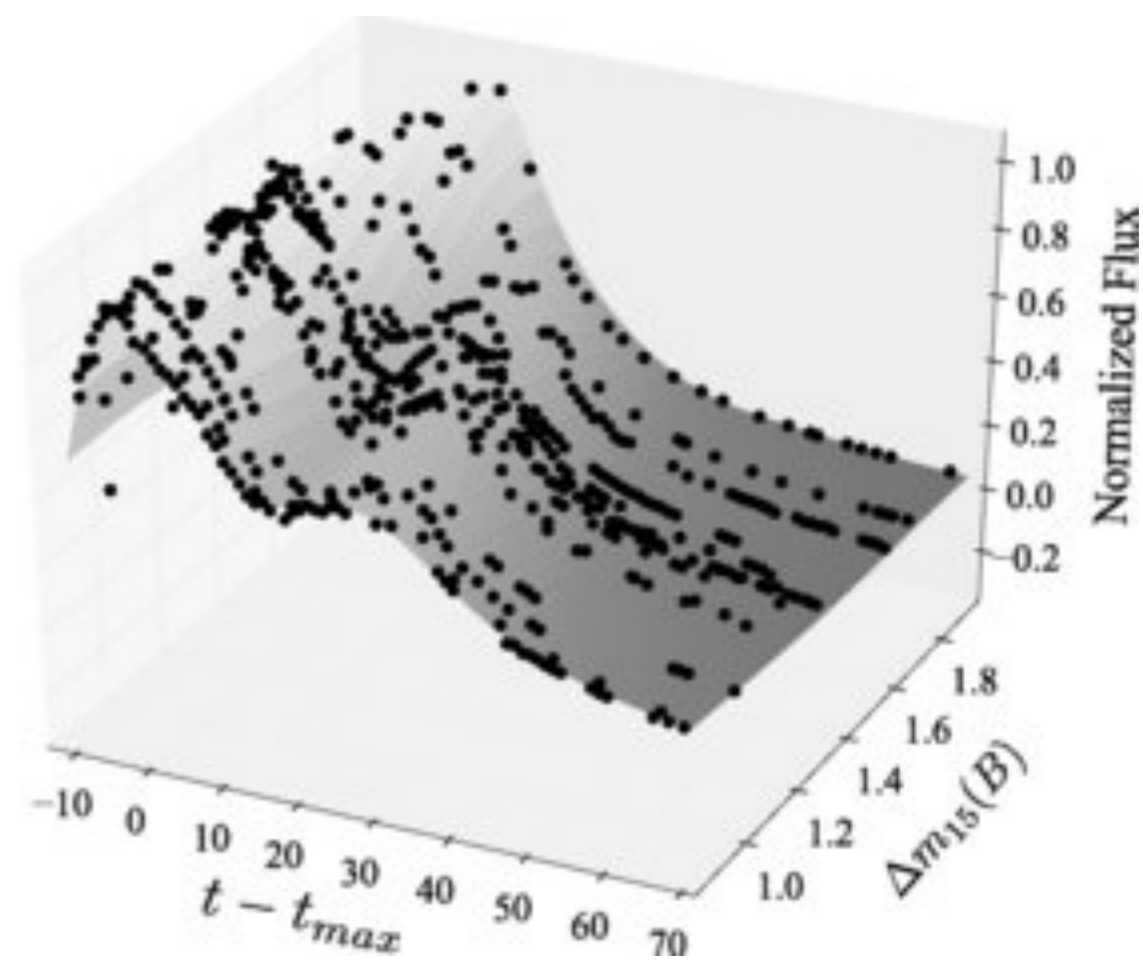
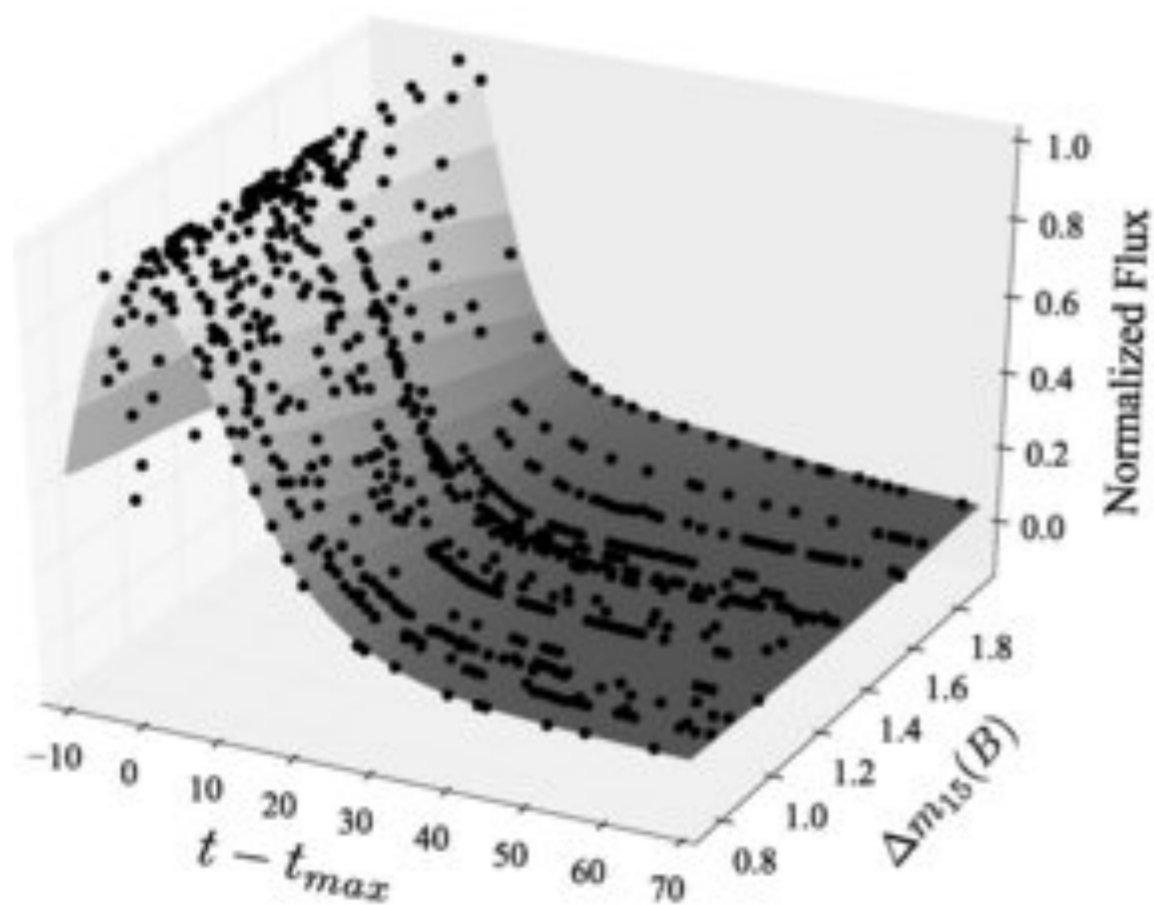
ABSTRACT

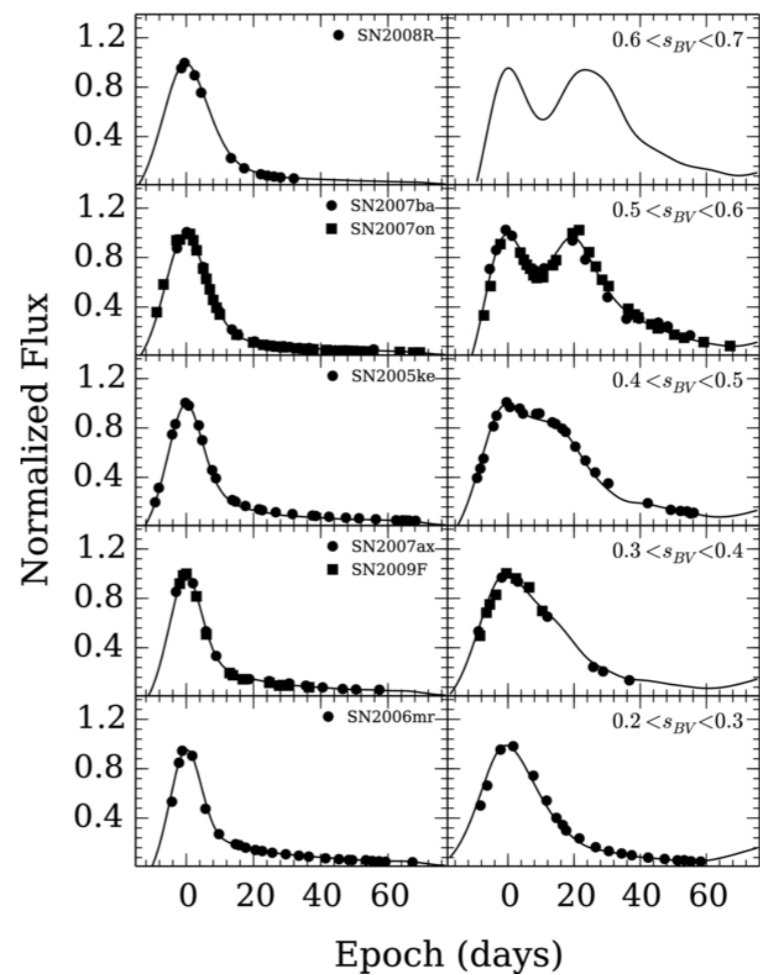
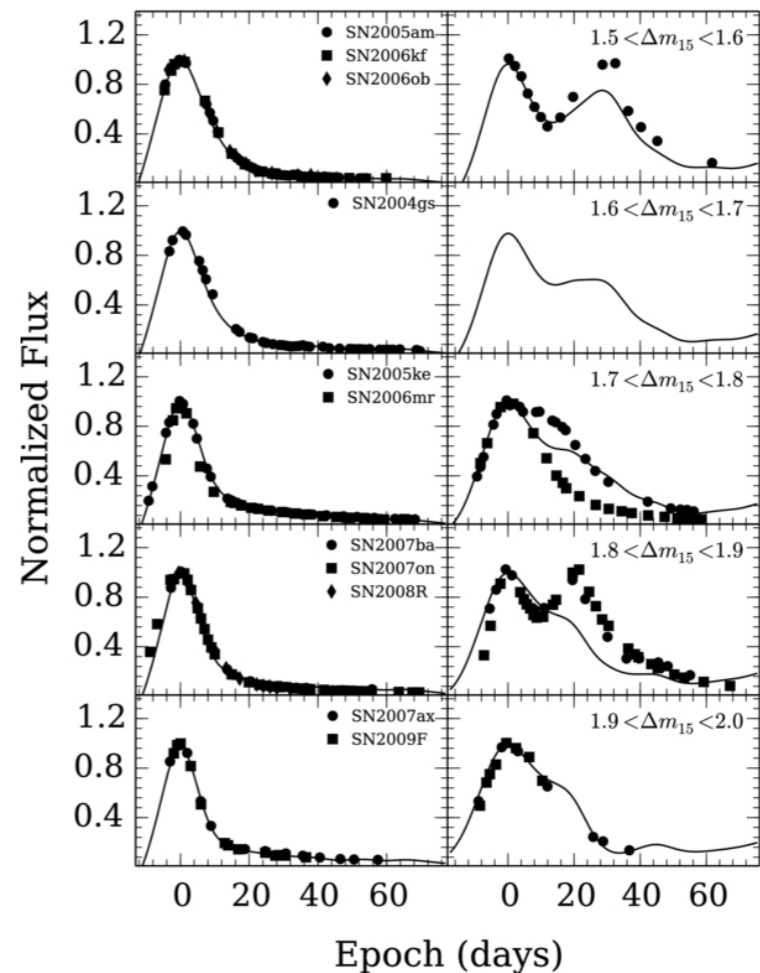
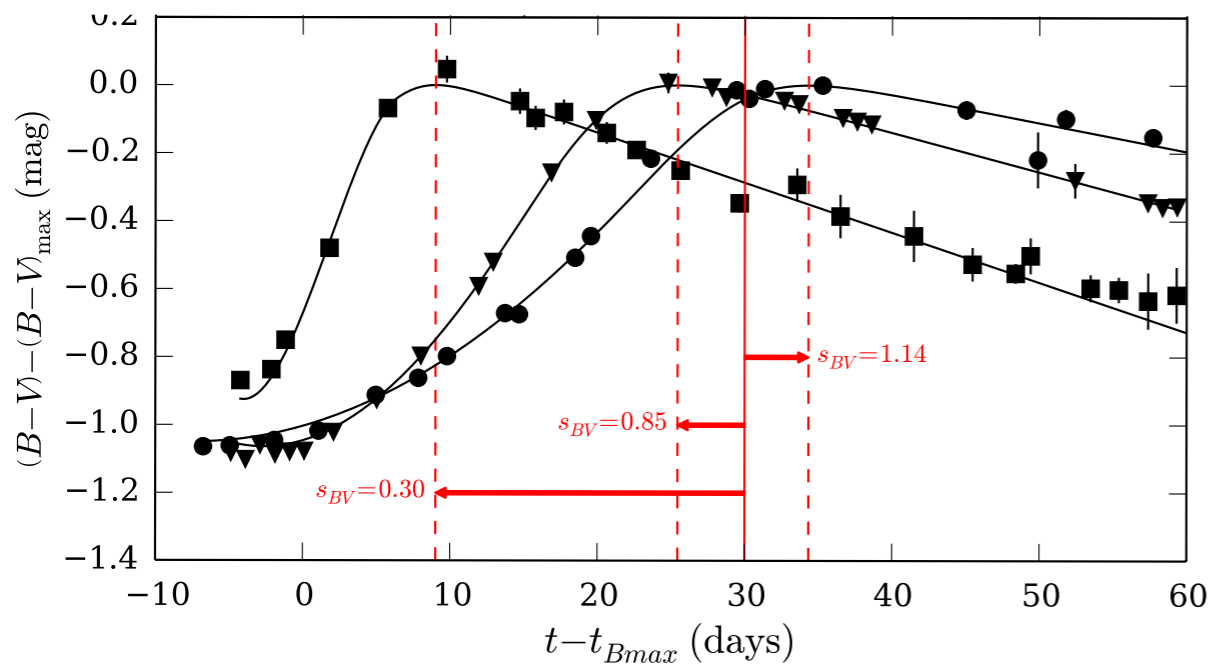
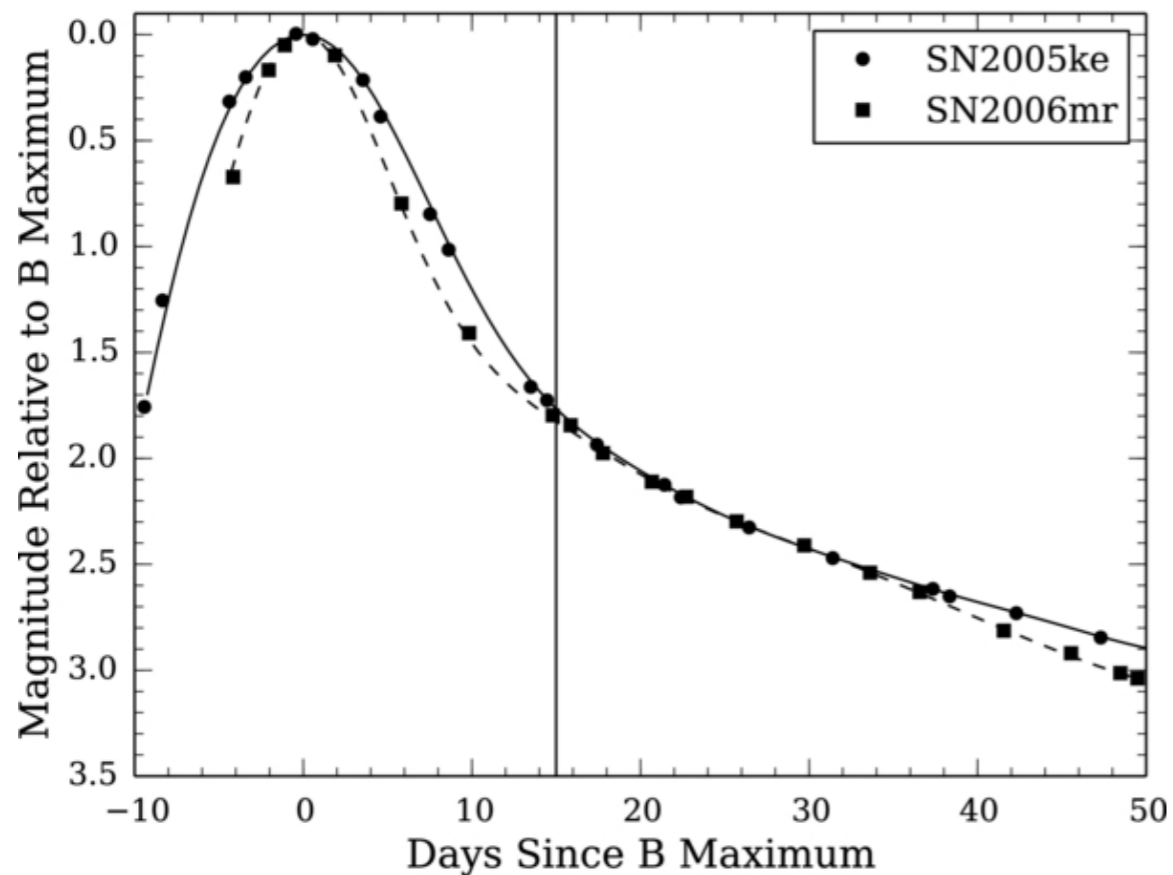
[Version of August 1, 2016. Not for general distribution. (**We are ready for the real abstract that Mark has promised.**)] We present previously unpublished photometry of all low redshift supernovae observed by the Carnegie Supernova Project (2004-2009). In all, 134 Type Ia supernovae were observed in optical bandpasses. NNN were measured in one or more near-infrared bands....

Observed Magnitudes for Campaigns 1-5



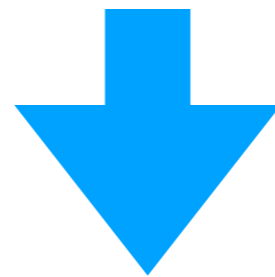
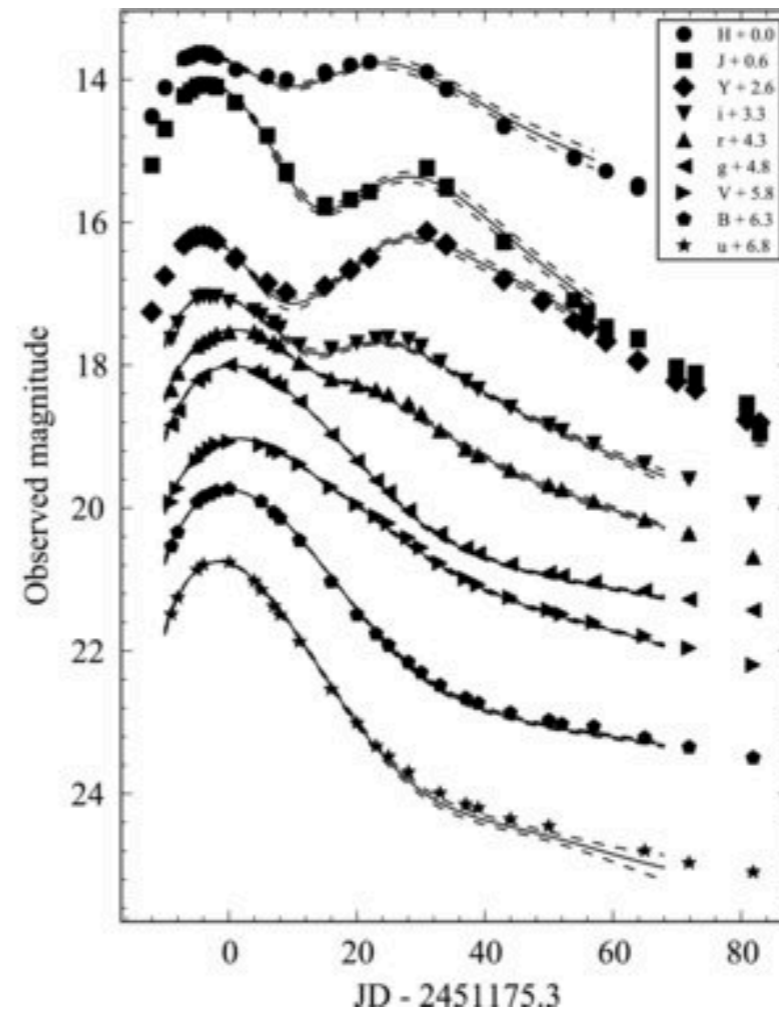
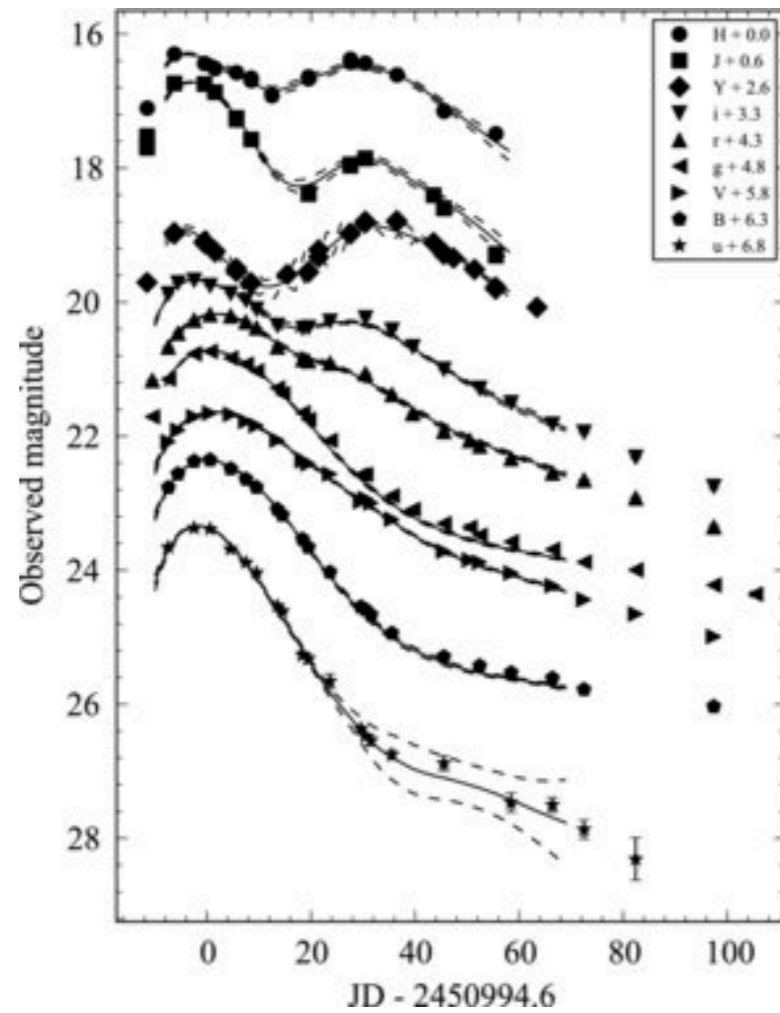
2.5-D Gaussian Process Surface





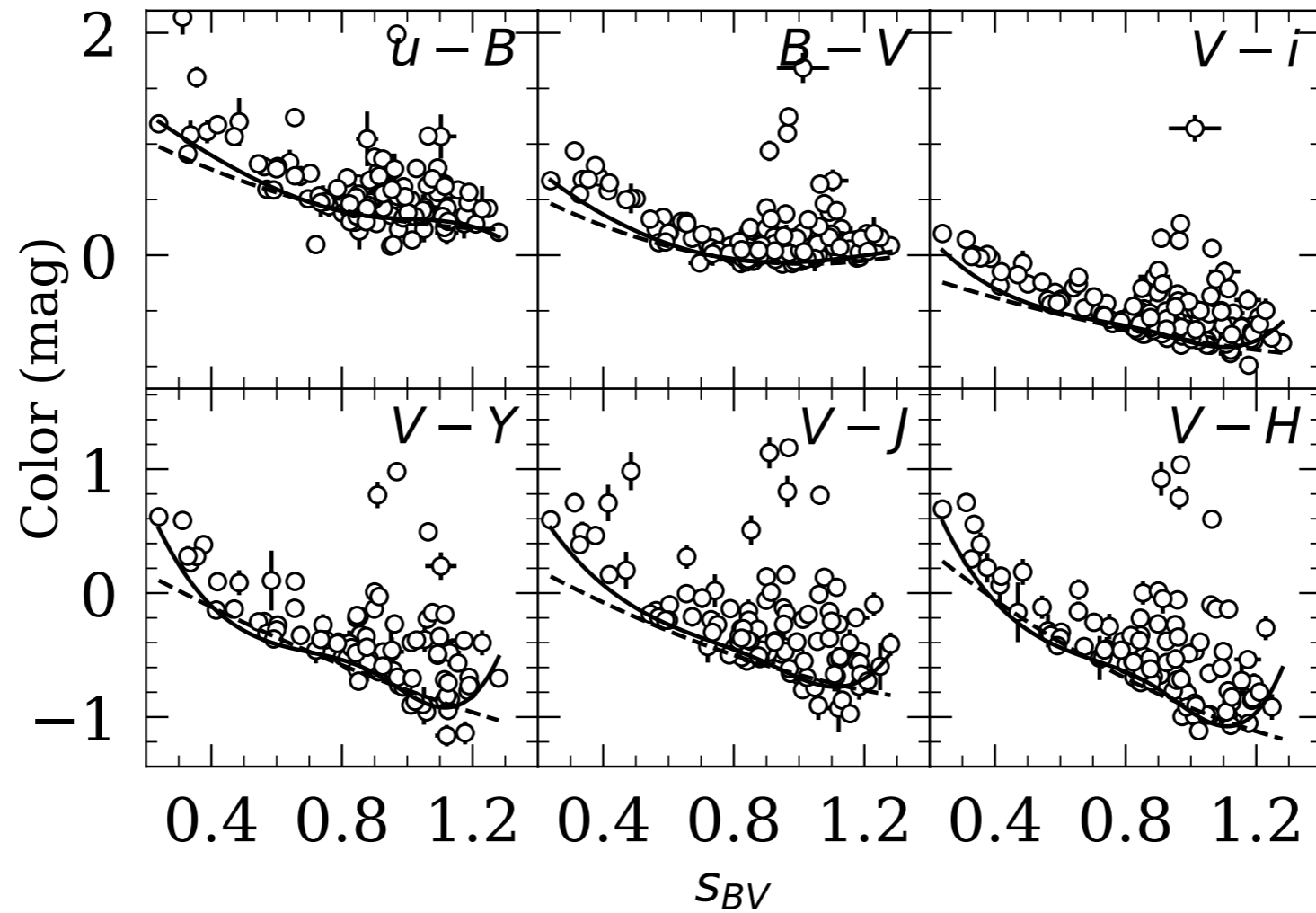


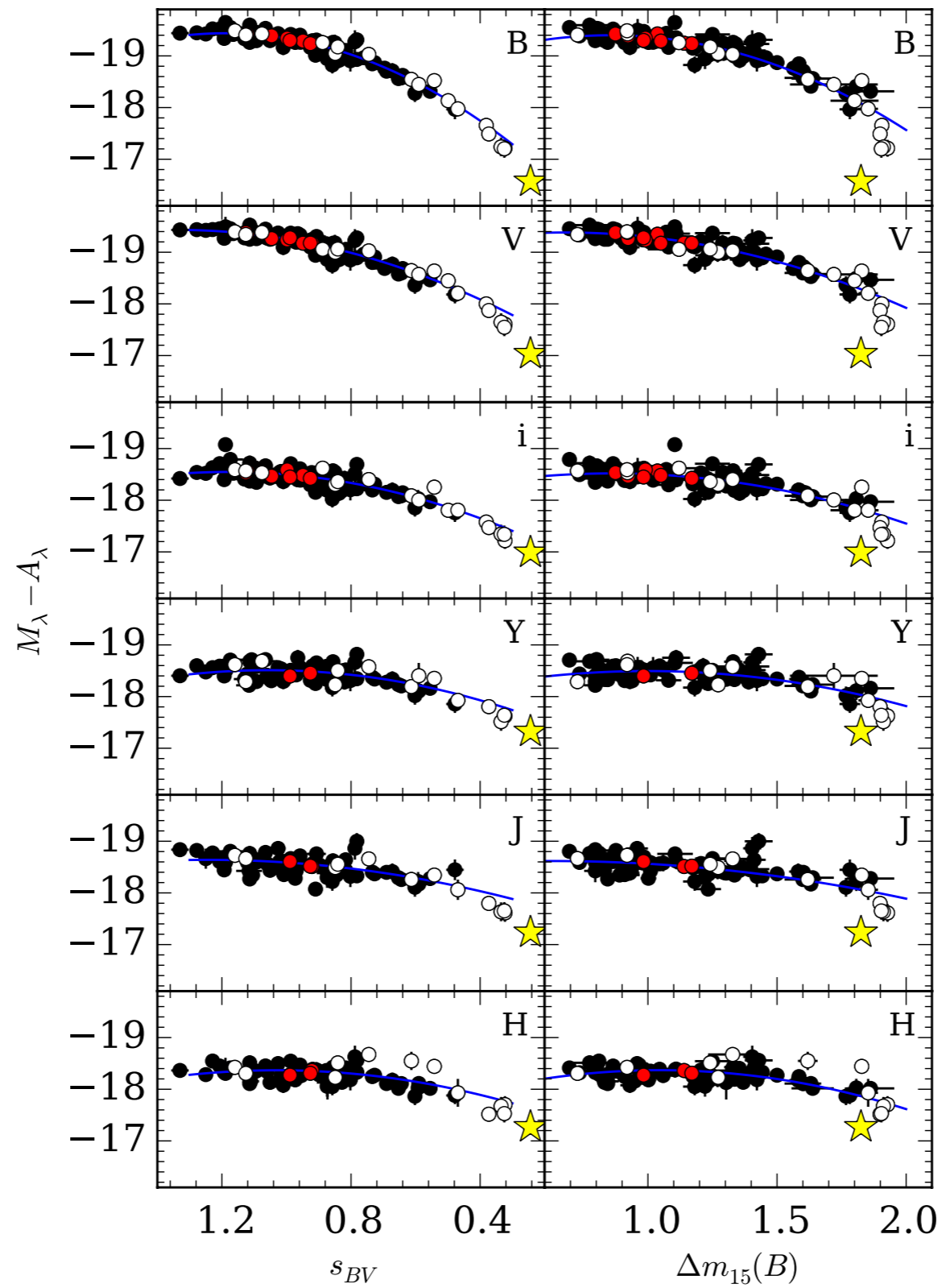
Fit for s_{BV} , $T_{\max}(B)$, $m_{\max}(\lambda)$





Intrinsic Colors and Extinctions



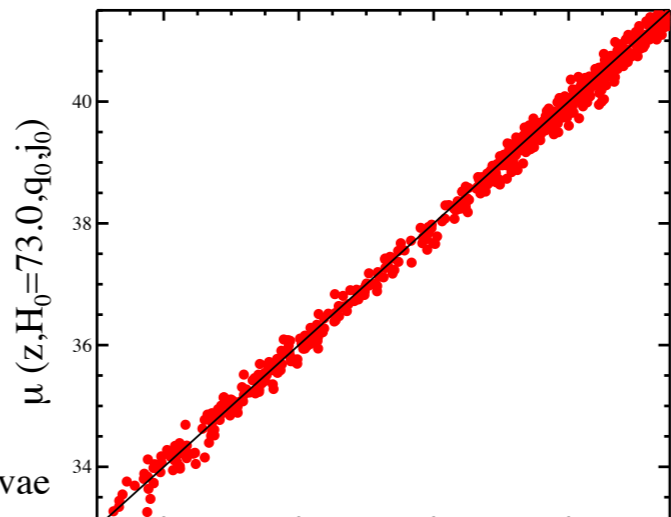


Assuming
 $H_0 = 72 \text{ km/s/Mpc}$

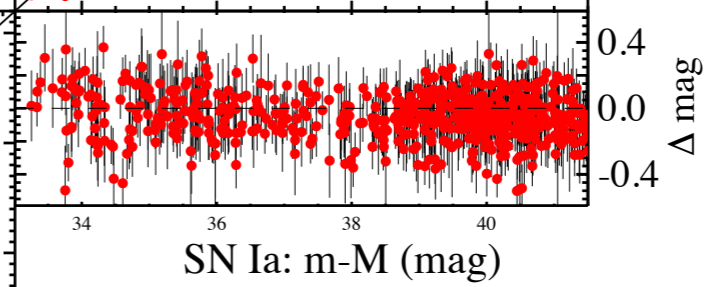
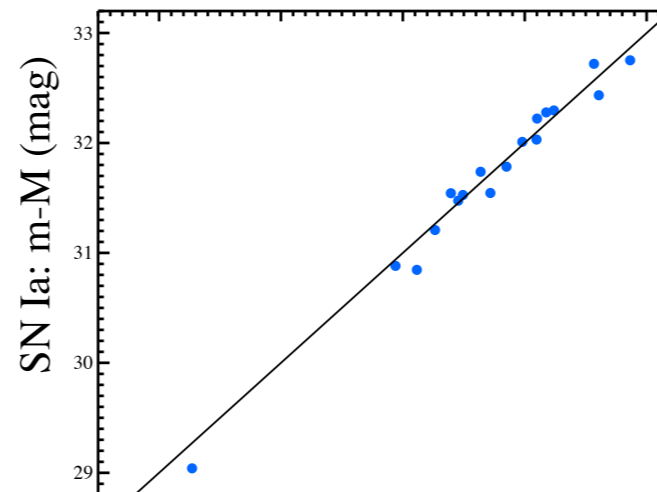


The Distance Ladder

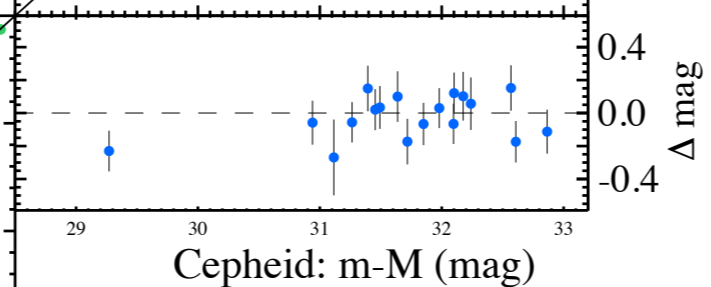
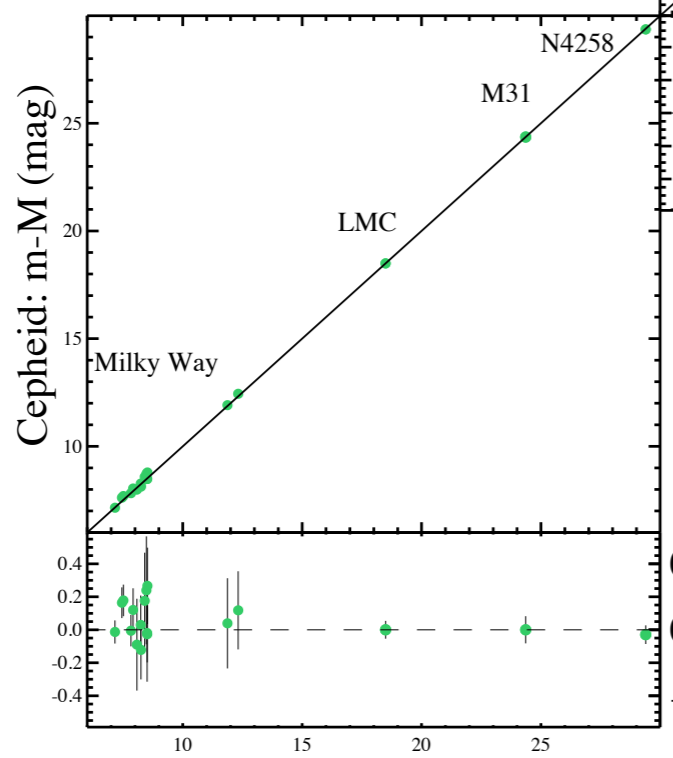
Type Ia Supernovae \rightarrow redshift(z)



Cepheids \rightarrow Type Ia Supernovae

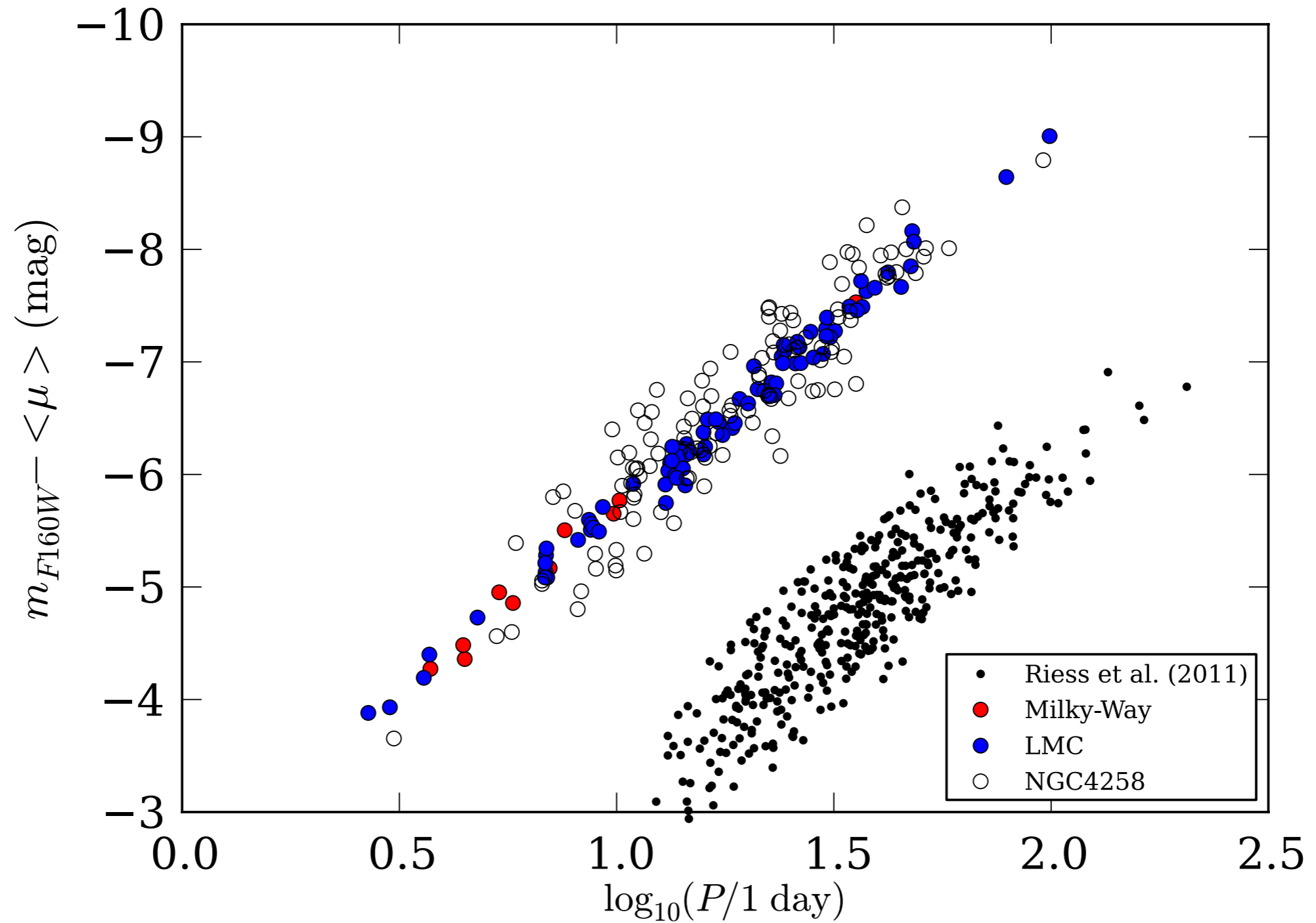


Geometry \rightarrow Cepheids



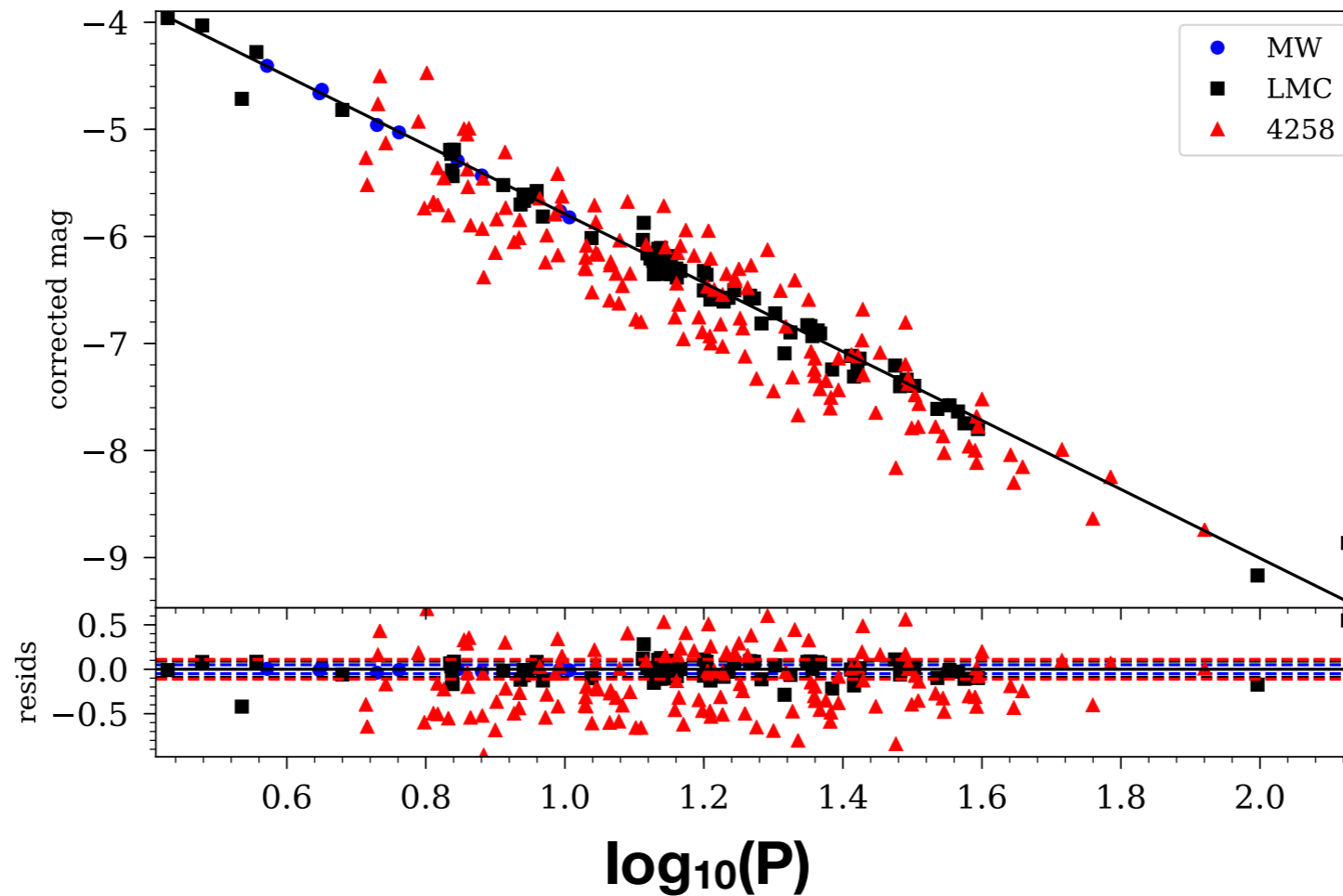
Geometry: $5 \log D$ [Mpc] + 25

Calibrating the Ia's with Cepheid distances

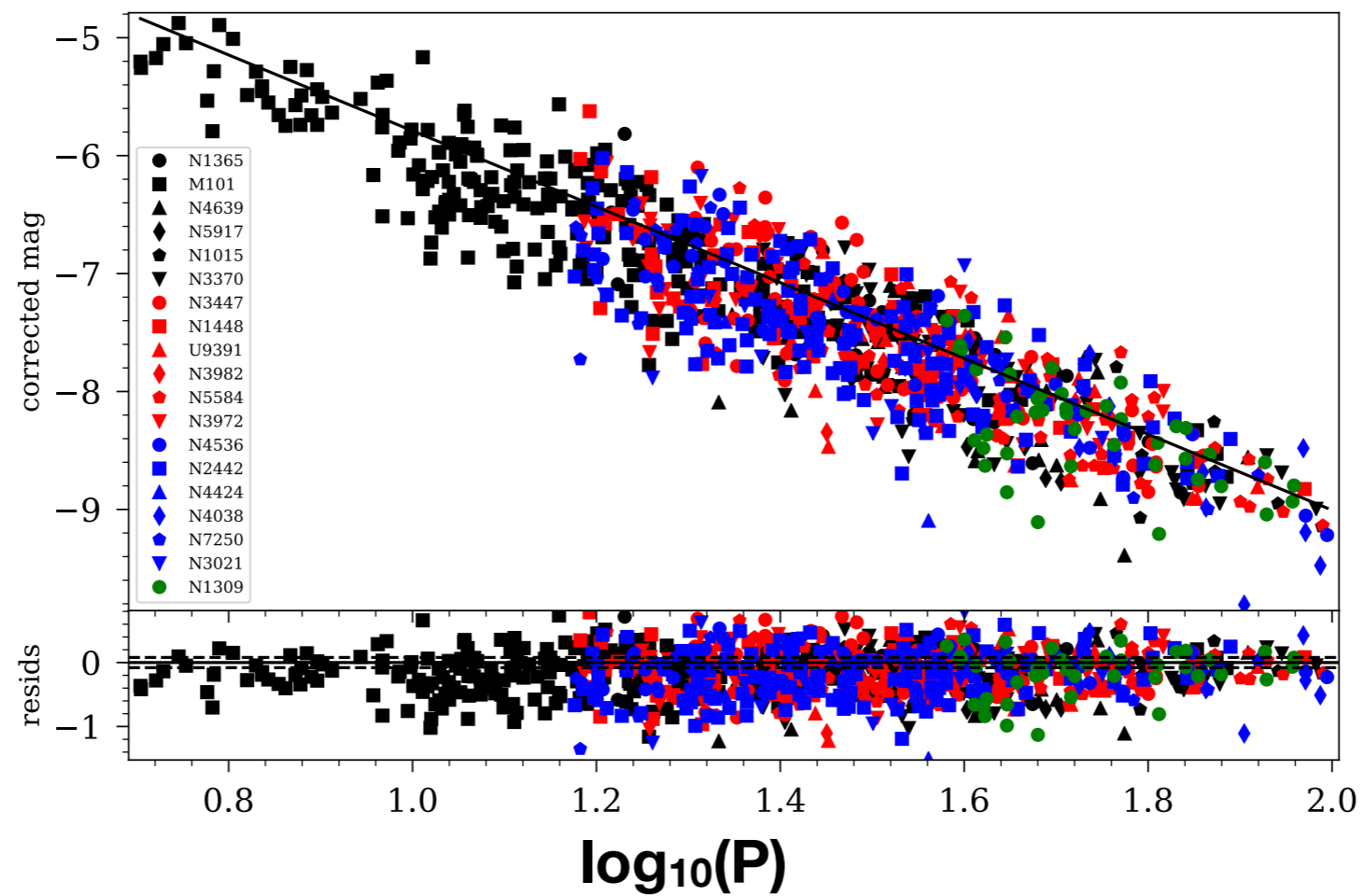


Anchors:

MW
LMC
NGC 4258

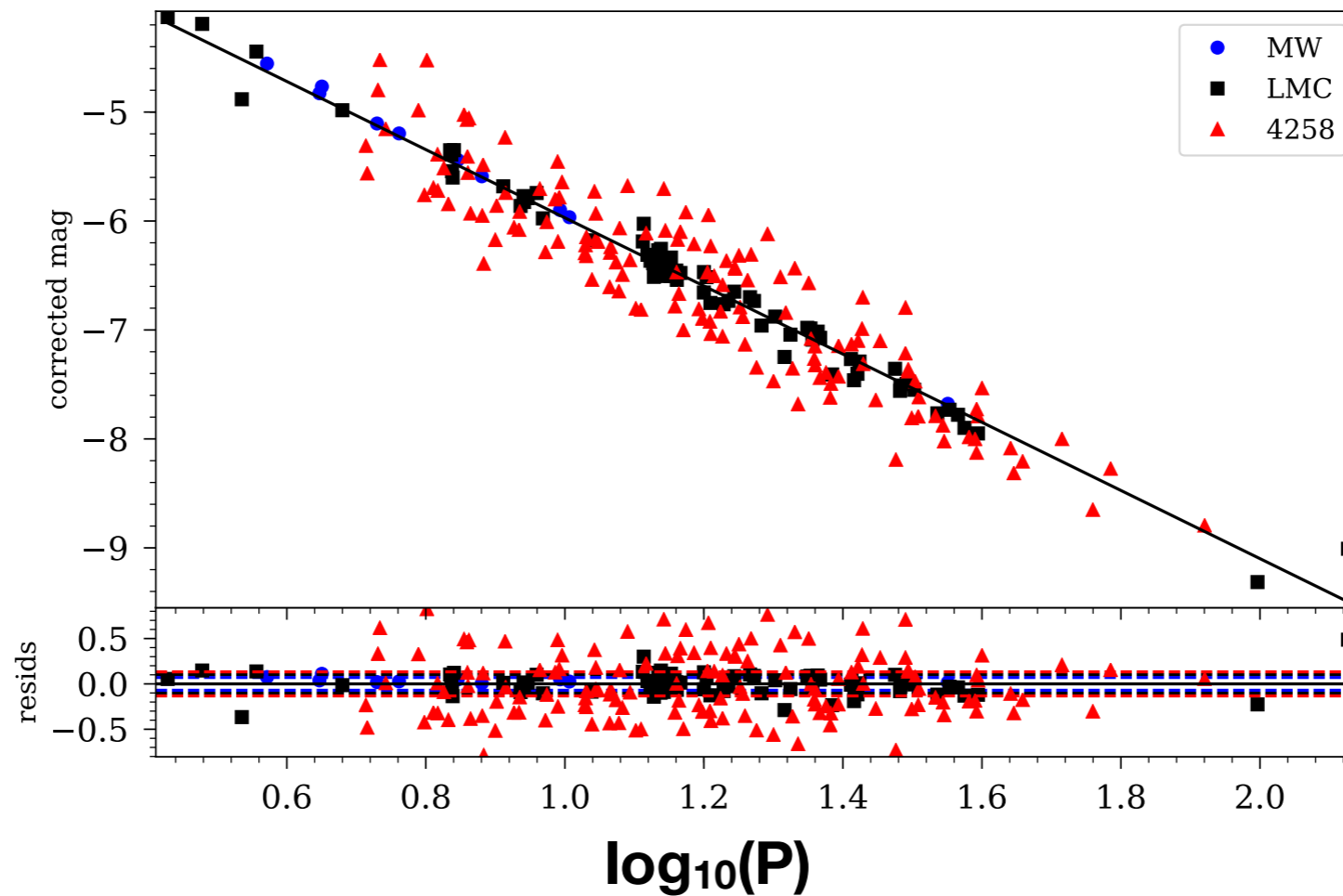


SN Hosts:

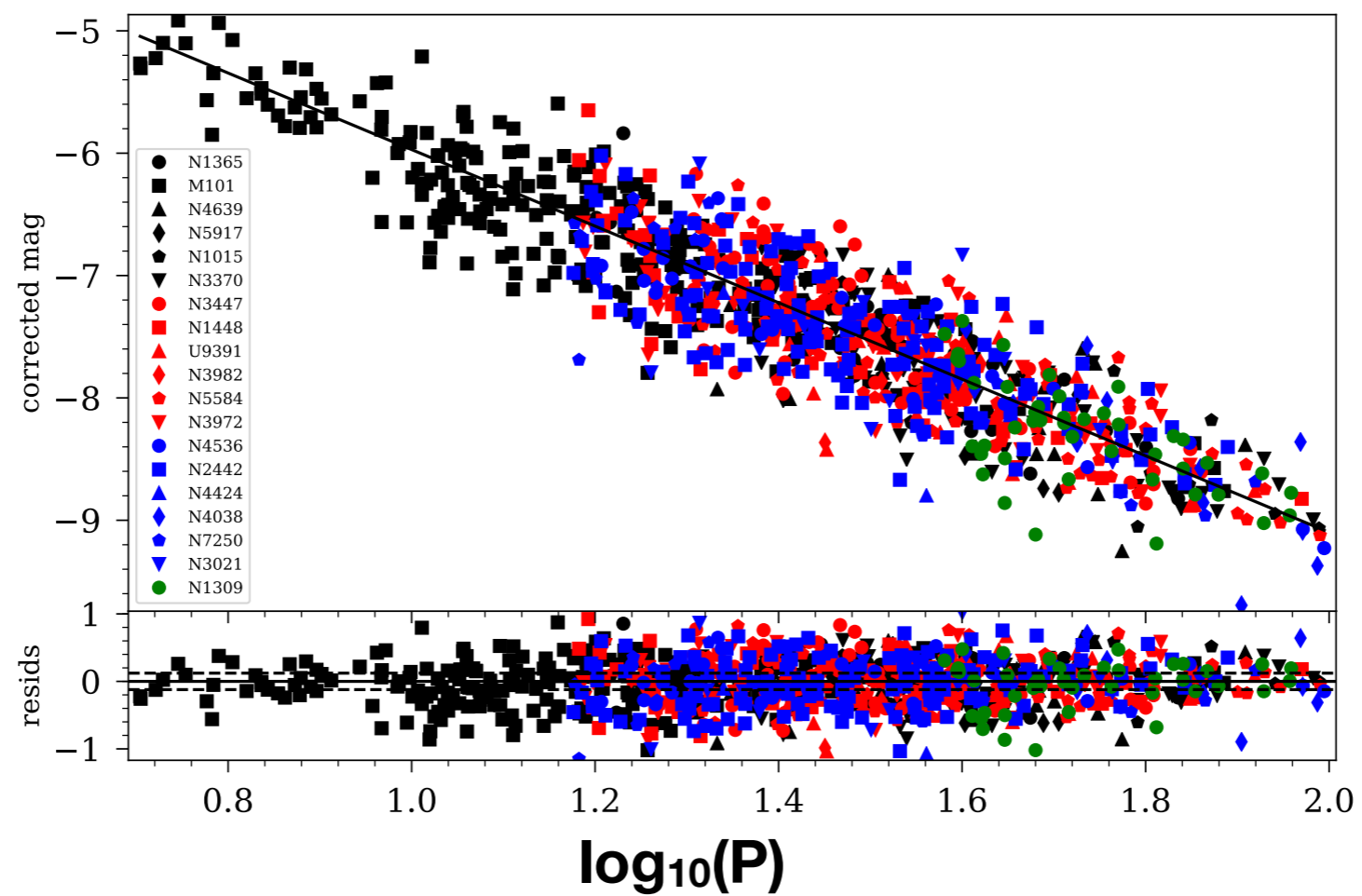


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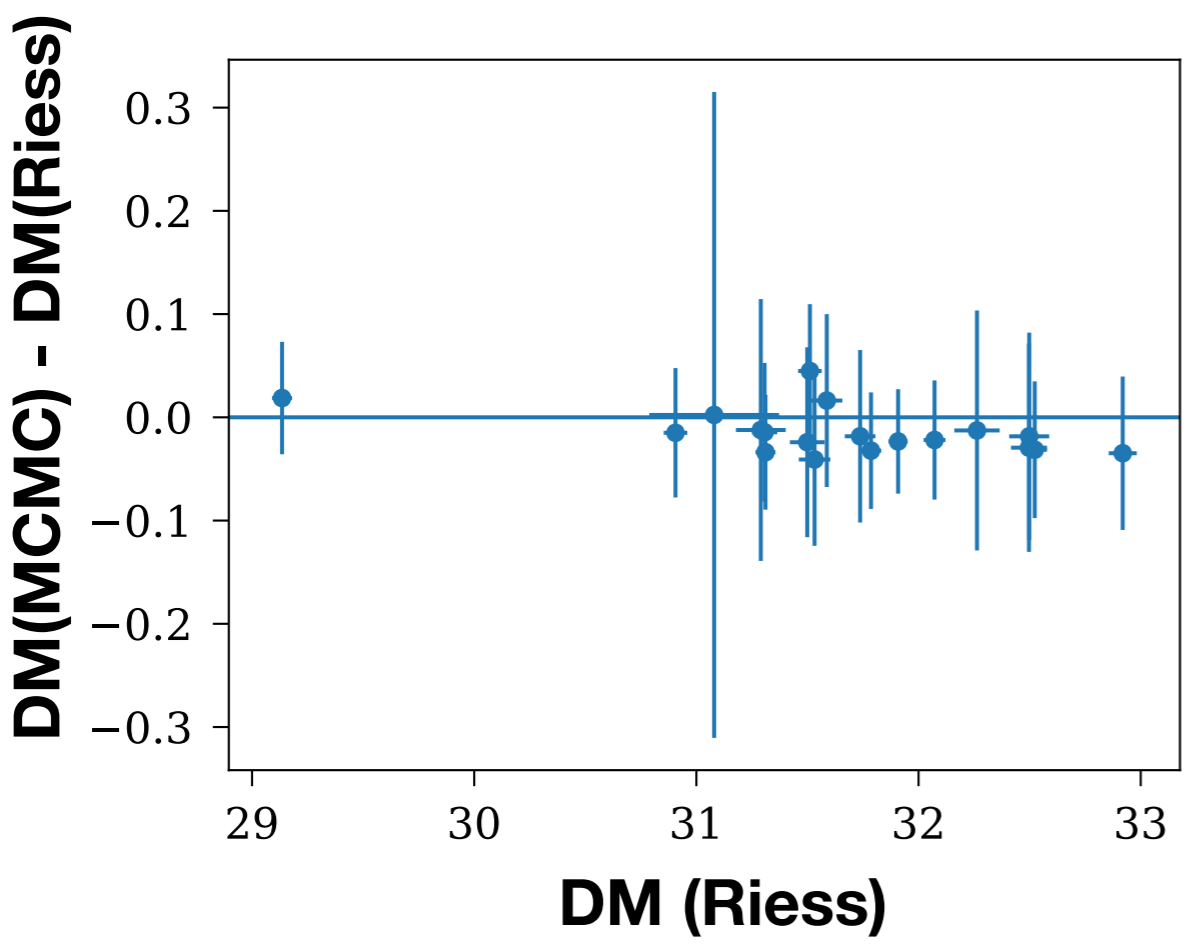
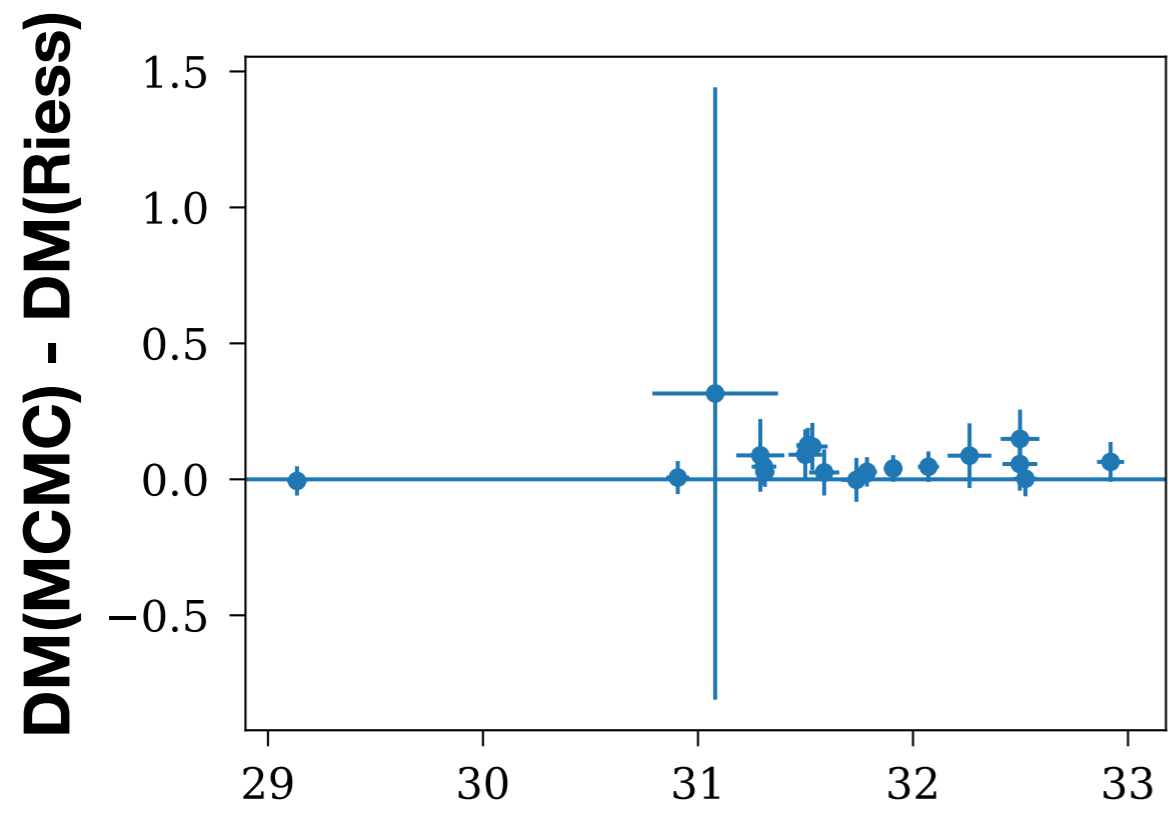
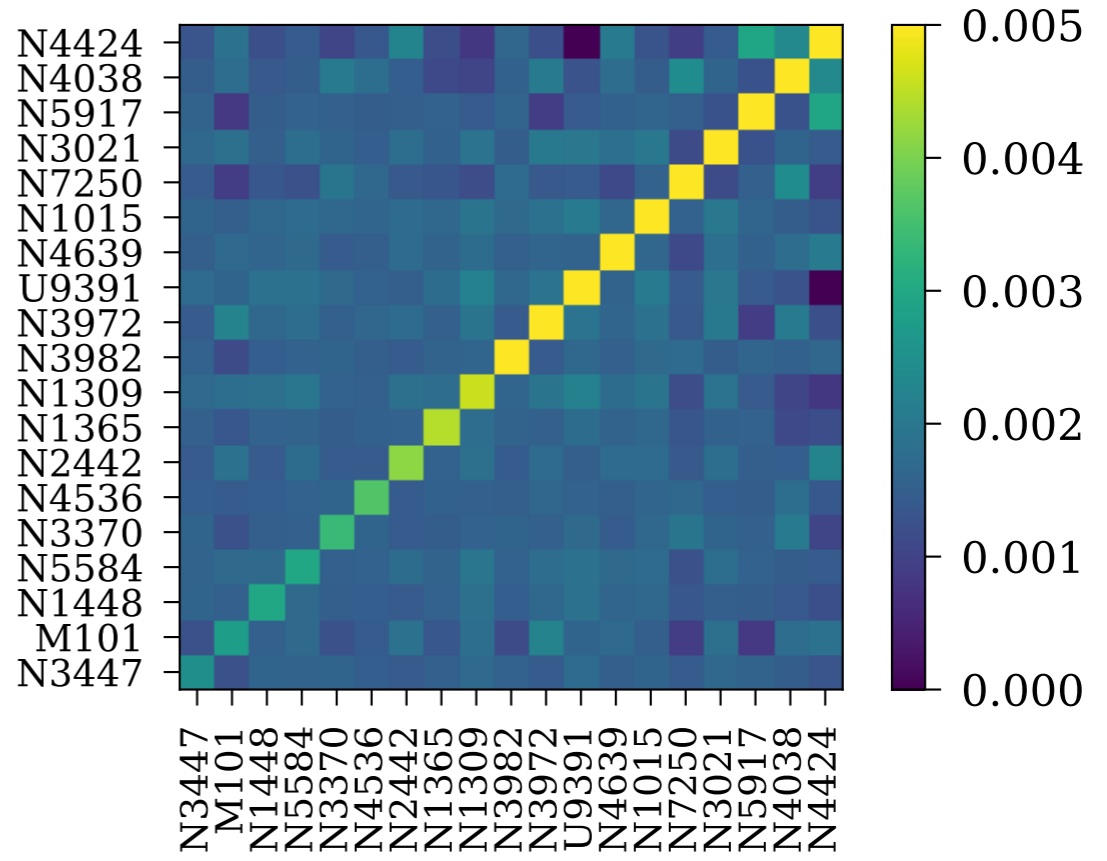
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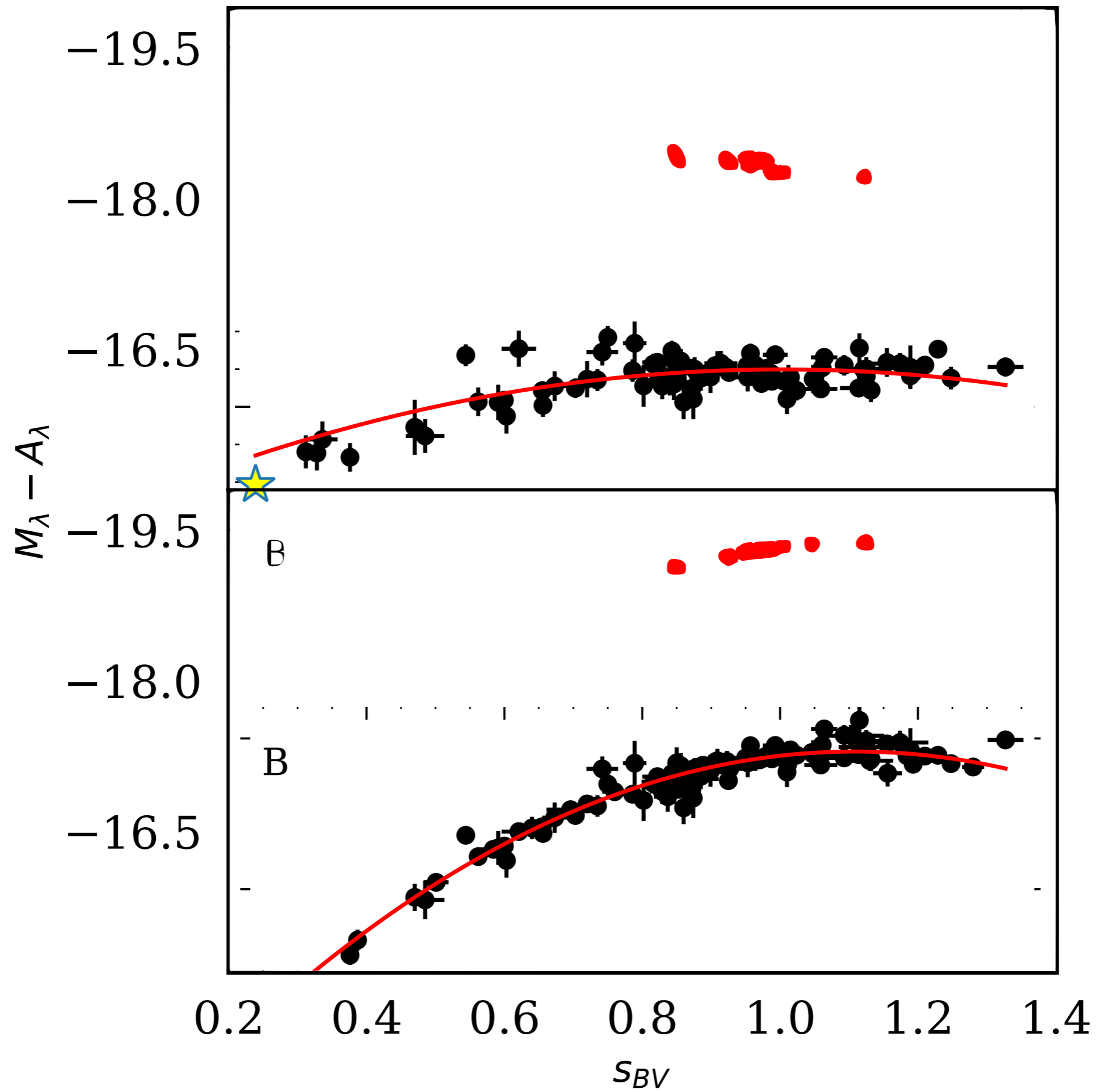


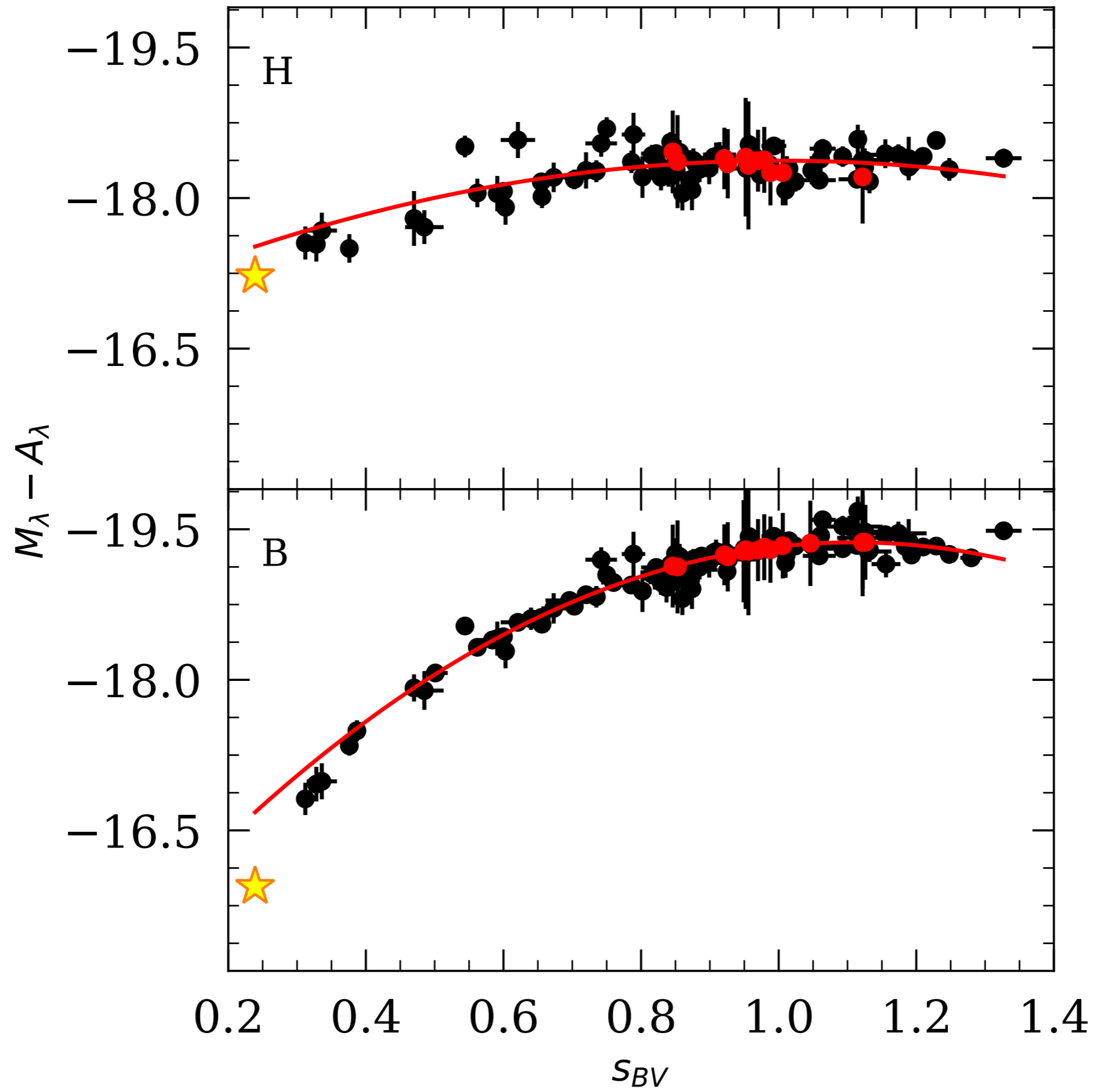
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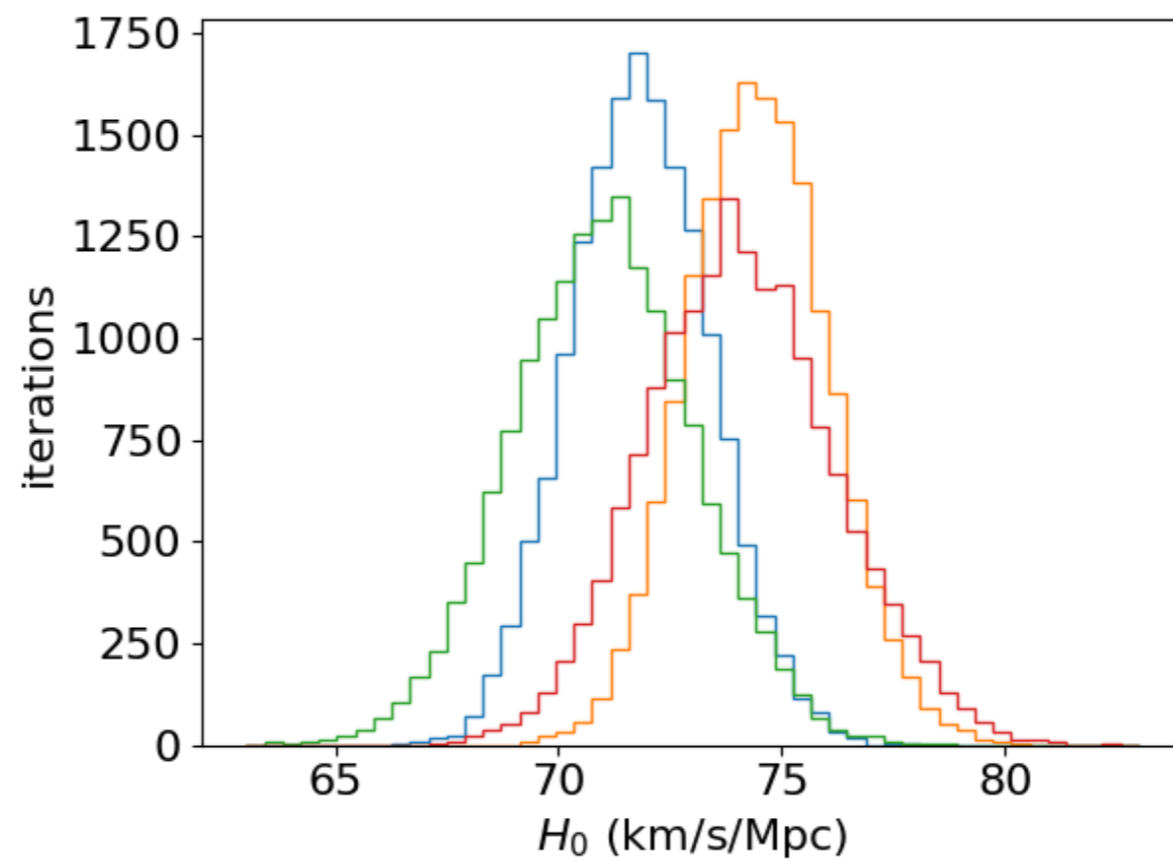
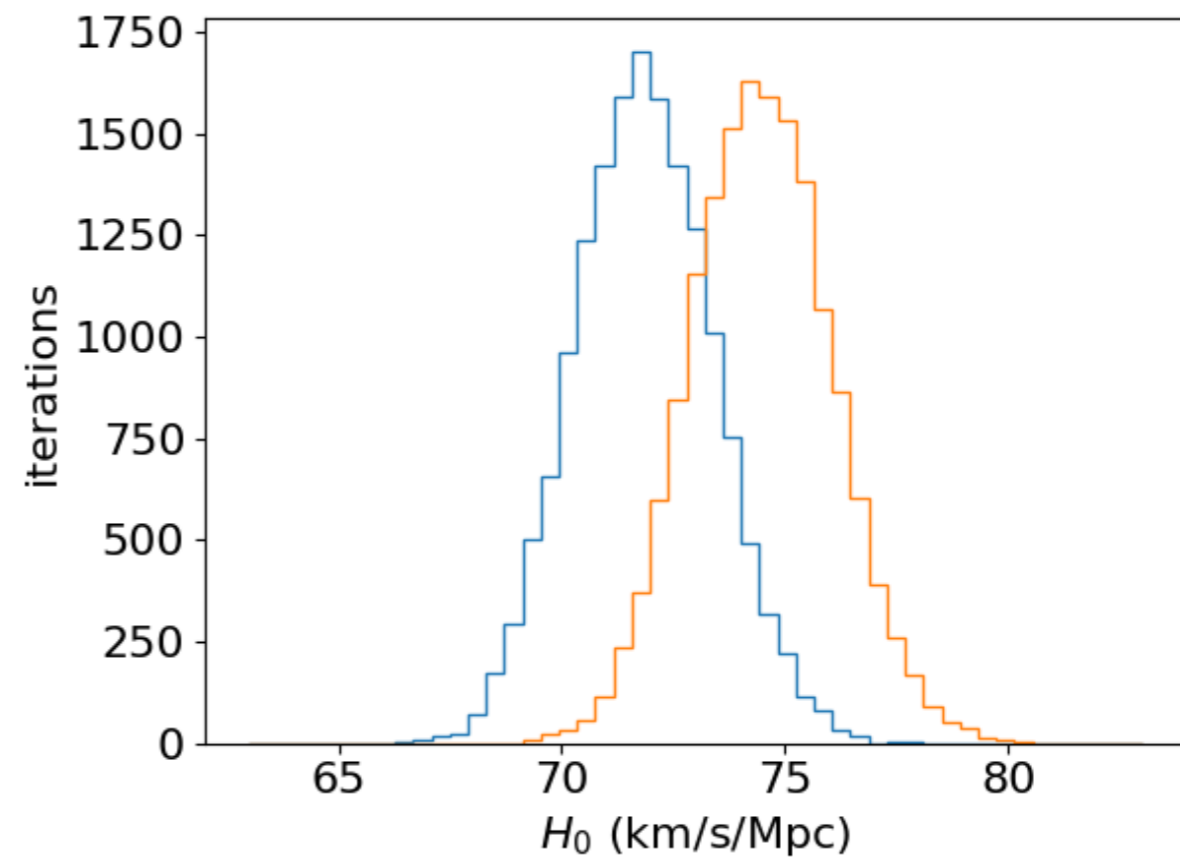
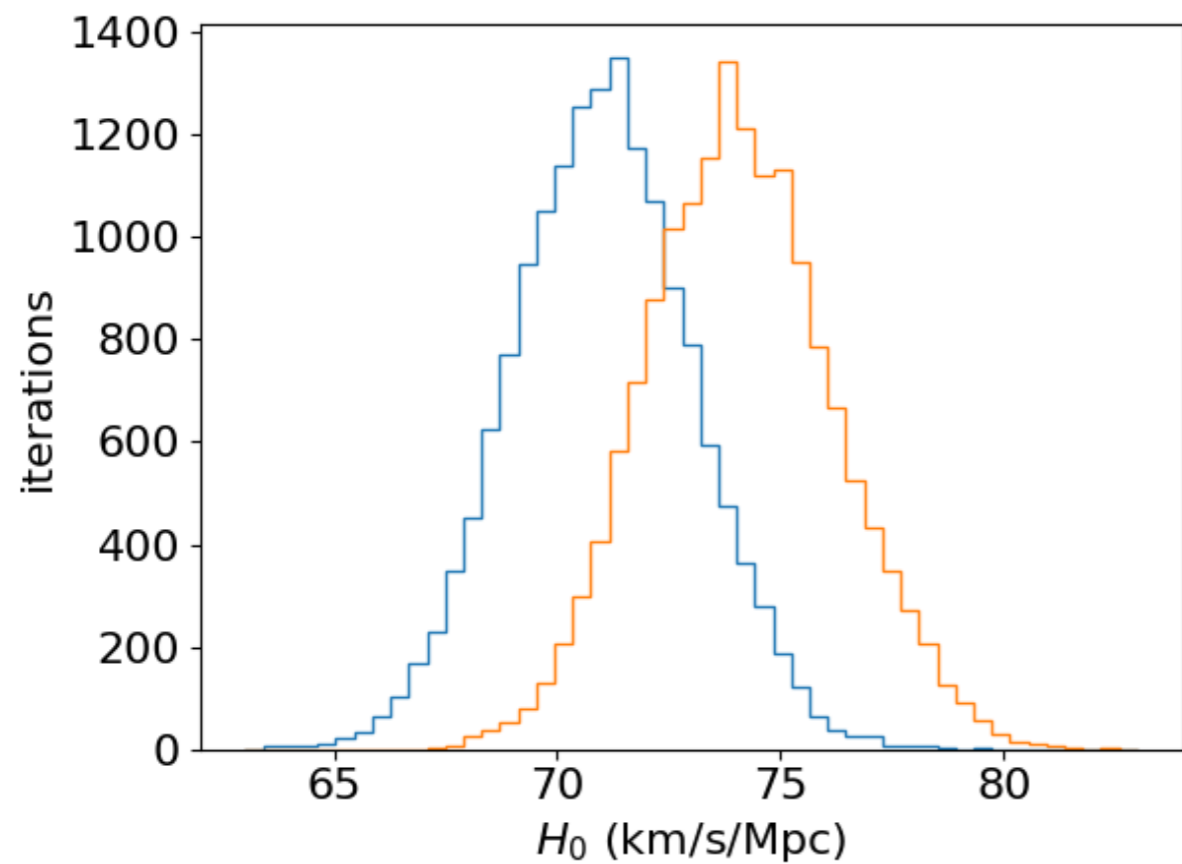


SN Host Covariance Matrix









What's Left

- Distribute first draft via overleaf
- Better Phillips relation fitting function (?)
- CSP zero-point errors
- Fresh eyes to tell me what's missing in the paper.